What potential science problem can we address using SAMBA-AMBER-MEASURE Network?

Endawoke Yizengaw Senior Scientist and PI of AMBER Institute for Scientific Research, Boston College











Outline

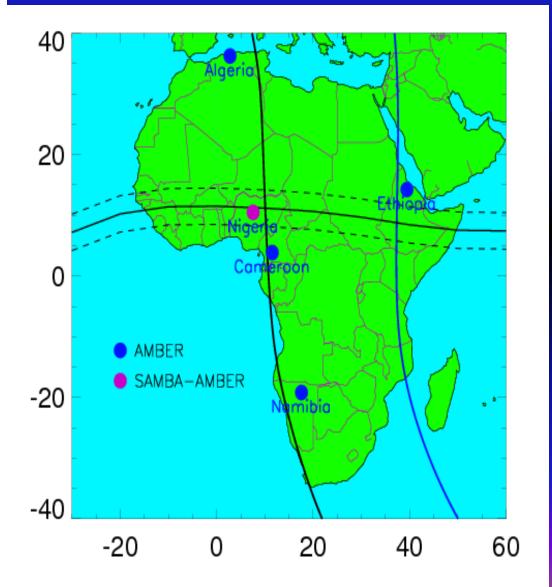
Introduction and objective of AMBER network

How SAMBA instruments we deployed in Chile is important to Chilean?

Potential Science to be done

What other instruments can we use to augment magnetometers' data?

Objective of AMBER Project AMBER (African Meridian B-field Education and Research)

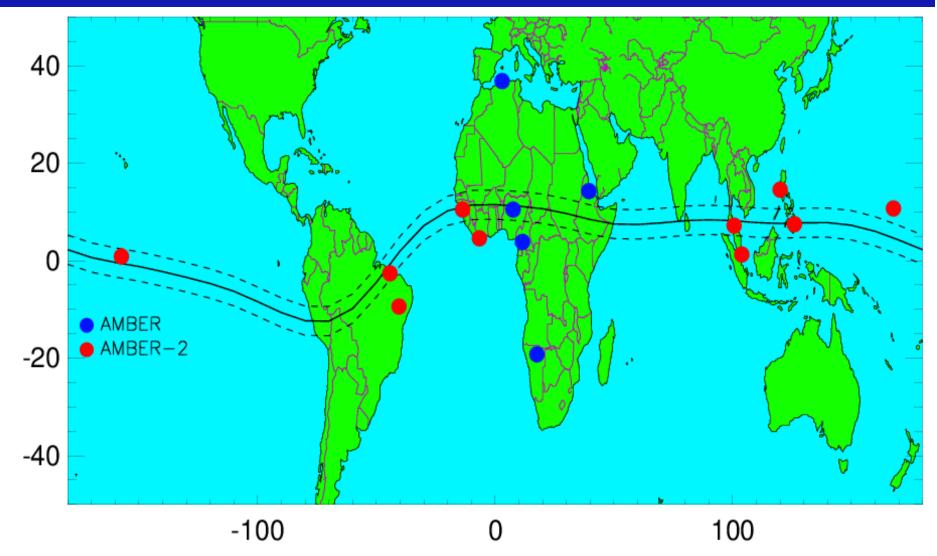


the processes governing electrodynamics of the equatorial ionosphere as a function of local time, longitude, magnetic activity, and season, and

ULF pulsation strength and its connection with equatorial electrojet strength at low/midlatitude regions.

Current AMBER Magnetometer Network

Team Members: E. Yizengaw (BC, PI); M. Moldwin (UM); E. Zesta (NASA); M. Magoun (BC); K. Hector (UCLA)



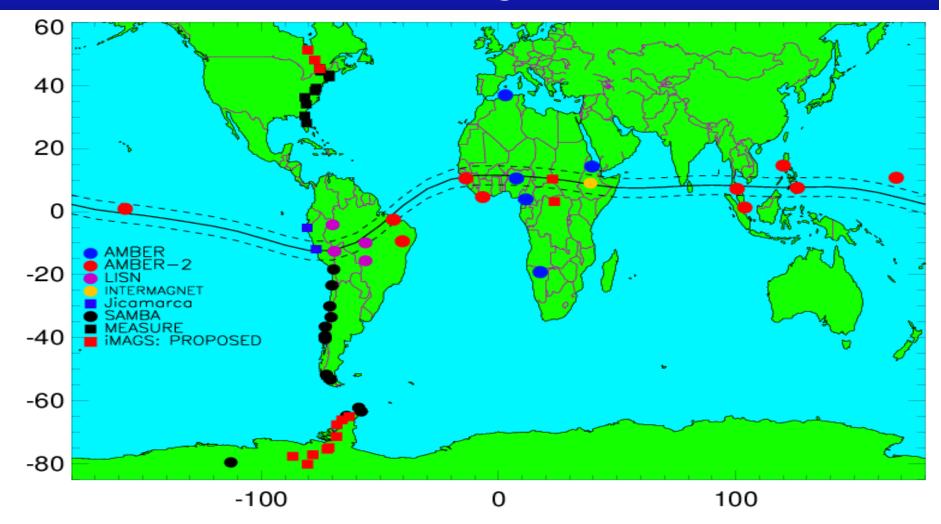
Where to get more information about AMBER and the SAMBA-AMBER database

Detail Information About AMBER https://www2.bc.edu/~kassie/AMBER.html

New AMBER/SAMBA/MEASURE database http://magnetometers.bc.edu/

Team Members: M. Moldwin (UM); E. Yizengaw (BC); E. Zesta (NASA); A. Boudouridis (SSI); M. Magoun (BC); K. Hector (UCLA)

Why we are here in Chile?



How SAMBA instruments we deployed in Chile is important to Chilean?

→ First it trains and exposes young Chileans to science & technology

It measures space currents induced to the magnetic field which is dangerous to the power lines because Chile has facility/cities at high latitudes

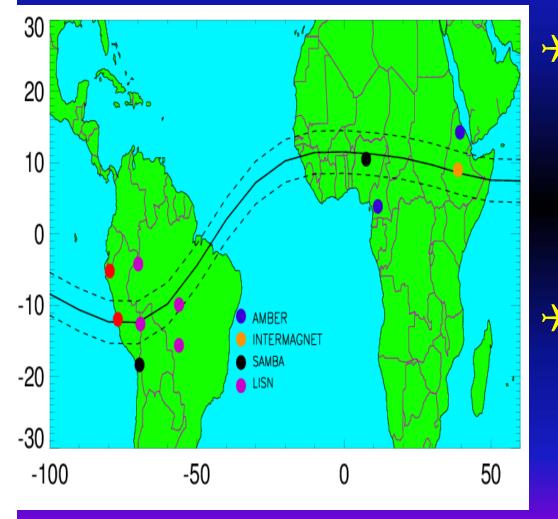
It opens opportunities for Chilean scientists to participate in the global space weather research activities

Potential Science to be done!

- What is the longitudinal variability of the equatorial vertical drift?
- → What is the impact of ULF wave penetration to lower latitudes? Is it related to scintillation activities?
- → Remotely monitoring the plasmasphere boundary layer location. Key input for radiation belt acceleration.
- → Is the magnetospheric plasma mass density and ionospheric electron density enhancements/depletion correlated? FLR studies?
- Why the ULF power and TEC have hemispheric asymmetries?
- → How strong the GIC current in Chile and what can we do about it?

What is the longitudinal variability of the equatorial vertical drift?

Electrodynamics Variablity



 Disturbances due to geomagnetic impact
 On H-components at the equator and off the equator

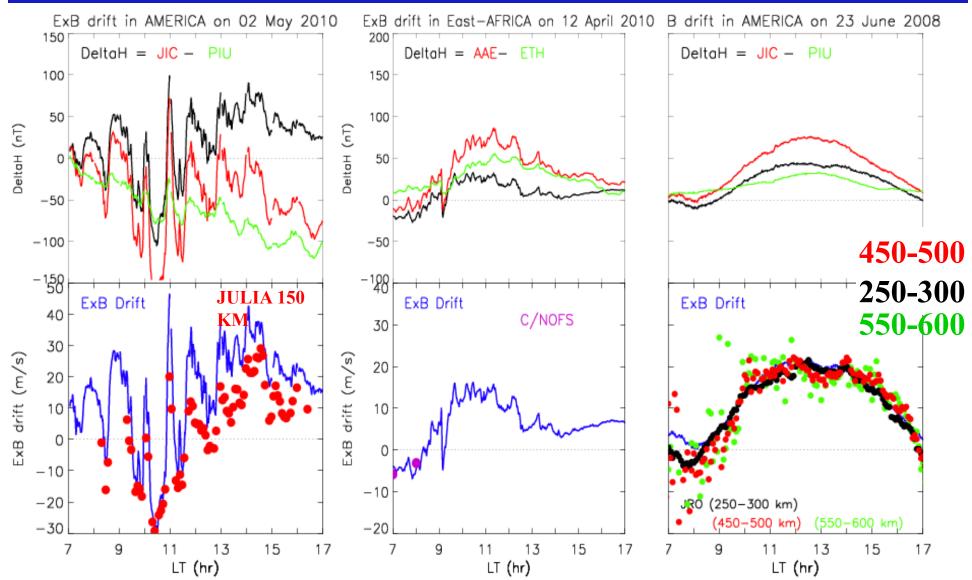
Disturbances due to
 EEJ
 Only on H-component at the equator

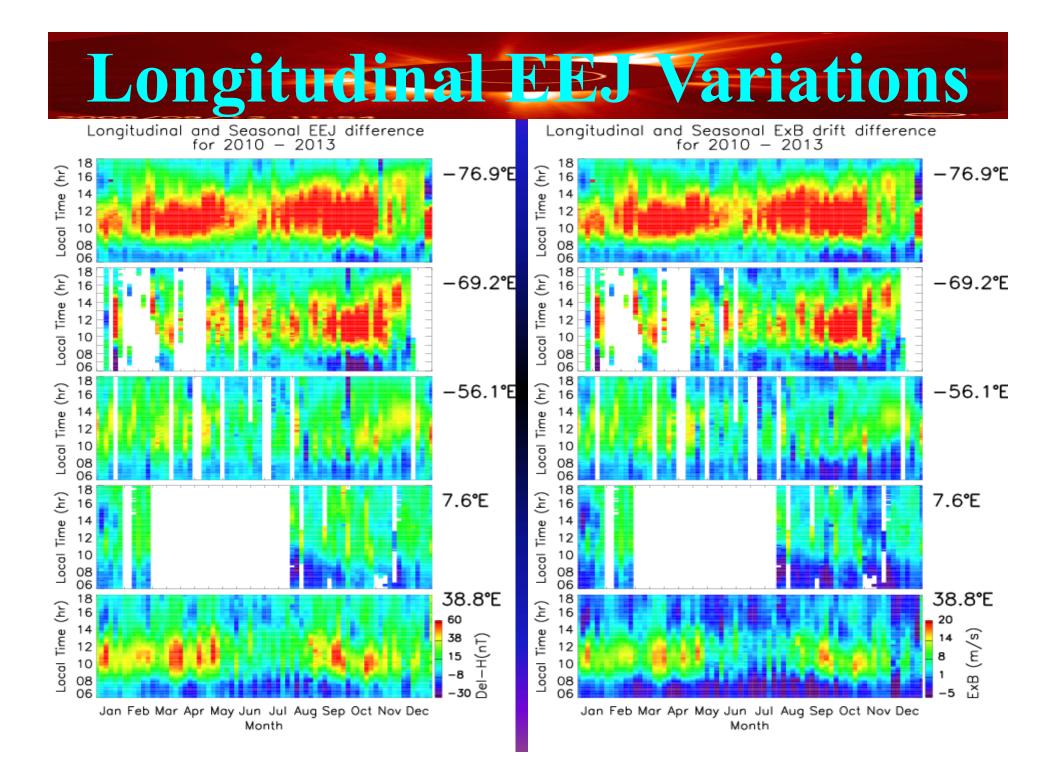
Comparison with other observations

with JULIA

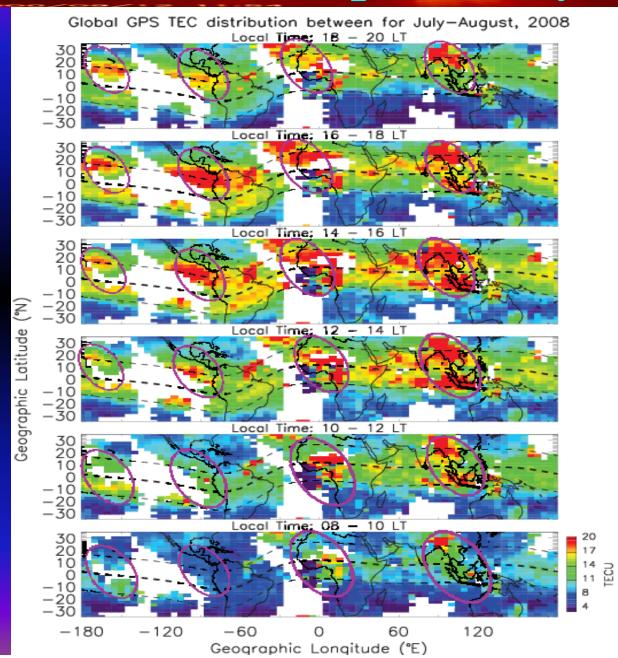
C/NOFS - Africa

ISR - America



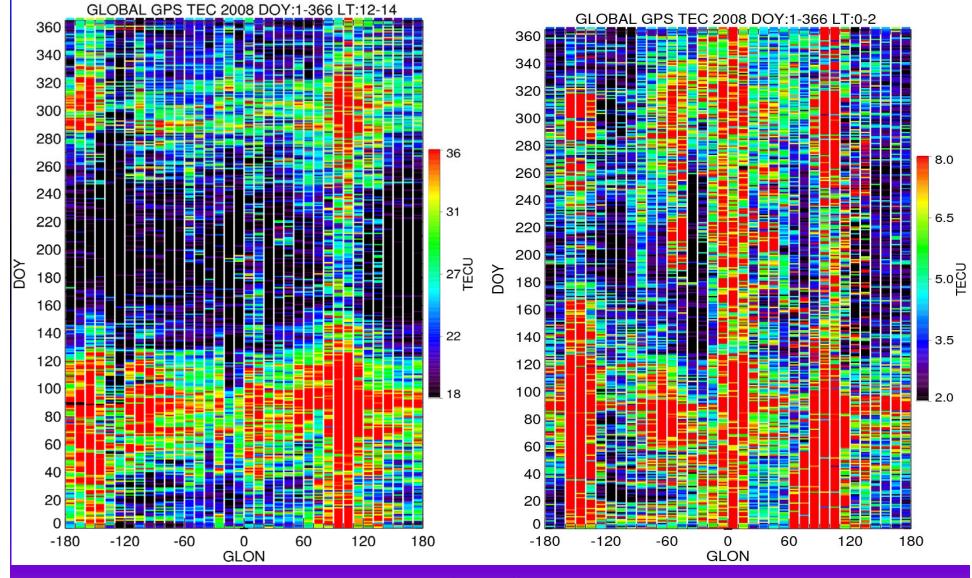


Global TEC map for July-August 2009



Yizengaw, IJG, 2012

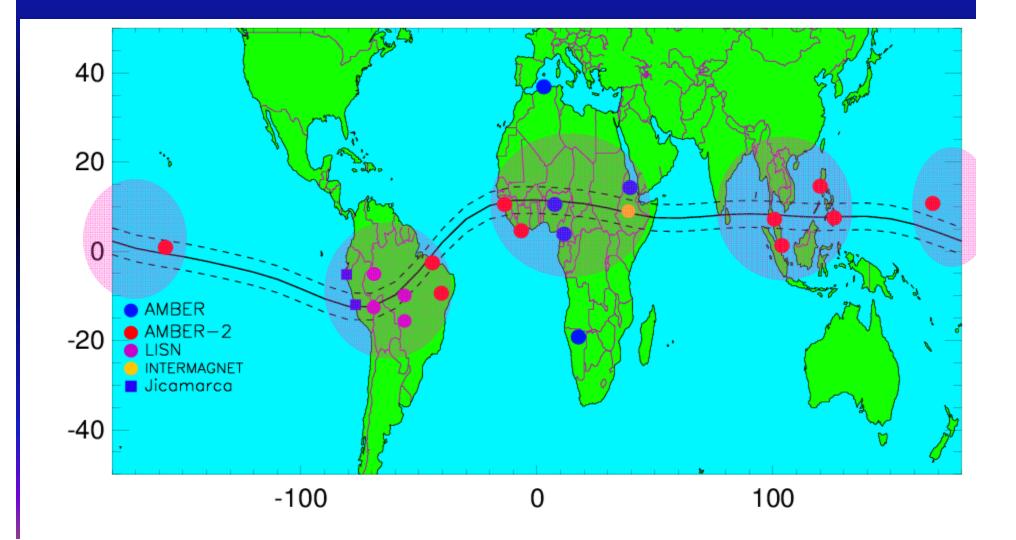
Day-to-day wave number 4 structure at different local time



Yizengaw and Pacheco, JGR, 2013

AMBER and other magnetometer networks

- → AMBER PI: Endawoke Yizengaw
- → LISN PI: Cesar Valladeres

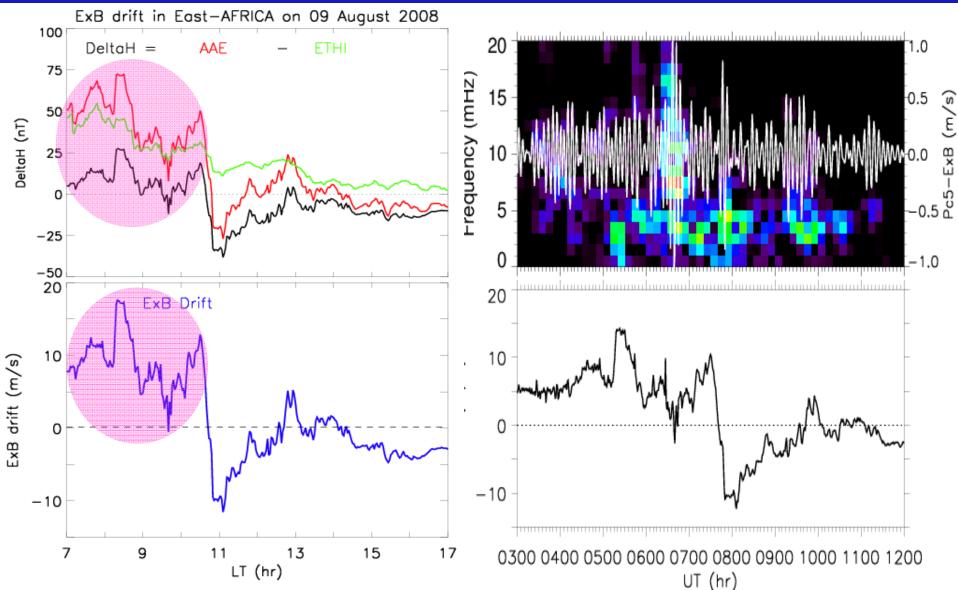


What is the impact of ULF wave penetration to lower latitudes? Is it related to scintillation activities?

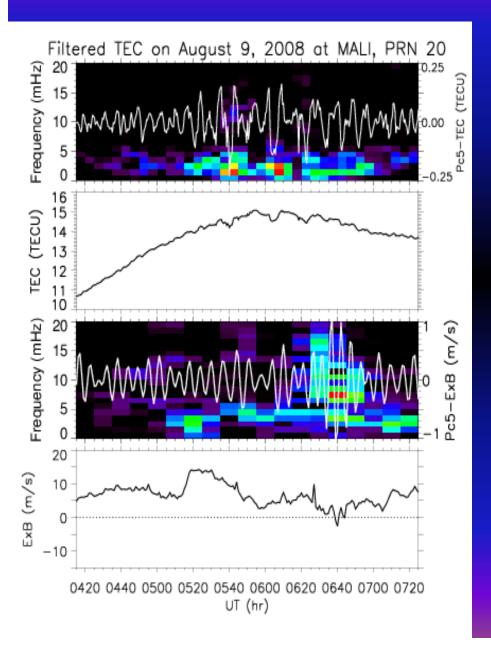
Equatorial Vertical Drift Fluctuation

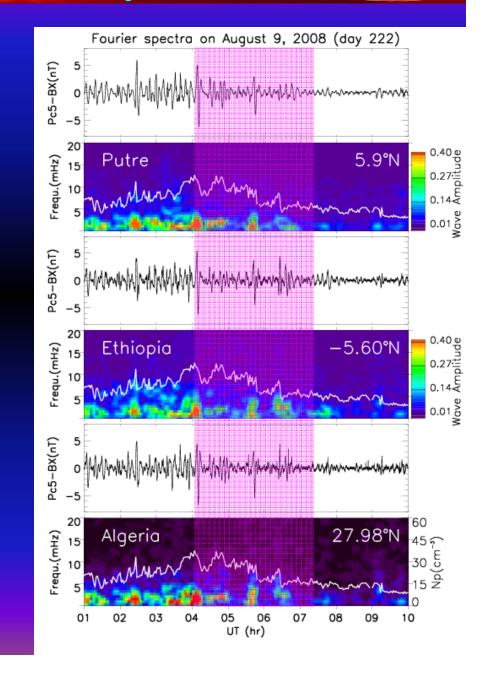
frican sector

Band-pass filtered drift shows ULF wave

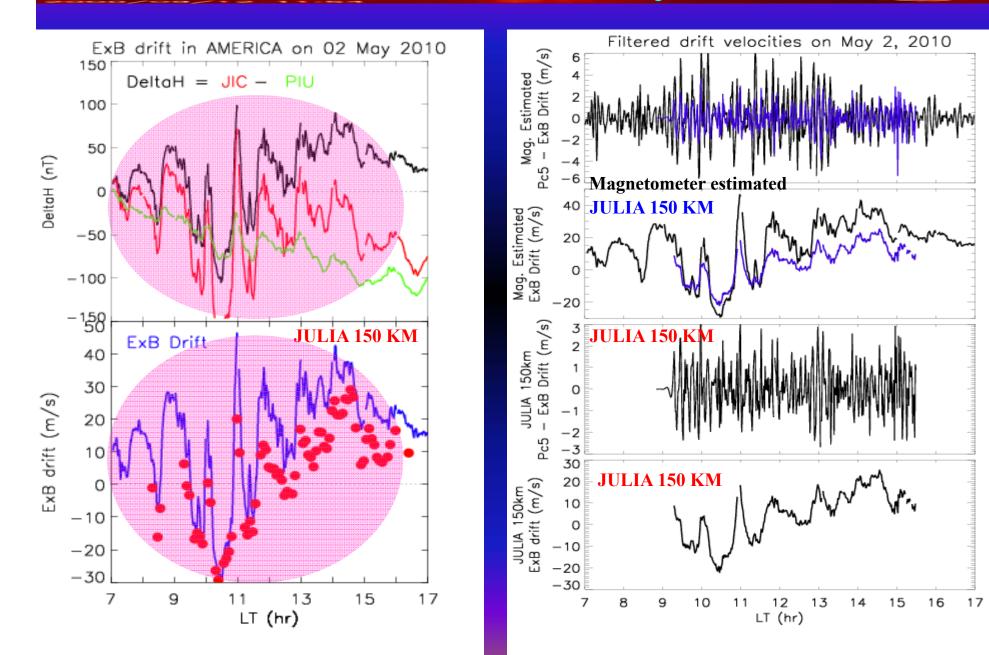


Vertical Drift and Density Fluctuation



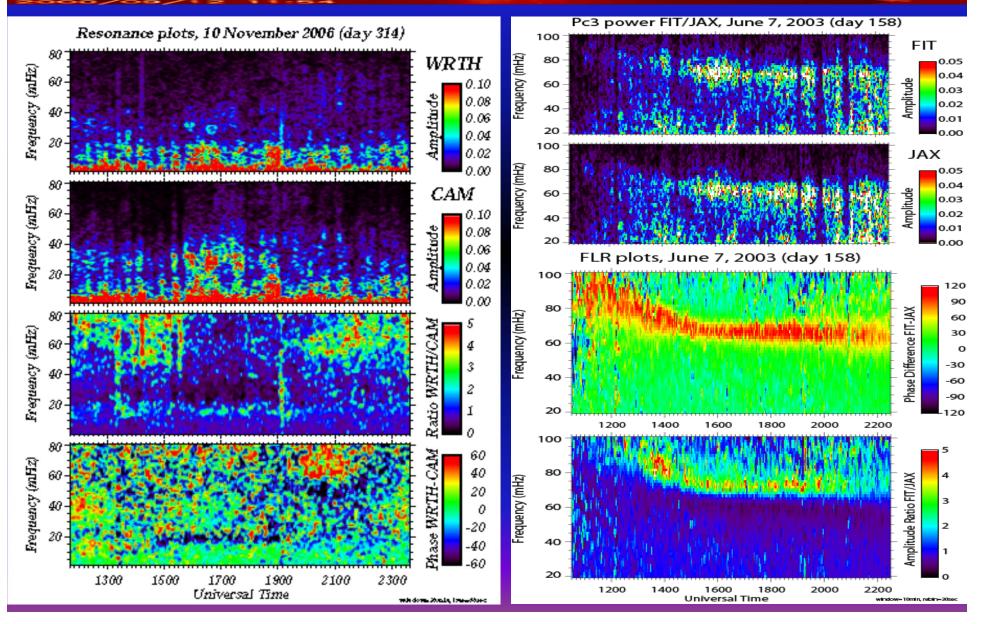


Vertical Drift and Density Fluctuation

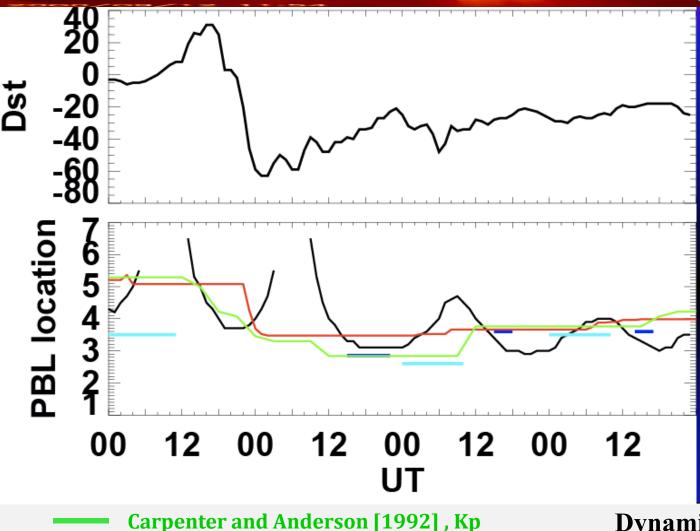


Is it possible to estimate the plasmapause location from the ground? Key input for radiation belt acceleration

Lpp from reverse Phase Difference Reported before by Menk et al. [2004]



PBL determination comparisons



O'Brien and Moldwin [2003], Dst

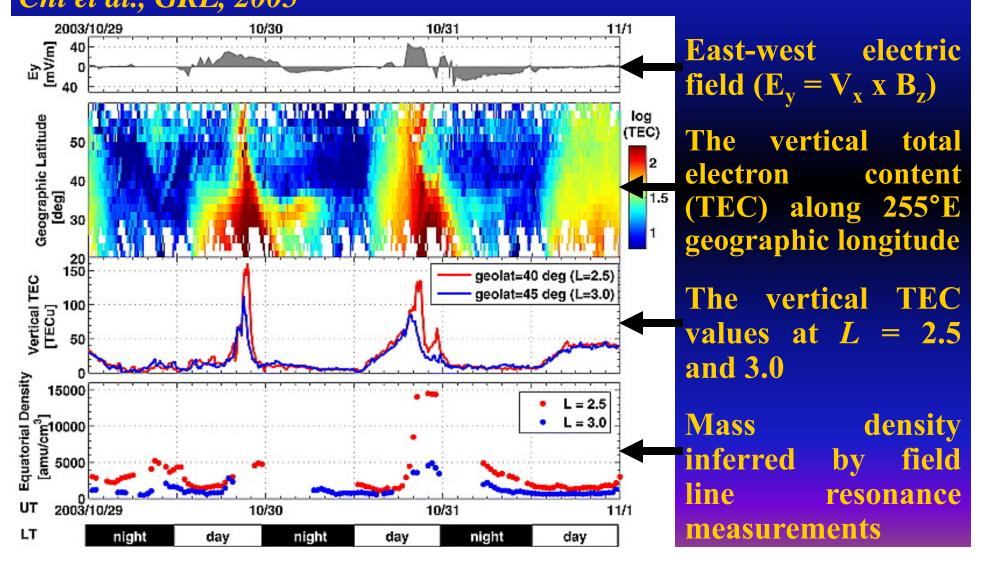
TEC profile from ground GPS receivers

Reverse PD from ground mags

Dynamic Global Core — Plasma Model (DGCPM), [Ober, 1997]. DGCPM is a 2D single-species model

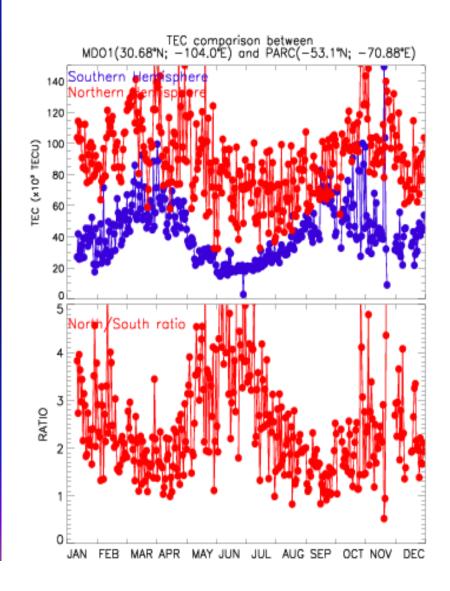
Is the magnetospheric plasma mass density and ionospheric electron density enhancements/depletion **correlated? FLR studies?**

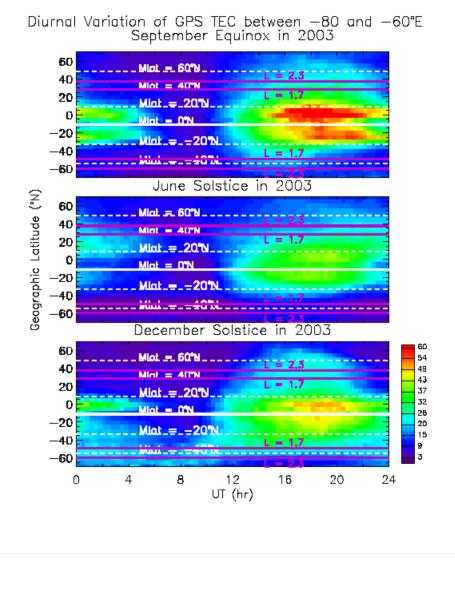
Simultaneous density enhancement in the plasmasphere and ionosphere *Chi et al., GRL, 2005*



Why the ULF power and TEC have hemispheric asymmetries?

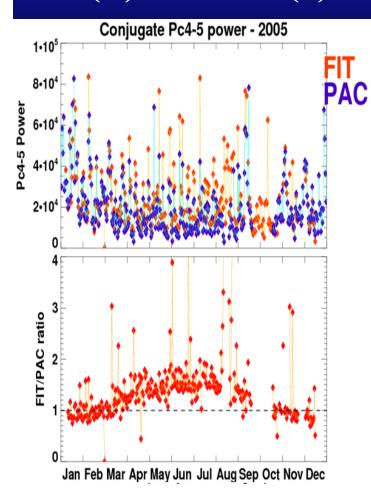
ULF power and TEC asymmetry



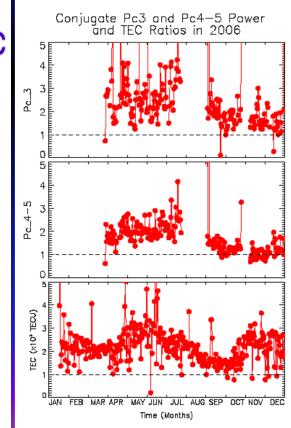


ULF power and TEC asymmetry

AMER (N) and OHI (S) 2006

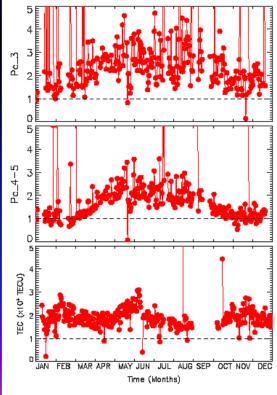


FIT (N) and PAC (S)



Conjugate Pc3 and Pc4-5 Power and TEC Ratios in 2007

2007



How strong the GIC current in Chile and what can we do about it?

GIC induction on long conducting cable

1956 trans Atlantic cable (yet metallic)



Power losses increased in the transformers Saturation

Other large economic events

- → San Francisco Earthquake …… 1906 ………\$ 500B
- → Annual loss from Electric interruption\$ 80B
- → North American Power Grid Blackout\$ 30B/day
 → GEO satellite revenue loss>\$25B
- → GEO satellite revenue loss\$ >25B
- → Blackout of East Coast 1955\$ 10B
- → Mt Lassen Volcanic Eruption .. 1915\$ 5B

Baker et al.

What other instruments can we use to augment magnetometers' data?

Where can we get calculated GPS TEC?

Madrigal data base (developed by MIT group) provide global calculated GPS TEC, interpolated 2 x 2 degree latitude and lonigtudes. The calculated TEC is averaged every 5 minutes.

Madrigal calculated TEC data base

http://madrigal.haystack.mit.edu/cgi-bin/madrigal/madInvent.cgi

TOPEX and JASON altimeter TEC can be used to augment GPS TEC that you have been discussing so far?

TOEX and JASON provide altimeter TEC (i.e, vertical TEC below 1360 km altitude) over the ocean.

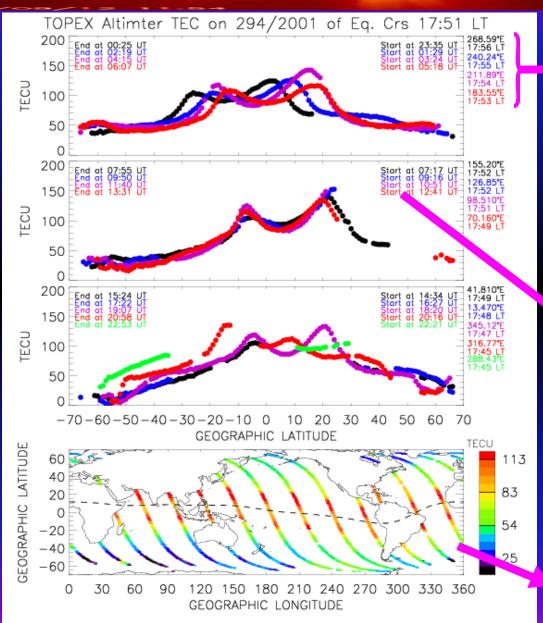
TOEX Altimeter TEC

ftp://podaac.jpl.nasa.gov/SeaSurfaceTopography/topex/L2/tp_ssha/

JASON Altimeter TEC

ftp://podaac.jpl.nasa.gov/SeaSurfaceTopography/jason1/L2/j1_ssha/

TOPEX altimeter TEC plot

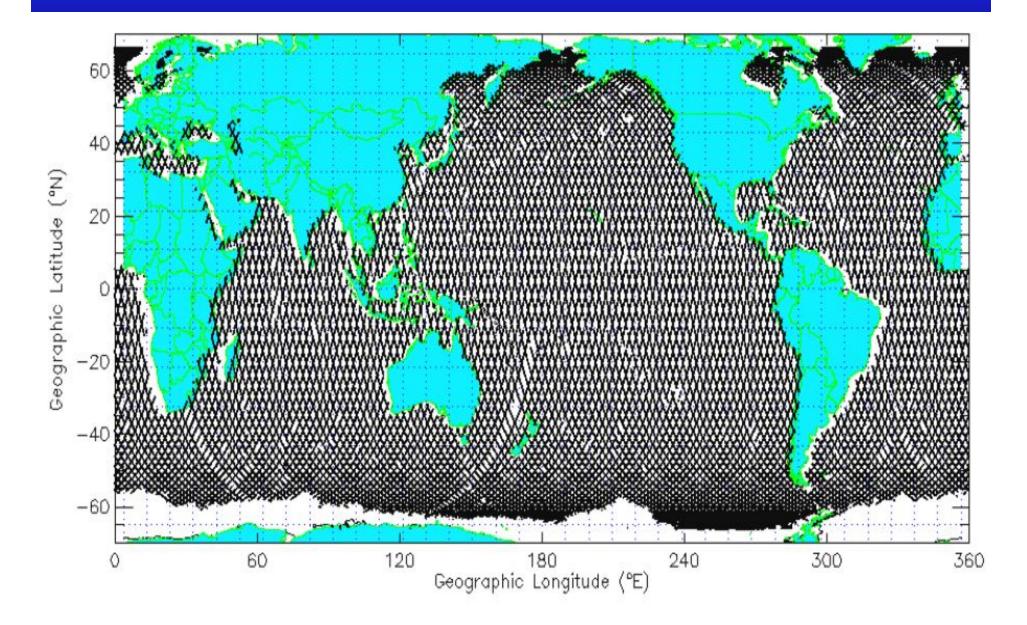


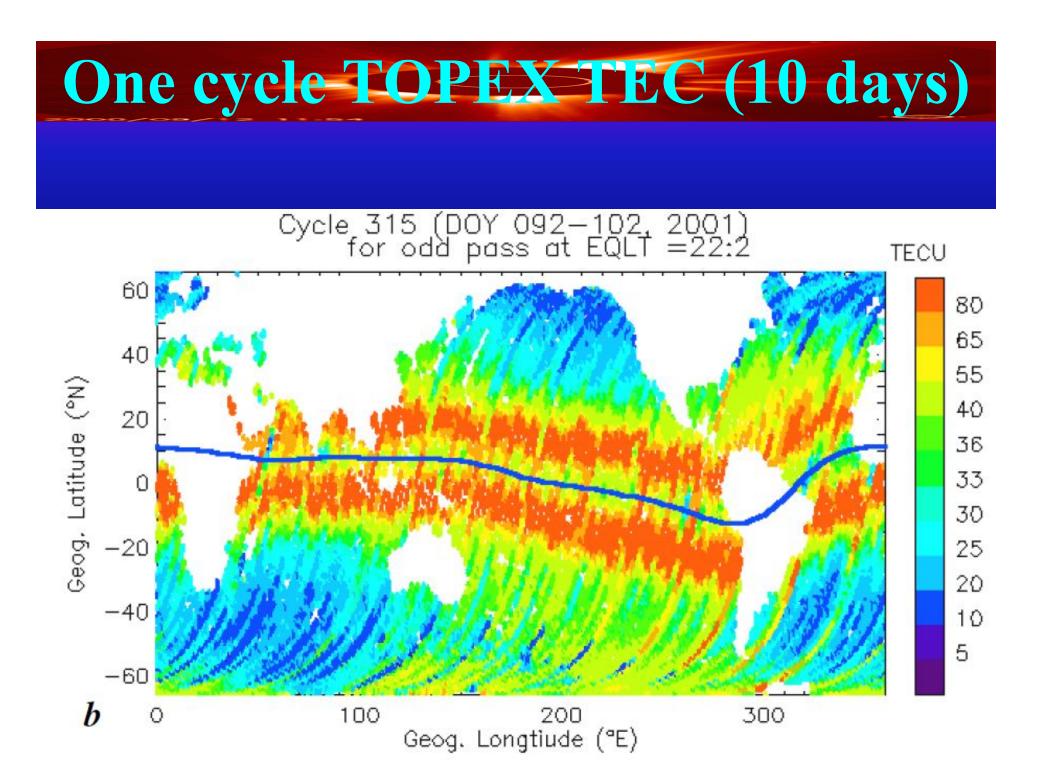
Equatorial crossing longitudes

Topex passes start and end times

Topex TEC contour plots along its ground tracks

One cycle TOPEX Passes (10 days)





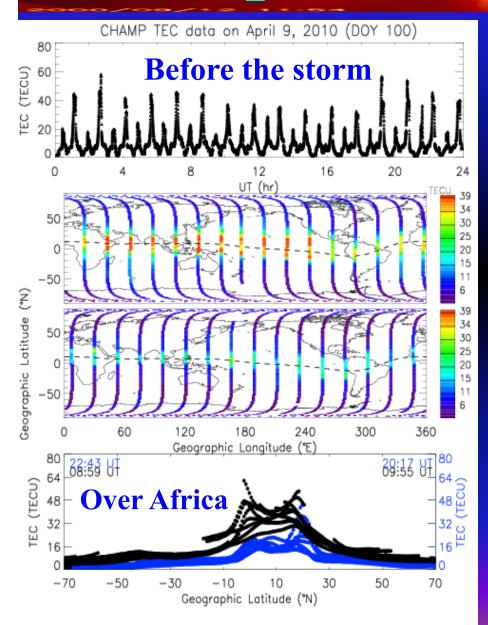
How about data from other LEO satellites?

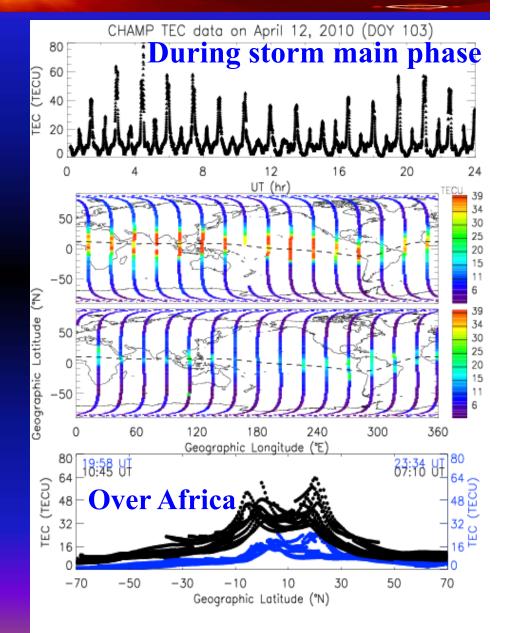
The COSMIC (six LEO satellites) and CHAMP LEO satellites provide two sets of very important data. (1) topside ionospheric and plasmaspheric GPS TEC (2) Occultation ionospheric density profiles (from about 70 km upto the altitude of the spacecraft).

COSMIC and CHAMP data base

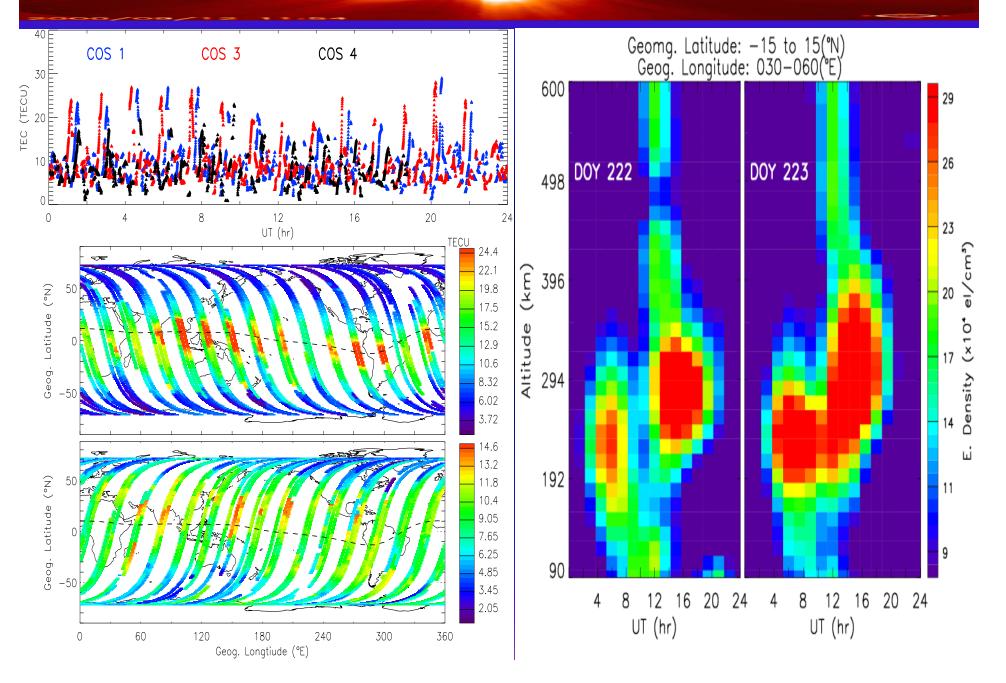
http://cosmic-io.cosmic.ucar.edu/cdaac/

Examples of LEO TEC data





Examples of LEO TEC density profile data



More In situ density and drift observation also avaliable from various space crafts?

DMSP satellite also provide various important data sets, including in situ density.

MDMASP database (organized by UT Dallas)

http://cindispace.utdallas.edu/DMSP/dmsp_data_at_utdallas.html

More space weather data can also be available

Here are some of the publicly available data bases.

Coordinated Data Analysis Web (CDAWeb)

http://cdaweb.gsfc.nasa.gov/cdaweb/istp_public/

OMNI data base

http://spdf.gsfc.nasa.gov/data_orbits.html

Where to find model run outputs?

Here are some of the publicly available model run data bases.

Community Coordinated Modeling Center (CCMC)

http://ccmc.gsfc.nasa.gov/

Thank you!

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