



Air Force Research Laboratory



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C/NOFS Thermospheric Research and Reentry EXperiment (T-RREX)

**Ionospheric Effects Symposium
12-14 May 2015**

**L. C. Gentile, C. G. Fesen
Air Force Research Laboratory Space
Vehicles Directorate**

R.A. Heelis

University of Texas at Dallas

R. F. Pfaff

NASA Goddard Space Flight Center

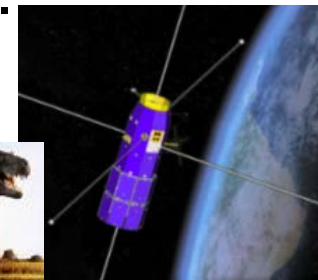




C/NOFS T-RREX



- As of 16 April 2015, C/NOFS has completed 7 years on orbit!
- Spacecraft systems are nominal; all payloads are working well;
- Apogee/perigee are now: ~495 km/353 km.
- Storm-time energy inputs on 17 Mar 2015 produced a global increase of ~250°K in thermosphere and ionosphere temperature; neutral density at fixed altitudes increased by a factor of 2.
- Initial observations below the peak of the F-layer show a consistent pattern; when plasma density is depleted in the bottomside, the ion motion (drift) is westward.





C/NOFS Status



Instrument	Partner	Provides	Status
NWM Neutral Wind Meter	NASA Small Explorer CINDI grant to U. Texas at Dallas	Neutral winds, neutral pressure	Cross-track ON
IVM Ion Velocity Meter		Ion velocities, temperatures, composition	ON
CORISS C/NOFS Occultation Receiver for Ionospheric Sensing and Specification	The Aerospace Corporation	Line-of-sight TEC, PNT	ON
VEFI Vector Electric Field Instrument	NASA	Ion velocities, electric and magnetic fields	ON
CERTO Coherent Electromagnetic Radio Tomography	NRL	TEC, electron density profiles, scintillation phase and amplitude	OFF
PLP Planar Langmuir Probe	AFRL/RVB	Ion and electron densities and fluctuations	Cycling in sun/eclipse

- Spacecraft systems are nominal.
- All payload instruments are working well.
- CERTO beacon is off; PLP is cycling in eclipse to conserve power while beta angle is marginal.
- As of 16 April 2015, C/NOFS has completed 7 years on orbit!



CINDI: Coupled Ion-Neutral Dynamics Investigation

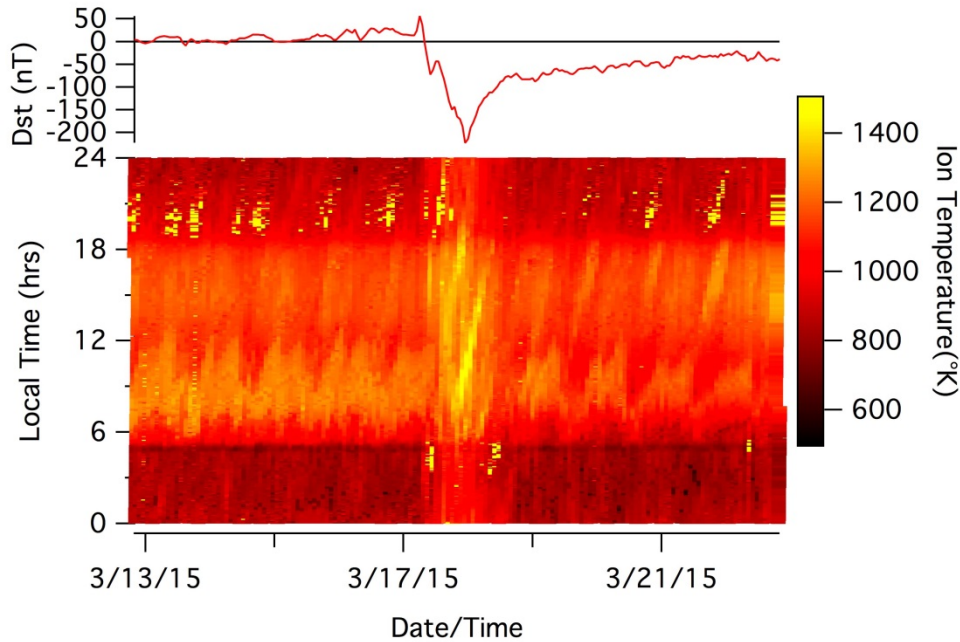


New Results: Coupled Ion Neutral Dynamics Investigation (CINDI) Observations

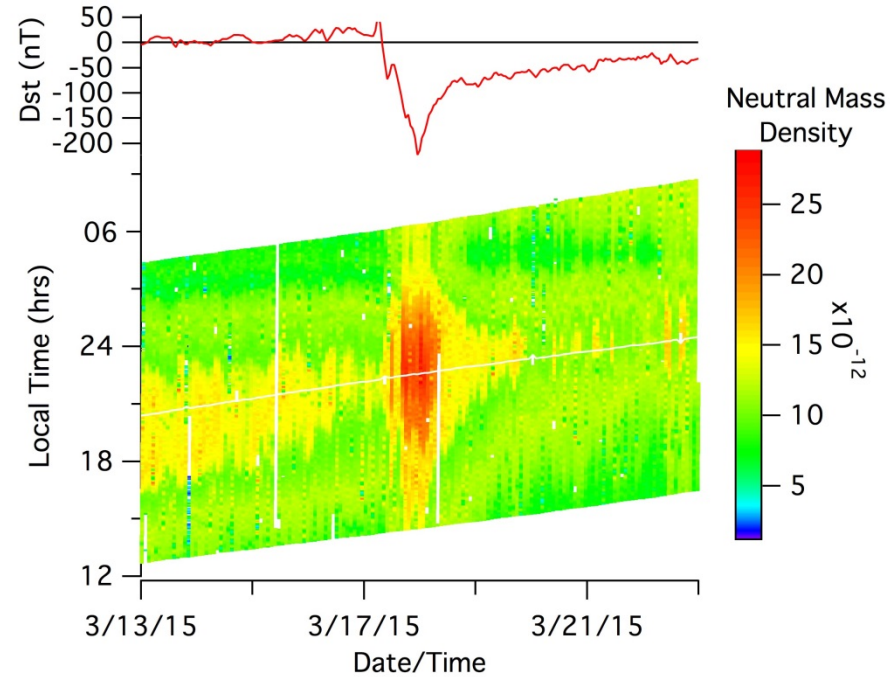


Response of Equatorial Ionosphere and Thermosphere to 17 March 2015 Storm

CINDI Storm-time Ion Temperature



CINDI C/NOFS Storm Time Neutral Density



Storm-time energy inputs produce global increase of $\sim 250^{\circ}\text{K}$ in thermosphere and ionosphere temperature.

Global storm-time increase in temperature produces expansion of the thermosphere; neutral density at fixed altitudes increased by a factor of 2.

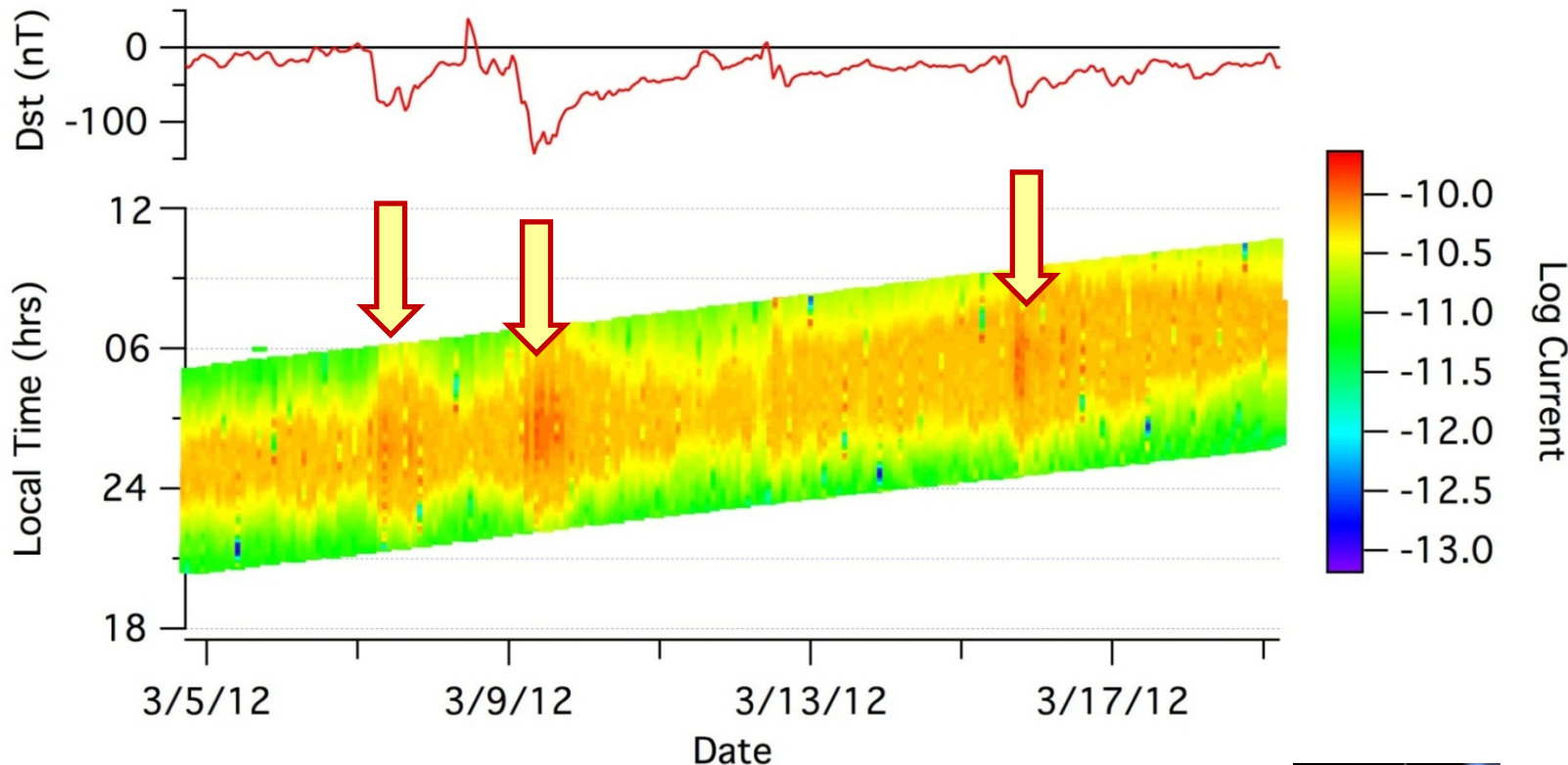




C/NOFS CINDI Neutral Density Variations with Local Time vs Dst



Density data are highly structured and clearly respond to geomagnetic activity. Effects persist for a day or more, well into the recovery phase of the storm.



Reentry predictions may be off 20° – 40° in latitude if low-altitude neutral density variations are not incorporated into predictive models.

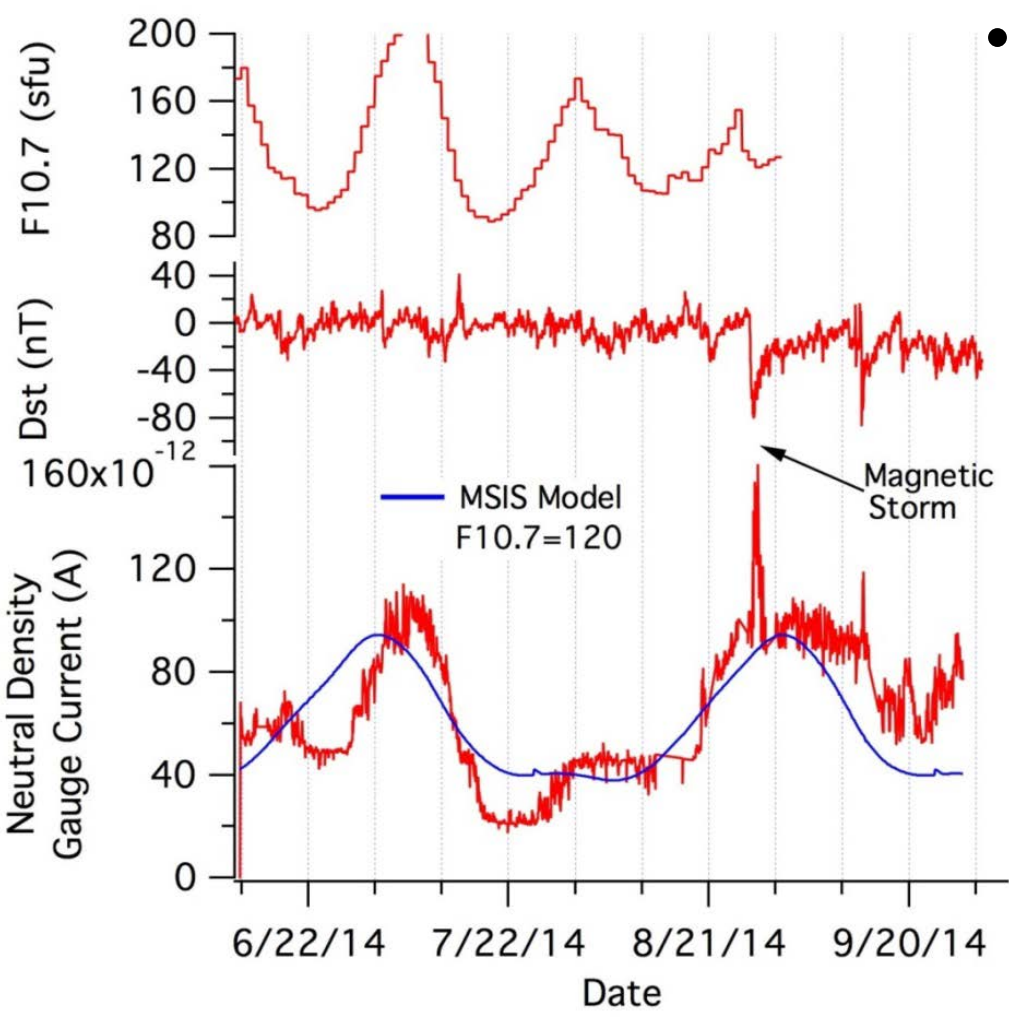




C/NOFS CINDI Neutral Density vs MSIS Model



C/NOFS CINDI NWM Neutral Density



Solar Activity (F10.7) Index

- Data reflect typical variability due to 27-day solar rotation and geomagnetic storms

Magnetic Activity (Dst) Index

- Typically varies between +/- 50 nT
- May fall below -400 nT in major storms
- Lowest Dst since launch: ~ -225 nT, so effects of major storms remain TBD

Neutral Density Variations

- Red: C/NOFS NWM observations
 - Clear response to solar activity
 - Large/impulsive responses to geomagnetic activity
- Blue: standard neutral wind model predictions

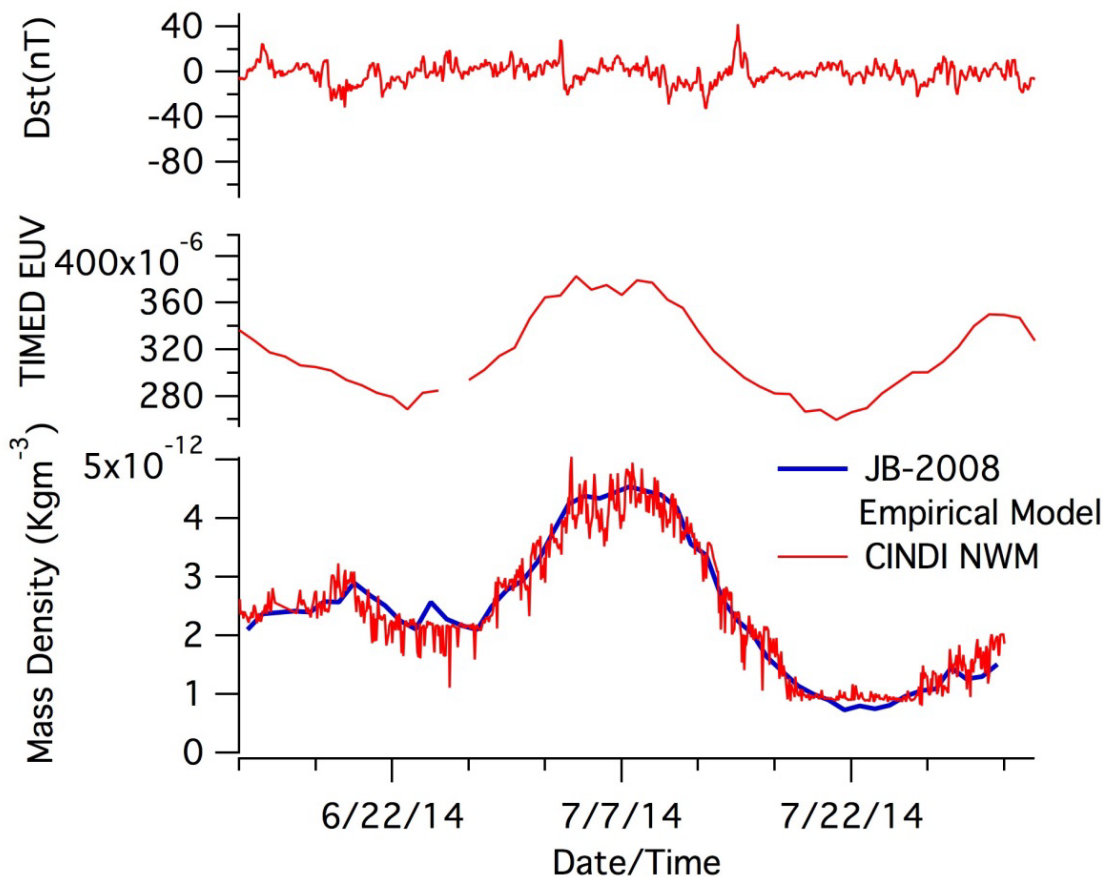




C/NOFS CINDI Neutral Density vs JB2008 Orbital Drag Model



CINDI C/NOFS NWM Neutral Pressure Variations



Magnetic Activity (Dst) Index

- Typical quiet time variations between +/- 40 nT

Solar EUV from TIMED

- Typical daily variations

Neutral Density: Quiet Time

- Red: C/NOFS observations
- Blue: Jacchia-Bowman 2008 AFSPC operational orbital drag model

Note: this is just one example for a very quiet time interval. There is much more analysis to be done under varying solar conditions.

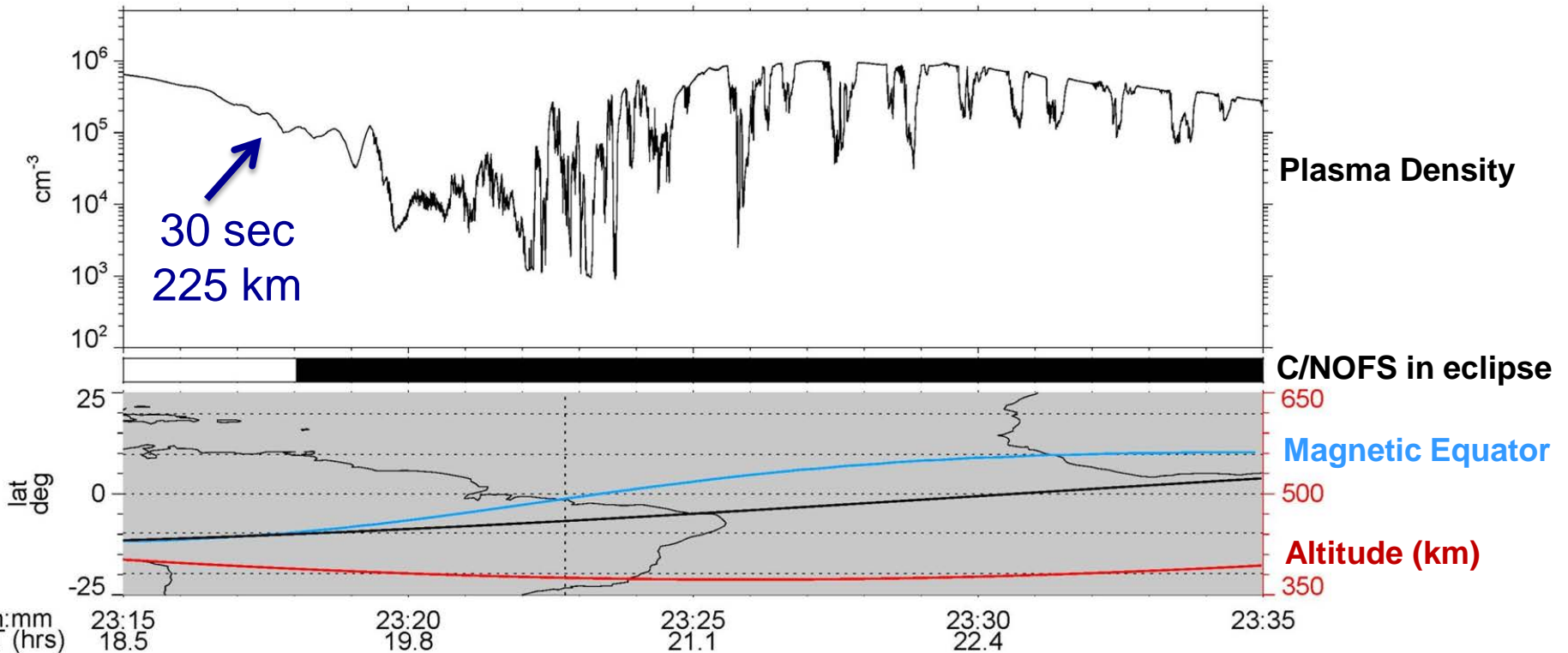




C/NOFS Vector Electric Field Instrument (VEFI) Observations of Bottomside Structures



VEFI Observations



VEFI observations below the peak of the F-layer show bottomside wave structures and plasma density irregularities.

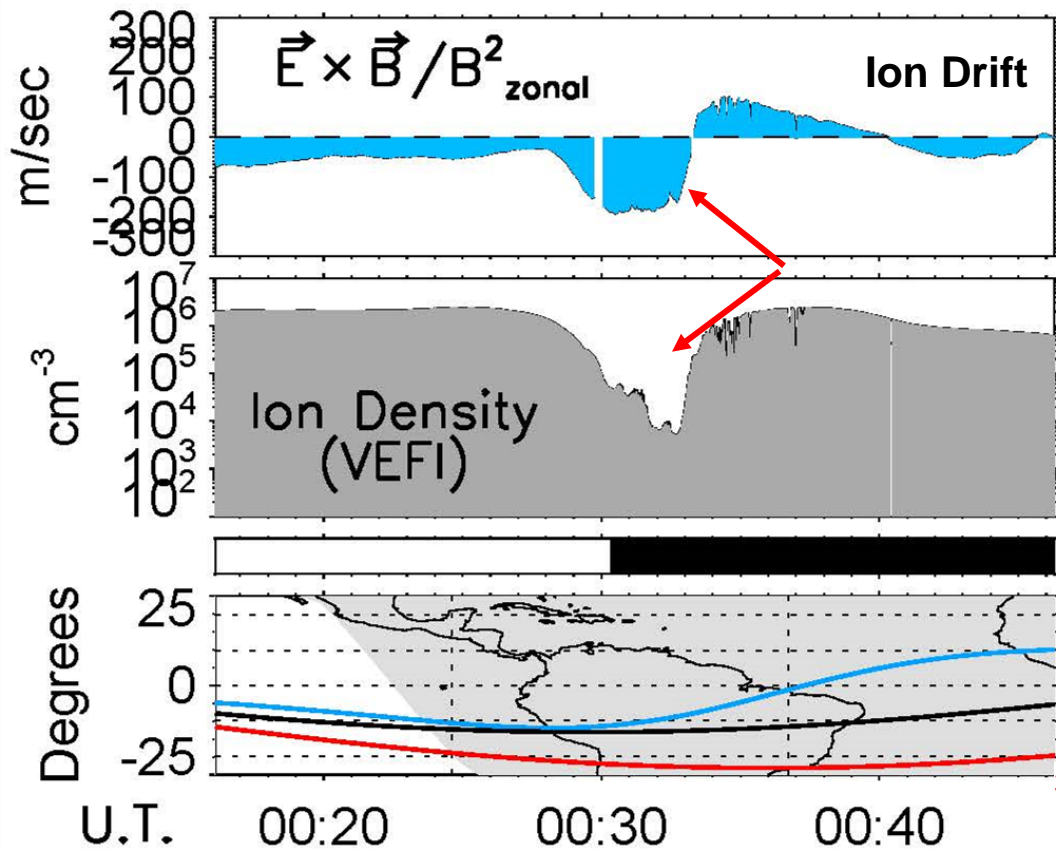




VEFI Shows Relationship Between Bottomside Density and Plasma Drift Motion



Orbit 36922
Jan 15, 2015



Initial analysis shows a consistent pattern in the bottomside observations; when plasma density is depleted in the bottomside, the ion motion (drift) is westward.

Plasma Density

C/NOFS in eclipse

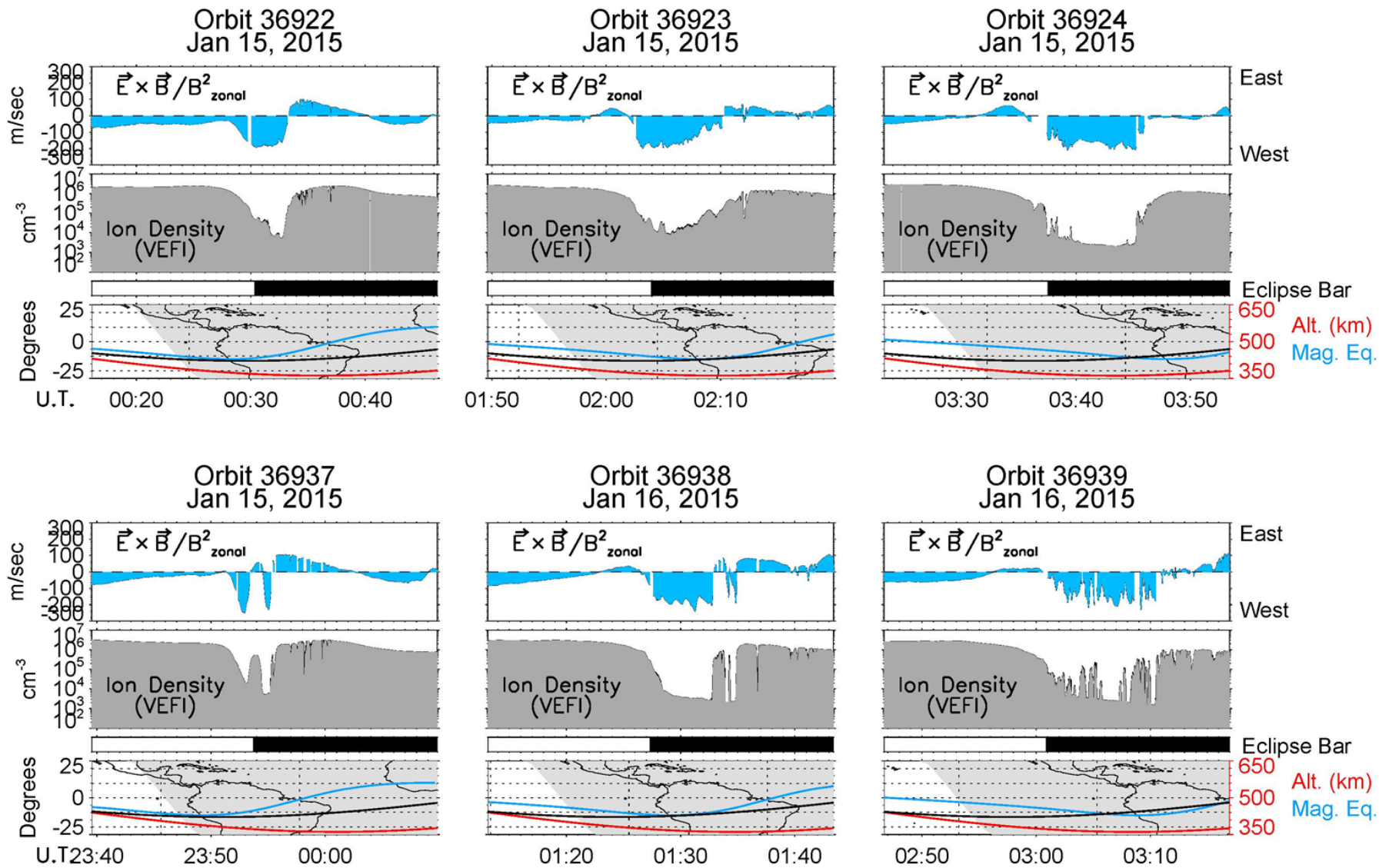
Magnetic Equator

Altitude (km)



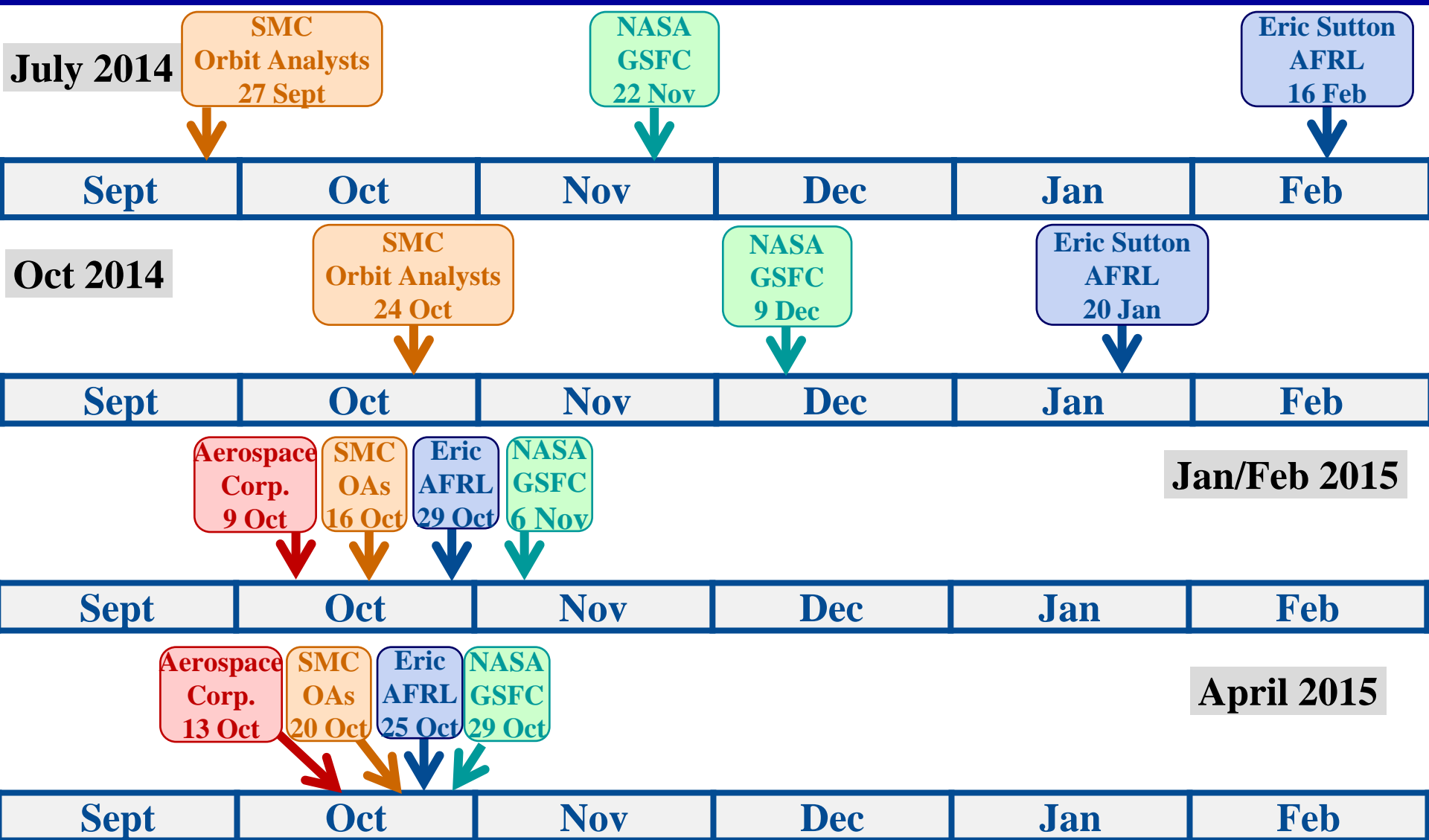


VEFI Shows Relationship Between Bottomside Density and Plasma Drift Motion





Reentry Prediction Mean Estimates Are Converging





C/NOFS T-RREX Summary



- C/NOFS is poised to gather significant new low-altitude observations in the low-latitude ionosphere.
- These data will advance our understanding of ionospheric coupling above and below the F peak and fill gaps in ionospheric models, providing a far more realistic picture of the low-latitude, low-altitude ionosphere.
- Initial results are quite promising, but there is still much work to be done in understanding the mechanisms that couple the topside and bottomside of the ionosphere.
- We have never had this opportunity before, and we may not have it again in the foreseeable future.
- We invite and encourage any collaborations that will leverage the results from this unique data set.

