Ionospheric Storm Effects in GPS Total Electron Content

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Geomagnetic Storms





Figure: Hourly *Dst* index values for the month of September 2011.

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- A geomagnetic storm is characterized by a disturbance in the horizontal (*H*) component of the Earth's magnetic field at the equator due to the changing intensity of the ring current [*Gonzalez et al.*, 1994].
- The Disturbance storm time (*Dst*) index is an hourly measure of the average global *H* variation obtained from low-latitude ground stations and is often used to indicate the occurrence of geomagnetic storms.

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Ionospheric Storm Effects ^{WVirginiaTech}

• The ionospheric electron density response to a geomagnetic storm is traditionally classified as either a positive (increase) or negative (decrease) storm effect [*Prölss*, 1995; *Buonsanto*, 1999].

• Initial positive storm effects are generally attributed to an uplift in the ionospheric *F* layer to regions of decreased recombination.

• Longer lived negative storm effects are attributed to neutral composition changes, specifically an increase in molecular gas (O_2 , N_2) and decrease in atomic oxygen density.



Motivation





• One increasingly popular approach for describing ionospheric behavior is by using measurements of vertically-integrated total electron content (TEC) from ground-based GPS receivers [*Mendillo*, 2006].

• The goal of this study is to statistically examine positive and negative storm effects at high spatial and temporal resolution using the densely-populated network of GPS receivers in North America to develop predictive capabilities.

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Motivation – WAAS







September 25th, 2011 19:51:57 UT

September 26th, 2011

19:51:57 UT

Figure: Wide Area Augmentation System (WAAS) availability over the U.S. before/after storm; small squares denote individual airport status; large squares denote grid ionosphere vertical error [*Wanner*, 2011].

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GPS Total Electron Content

• Original 5 min GPS TEC values are re-binned from $1^{\circ} \times 1^{\circ}$ geographic lat/lon cells to 30 min averages in a new $2^{\circ} \times 4^{\circ}$ geographic lat/lon grid over North America.



• These values are then compiled into a database spanning 13 years (2001-2013), comprising a total of 41,022,720 TEC measurements.

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Storm Identification





• Using *Dst* values from the World Data Center for Geomagnetism in Kyoto, we have identified all geomagnetic storms during the current solar cycle maximum (2010-2013) reaching a magnitude of at least -40 nT (total of 69 events).

• In this study, we use the time of storm main phase onset (sharp decrease to negative values) rather than the storm sudden commencement (initial storm signature).

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Superposed Epoch Analysis

• Instantaneous (TEC) and 27-day median (TECq) values were recorded in each geographic latitude/longitude bin for the interval ranging from 1 day before to 3 days after storm onset.

 The storm time change in total electron content (RTEC) defined as RTEC = (TEC – TECq)/TECq

was calculated for each bin and then organized by magnetic latitude (MLAT) and magnetic local time (MLT) [e.g. *Biqiang et al.*, 2007].

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Superposed Epoch Analysis



• Finally, median values of RTEC within each MLAT/MLT bin were then calculated using data from all storm events.

• Here we use a relative change (RTEC) rather than an absolute difference because of dependence of the background TEC magnitude on solar cycle variability.

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Results – Global RTEC





Key Features:

- Positive phase spanning 06-24 MLT observed at all latitudes at storm onset
- Negative phase starts in dawn sector within 6 hours
- Complete change from positive to negative phase at all MLTs within 24 hours
 - Auroral oval seen as early positive response from 24-06 MLT at high latitudes
- Trough seen as negative response from 18-24 MLT about 4-24 hours after onset

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Results – RTEC at 12 MLT





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Future Work – Storm List



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Future Work – EOF Model

First EOF Basis Function E1 0.06 0.07 0.07 0.08 0.08 0.08 0.09



1st Basis Function: variation with geomagnetic latitude and solar EUV

 $E_1 \times P_1 = 94.28\%$ of variance

Second EOF Basis Function E₂ -0.11-0.08-0.04 0.00 0.04 0.08 0.11



2nd Basis Function: variation with declination (neutral winds) Third EOF Basis Function E₃ -0.11-0.08-0.04 0.00 0.04 0.08 0.11

3rd Basis Function: variation with dipole tilt offset + geomagnetic latitude?

 $E_3 \times P_3 = 0.61\%$ of variance

$$E_2 \times P_2 = 3.90\%$$
 of variance

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Future Work – EOF Model

First EOF Basis Function E1 0.06 0.07 0.07 0.08 0.08 0.08 0.09



1st Basis Function: variation with geomagnetic latitude and solar EUV

 $E_1 \times P_1 = 94.28\%$ of variance

Second EOF Basis Function E2 -0.11-0.08-0.04 0.00 0.04 0.08 0.11



2nd Basis Function: variation with declination (neutral winds)

 $E_2 \times P_2 = 3.90\%$ of variance

Fourth EOF Basis Function E4 -0.15-0.10-0.05 0.00 0.05 0.10 0.15

4th Basis Function: "Kansas Anomaly" (bad GPS receiver biases in MIT data processing)

 $E_4 \times P_4 = 0.36\%$ of variance (cumulative variance = 99.15%)

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Summary



• We have performed a superposed epoch analysis of 69 geomagnetic storms from the recent solar cycle maximum period (2010-2013) to gain a better understanding of the average GPS TEC response in the North American sector.

• A new effort is underway to automatically identify geomagnetic storm onset times and durations for the full 2001-2013 period to improve statistics of seasonal / local time / etc. effects.

• Future work will also include modeling of TEC storm effects using Empirical Orthogonal Function (EOF) techniques for potential predictive or operational capabilities.

References



- Biqiang, Z., W. Weixing, L. Libo, and M. Tian (2007), Morphology in the total electron content under geomagnetic disturbed conditions: Results from global ionosphere maps, *Ann. Geophys.*, 25, 1555–1568.
- Buonsanto, M. J. (1999), Ionospheric Storms A Review, Space Sci. Rev., 88, 563–601.
- Gonzalez, W. D., J. A. Joselyn, Y. Kamide, H. W. Kroehl, G. Rostoker, B. T. Tsurutani, and V. M. Vasyliunas (1994), What is a geomagnetic storm?, *J. Geophys. Res.*, 99(A4), 5771–5792.
- Mendillo, M. (2006), Storms in the ionosphere: Patterns and processes for total electron content, *Rev. Geophys.*, 44, RG4001.
- Prölss, G. W. (1995), Ionospheric F-region storms, in *Handbook of Atmospheric Electrodynamics*, vol. 2, edited by H. Volland, chap. 8, pp. 195-248, CRC Press, Boca Raton, Fla.
- Rideout, W., and A. Coster (2006), Automated GPS processing for global total electron content data, *GPS Solutions*, 10(3), 219–228.
- Wanner, B. (2011), DR 104 WAAS Reaction to Iono Activity September 26 2011, *William J Hughes Technical Center WAAS Test Team*, http://www.nstb.tc.faa.gov/DisplayDiscrepancyReport.htm.



Thank You!

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