



## RELATIONS BETWEEN THE EQUATORIAL VERTICAL DRIFTS, ELECTROJET, GPS-TEC AND SCINTILLATION DURING THE 2008-09 SOLAR MINIMUM

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# Outlines

## ➤ Background

Equatorial Ionosphere: EEJ, EIA & EPB  
Ionospheric Parameters

## ➤ Methodology

Instruments/ Data Locations  
Artificial Neural Network

## ➤ Data Analysis

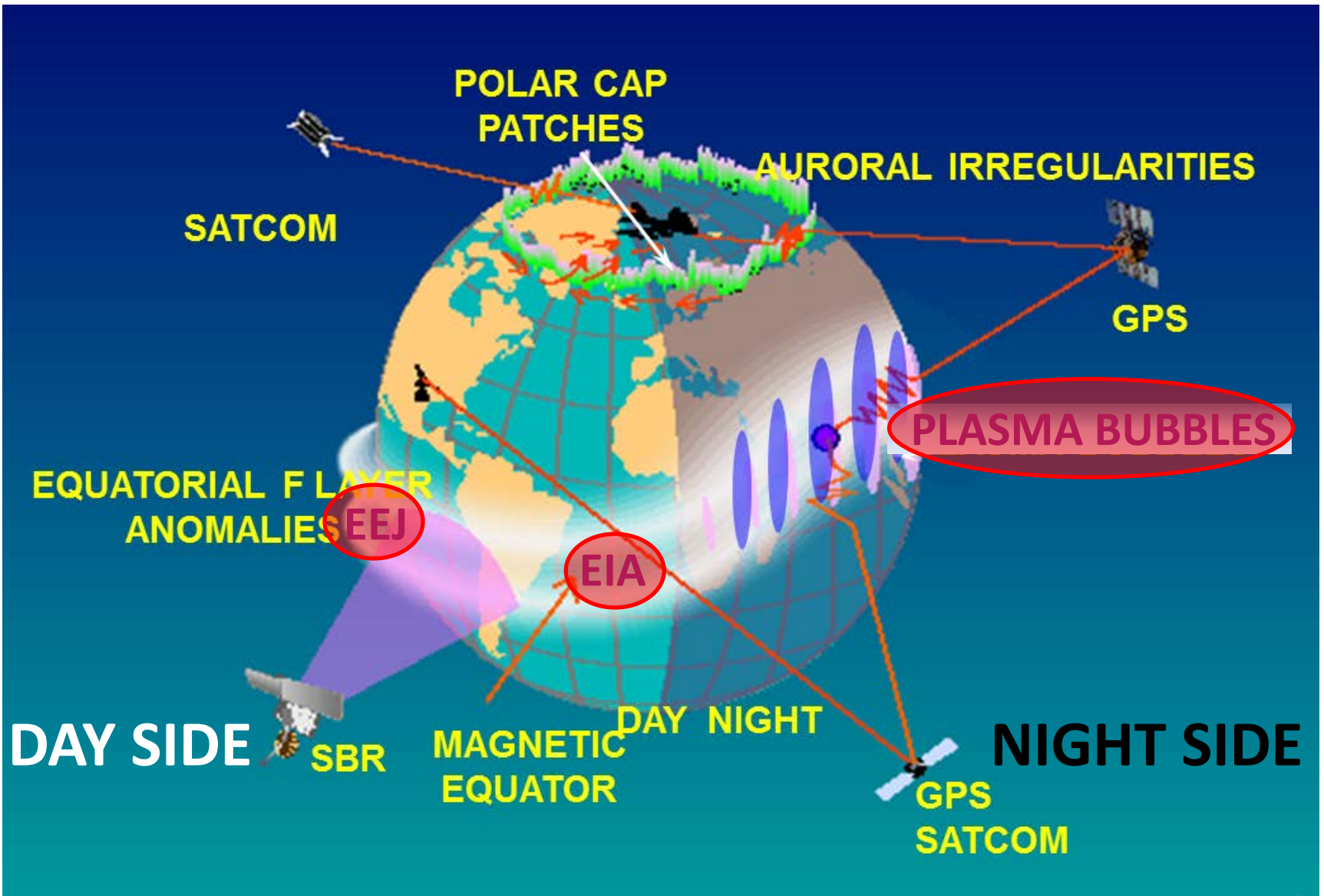
Data Selection  
Neural Network Inputs  
Scintillation Threshold

## ➤ Results

## ➤ Summary



# Background





# Ionospheric Parameters: TEC & S<sub>4</sub> Index

## Total Electron Content (TEC) Measurement

- Number of free electrons in a rectangular solid with a one-square-meter cross section extending from the receiver to the satellite.

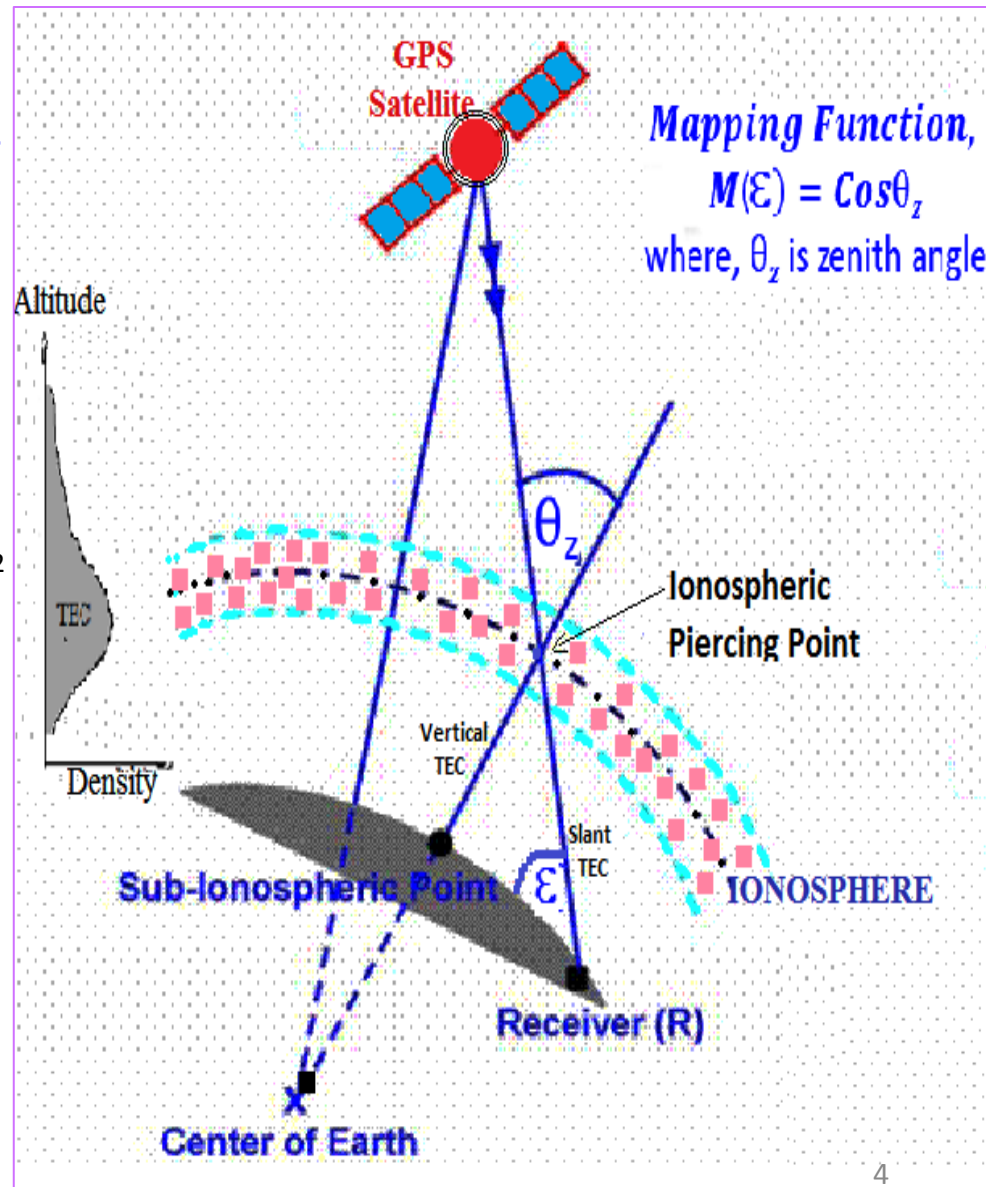
$$TEC = \int_{Receiver}^{Satellite} n(h) dh$$

$$1TECU = 10^{16} \text{ electrons/m}^2$$

## Ionospheric Scintillations Measurement

- Rapid fluctuation of the phase and intensity of signal that passed through ionosphere.
- **S<sub>4</sub> index**: Normalized standard deviation of signal intensity,

$$S_4 = \frac{\sqrt{(\langle I^2 \rangle - \langle I \rangle^2)}}{\langle I \rangle}$$



# Methodology

- **Sensor Techniques**

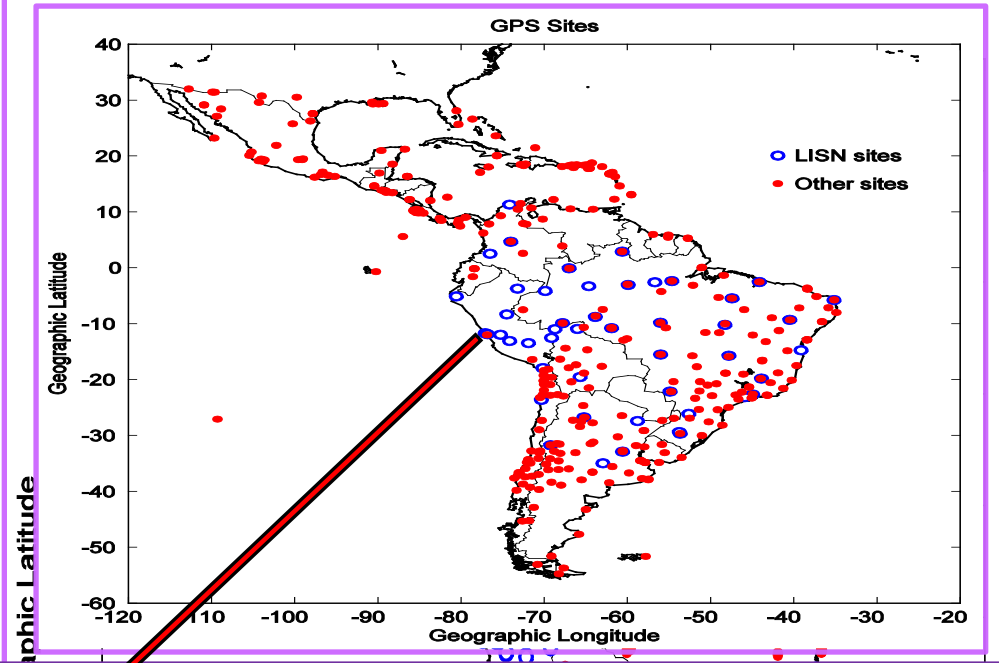
- Magnetometer

- **Radio Techniques**

- Radars (CS, IS & UHF)

- GPS

- Ionosondes



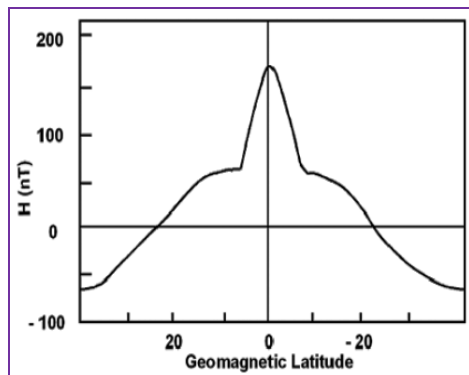
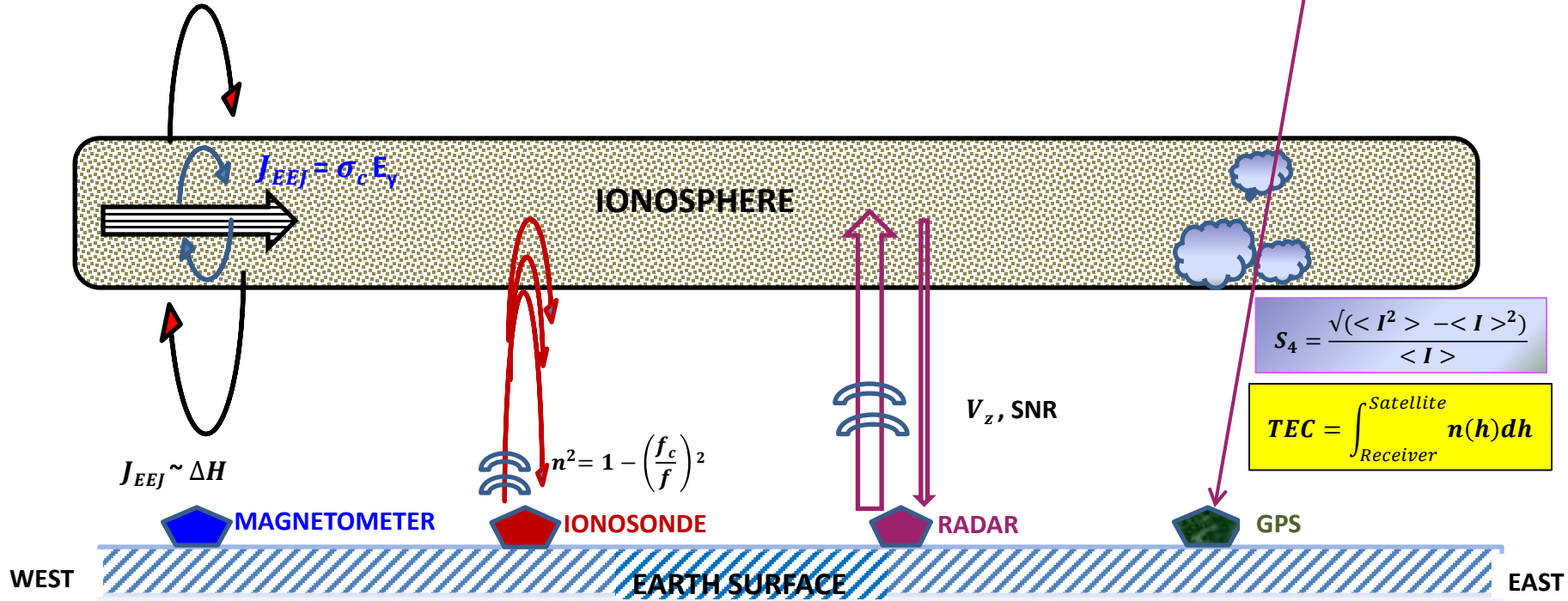
CSR: Coherent Scatter Radar

ISR: Incoherent Scatter Radar

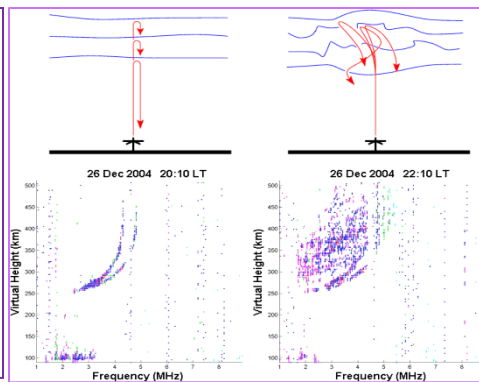
**Jicamarca Radio Observatory, Peru @ Magnetic Equator of the Earth**



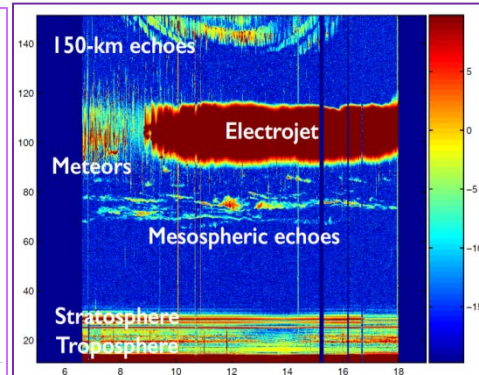
# Ground Based Observations of Space Dynamics



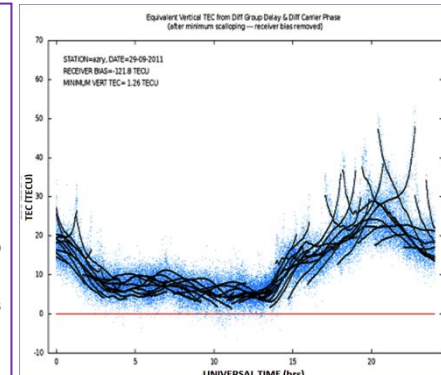
H ENHANCEMENT AT NOON



IONOGRAM



RADAR



TOTAL ELECTRON CONTENT



# Artificial Neural Network Approach

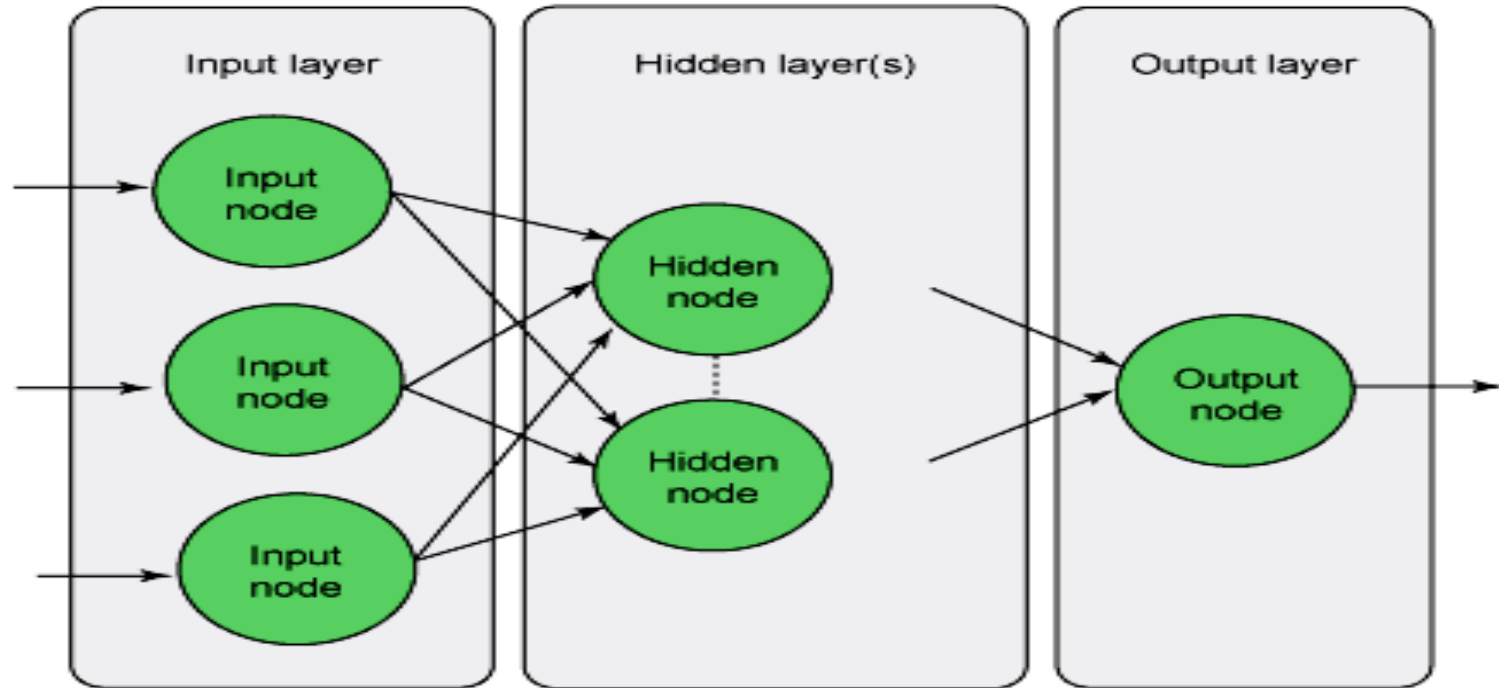
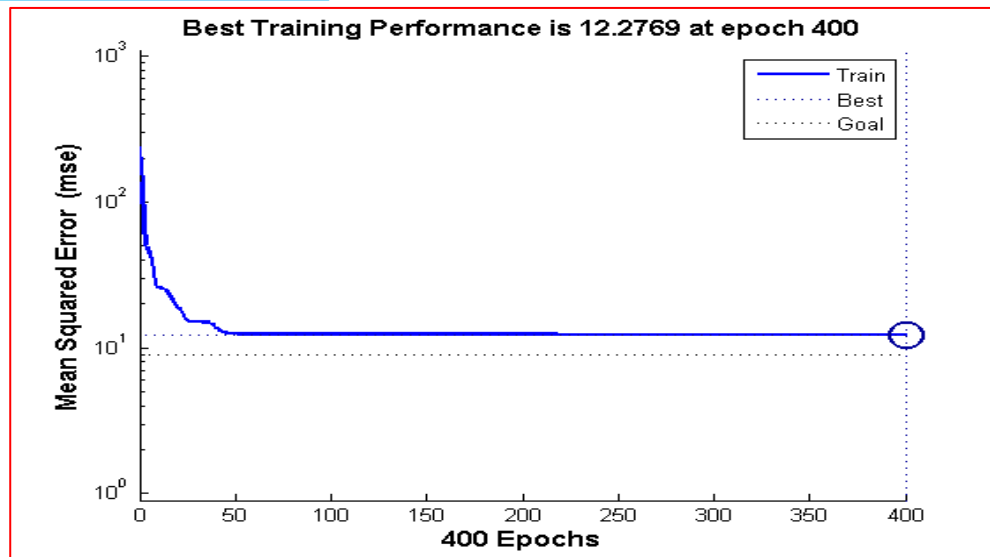


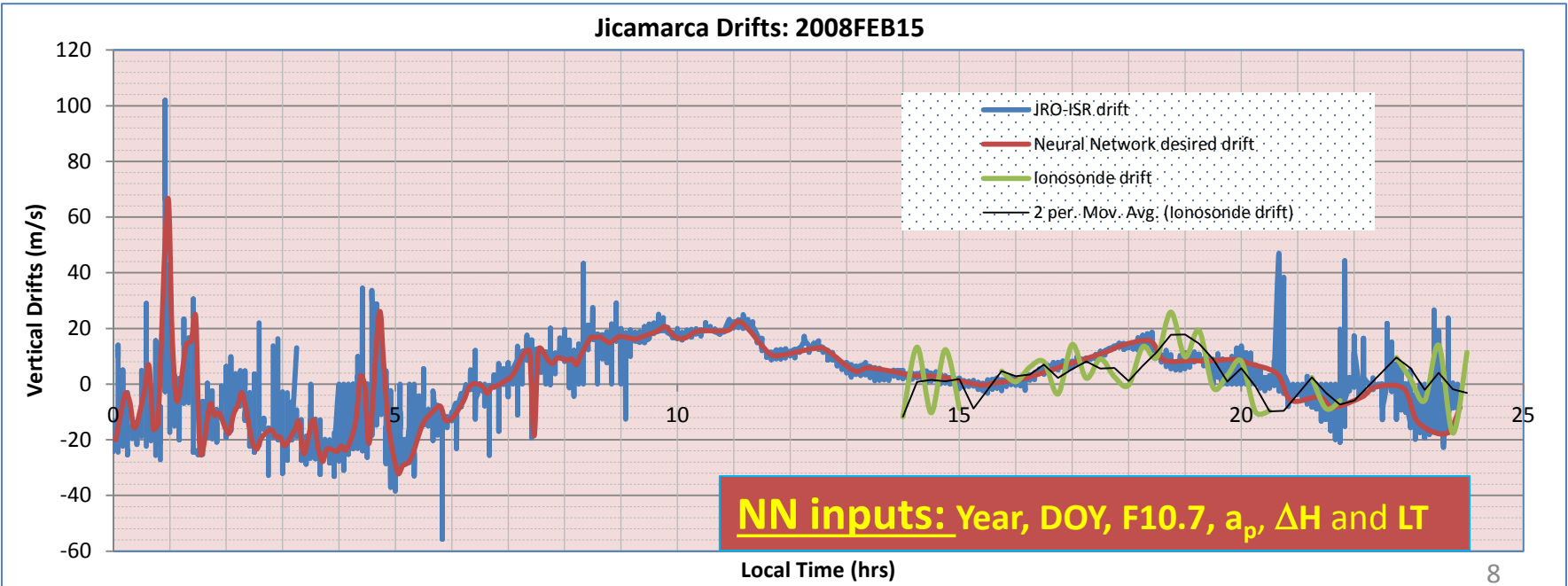
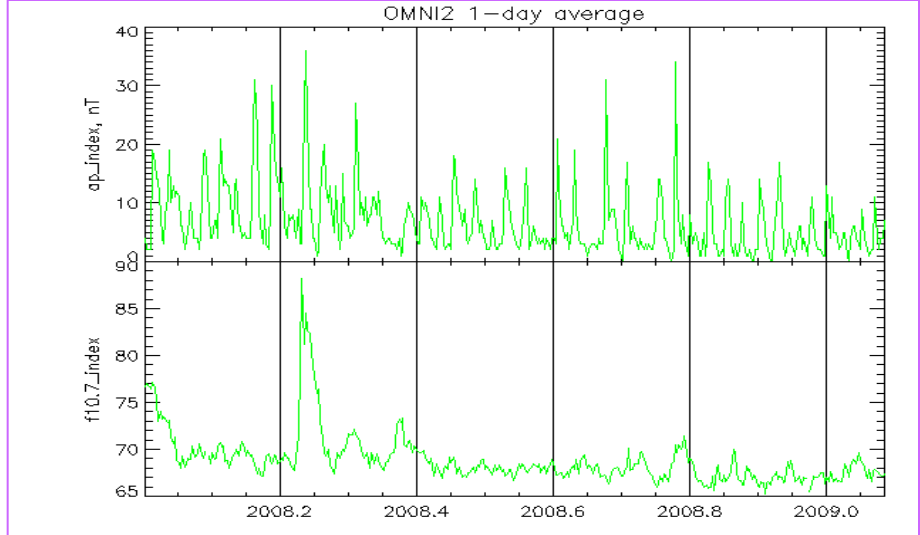
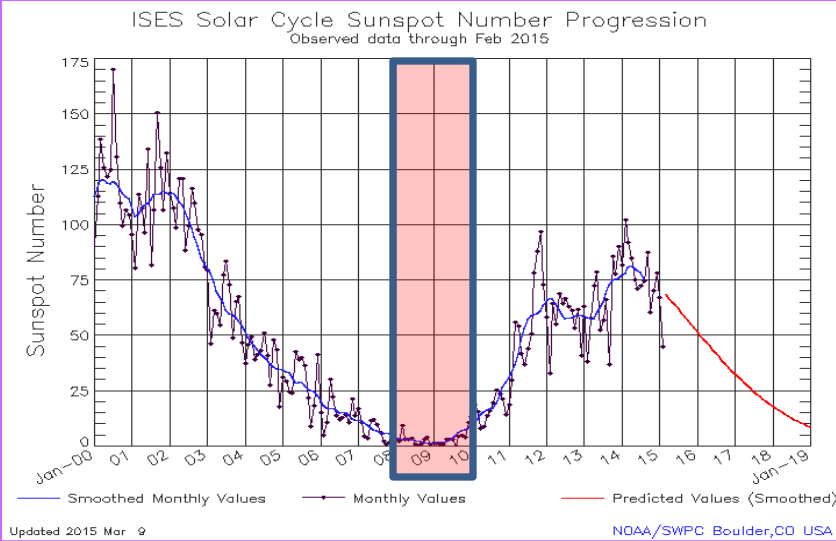
Fig.: Neural network schematic

| Least Squares Method (LSM) and Neural Network (NN)  |                |
|---|----------------|
| Approach  | RMS Error, m/s |
| $E \times B$ (LSM) = $a_0 + a_1\Delta H$  | 3.82           |
| $E \times B$ (NN) = 1 input ( $\Delta H$ )  | 3.76           |
| $E \times B$ (LSM) = $a_0 + a_1\Delta H + a_2\Delta H^2 + a_3\Delta H^3$                                | 3.79           |
| $E \times B$ (NN) = 1 input ( $\Delta H$ )  | 3.76           |
| $E \times B$ (LSM) = $a_0 + a_1\Delta H + a_2\Delta H^2 + a_3\Delta H^3 + a_4LT$                        | 3.75           |
| $E \times B$ (NN) = 2 inputs ( $\Delta H, LT$ )   | 3.75           |
| $E \times B$ (LSM) = $a_0 + a_1\Delta H + a_2\Delta H^2 + a_3\Delta H^3 + a_4Kp + a_5Ap + a_6LT$        | 3.68           |
| $E \times B$ (NN) = 4 inputs ( $\Delta H, LT, Ap, Kp$ )   | 3.67           |
| $E \times B$ (LSM) = $a_0 + a_1\Delta H + a_2\Delta H^2 + a_3\Delta H^3 + a_4DOY + a_5year + a_6LT$     | 3.39           |
| $E \times B$ (NN) = 4 inputs ( $\Delta H, LT, year, DOY$ )  | 3.24           |
| $E \times B$ (LSM) = $a_0 + a_1\Delta H + a_2\Delta H^2 + a_3\Delta H^3 + a_4F10.7 + a_5F10.7a + a_6LT$ | 3.25           |
| $E \times B$ (NN) = 4 inputs ( $\Delta H, LT, F10.7, F10.7A$ )  | 3.02           |

Anderson et al., 2004



# Neural Network Inputs





# Instruments & Data Locations

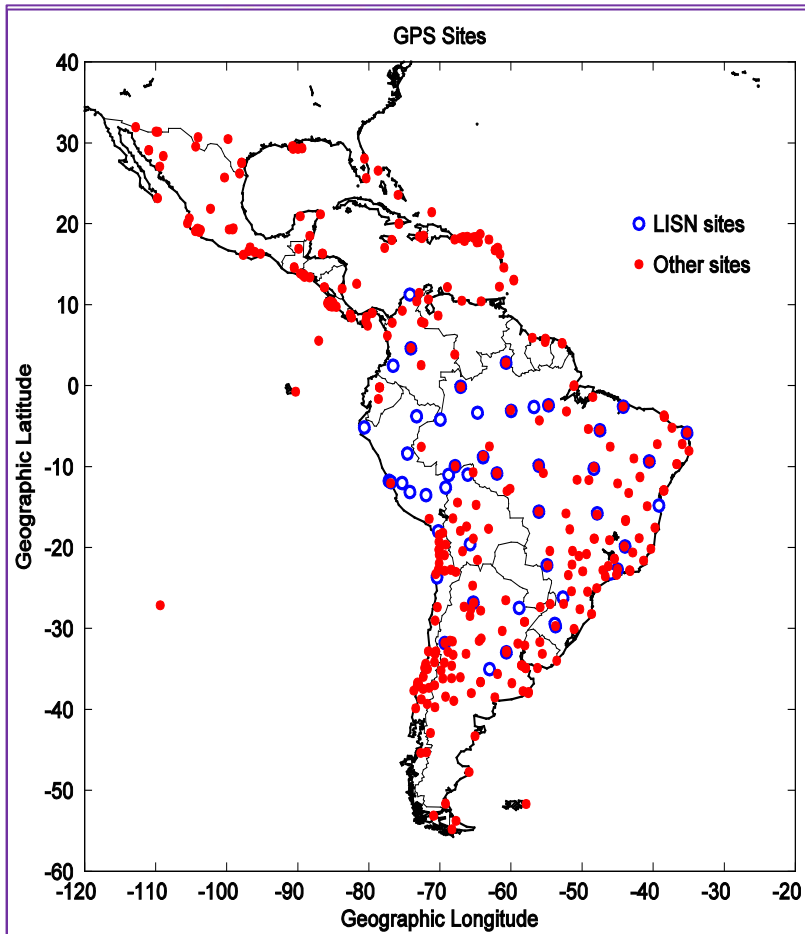
## Locations

**Jicamarca:-** Geographic:  $11.92^{\circ}\text{S}$ ;  $283.13^{\circ}\text{E}$

Geomagnetic:  $0.8^{\circ}\text{N}$

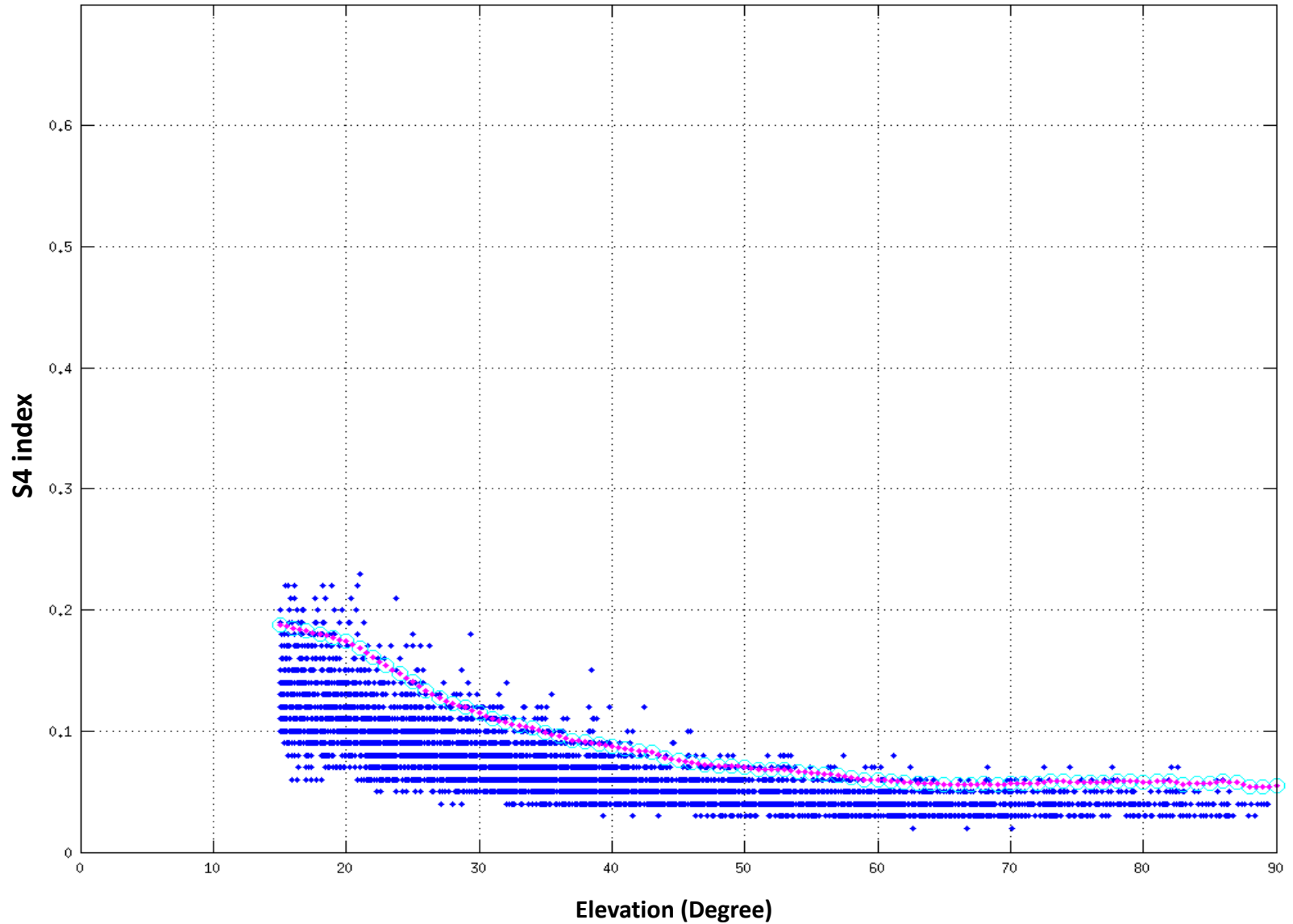
**Piura:-** Geographic:  $5.18^{\circ}\text{S}$ ;  $279.36^{\circ}\text{E}$

Geomagnetic:  $6.8^{\circ}\text{N}$

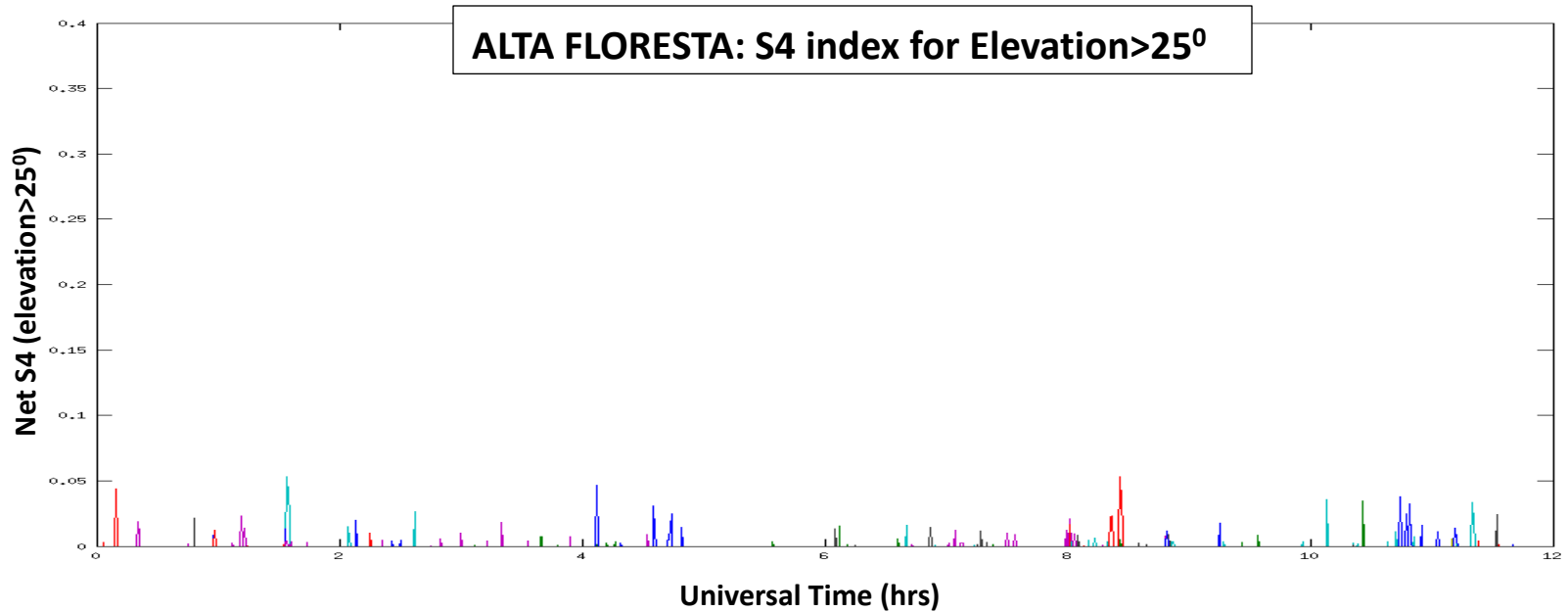
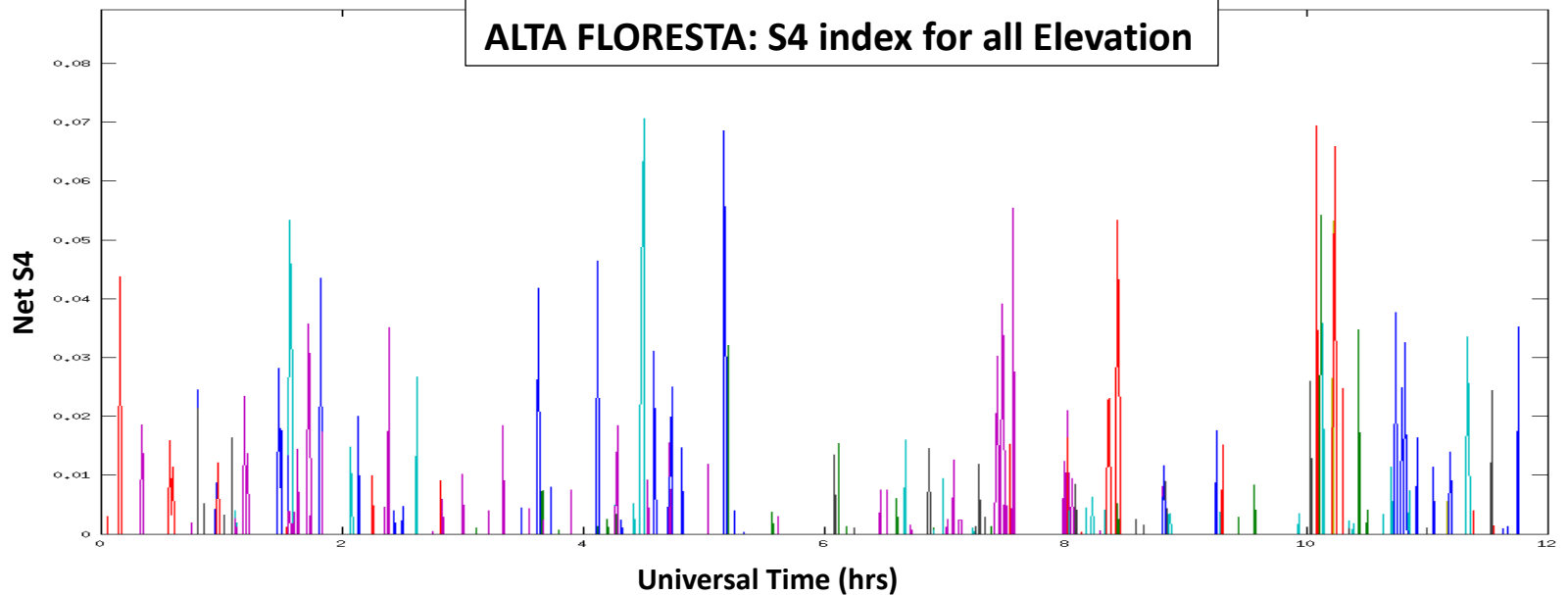


# Data Analysis

## Scintillation Threshold Model



# Net Scintillation

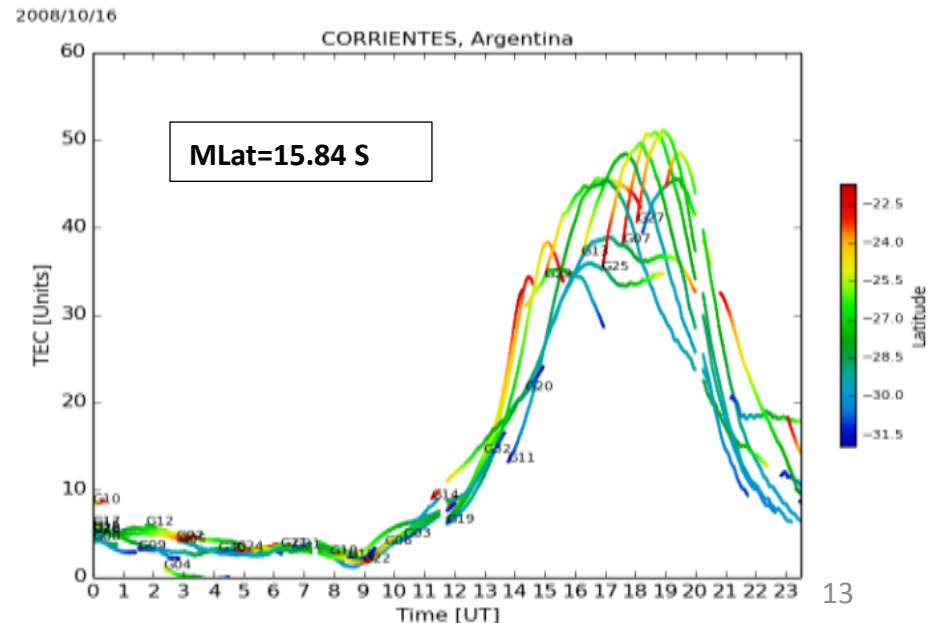
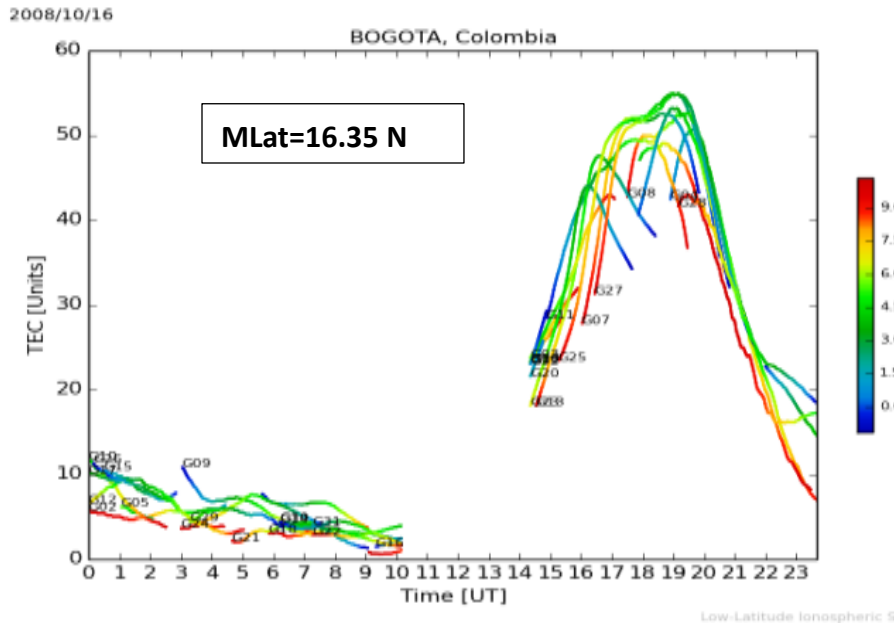
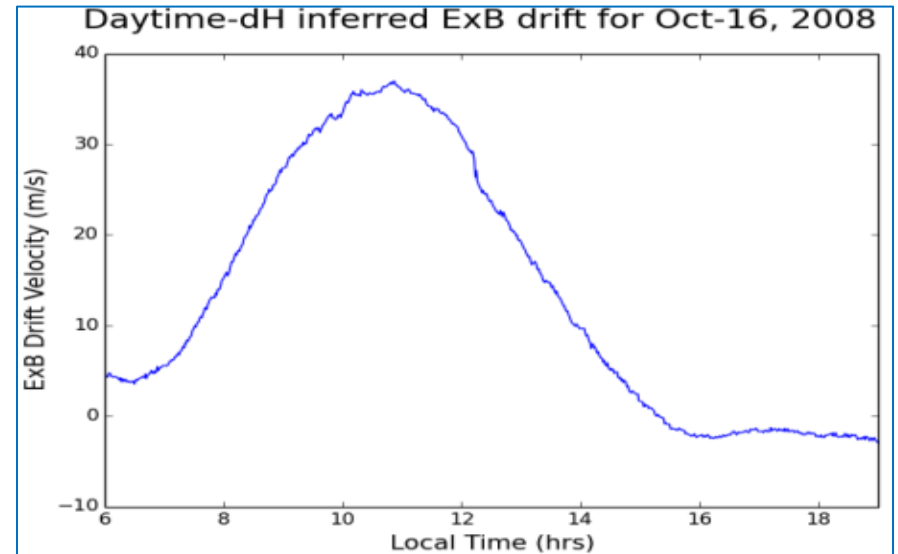
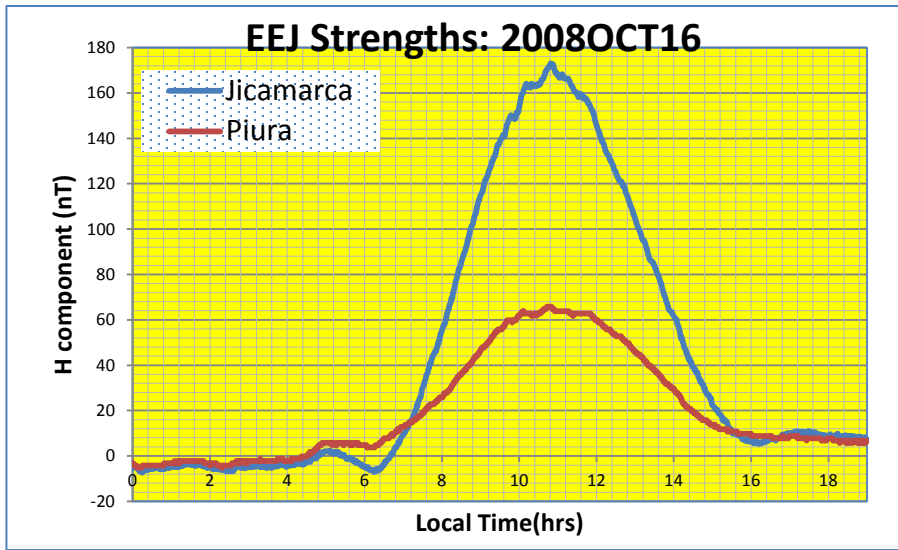


# Results

| Kp | Date     | EEJ Strength (nT)<br>(Jicamarca-Piura) | Vz,<br>m/s | TEC,<br>TECU         | GPS S4 index   | Ancon<br>UHF, S4 | EIA Location:<br>G.Latitude |
|----|----------|--|------------|----------------------|--|------------------|-----------------------------|
| <3 | OCT16'08 | @10.81LT<br>173.02-65.79=107.23        | ~37        | BOGT~ 55<br>CORR~ 51 | CUZC: 0.13; HUAN: 0.03<br>JICA: 0.10; PIUR: 0.11<br>PUER: 0.16; TACN: 0.13 | 0.2              | 34                          |
| <3 | SEP29'08 | @11.15LT<br>138.2-43.47= 94.73         | ~34        | BOGT~ 34<br>CORR~ 25 | CUZC: ; HUAN: 0.07<br>JICA: ; PIUR: 0.14<br>PUER: 0.09; TACN: 0.09         | 1                | 28                          |
| <3 | NOV07'09 | @11.76LT<br>89-59.92= 29.08            | ~17        | BOGT~ 35<br>CORR~ 33 | CUZC: ; HUAN:<br>JICA: ; PIUR:<br>PUER: ; TACN:                            |                  | 24                          |
| =4 | NOV08'08 | @9.53LT<br>57.92-49.17= 8.75           | ~5         | ANTF~ 28<br>CORR~ 25 | CUZC: 0.12; HUAN: 0.4<br>JICA: 0.2; PIUR: 0.4<br>PUER: 0.19; TACN:         | 0.88             | 10                          |
| =4 | JUL23'08 | @11.61LT<br>85.74-33.73= 52.01         | ~18        | BOGT~ 23<br>CORR~ 14 | CUZC: 0.08; HUAN: 0.07<br>JICA: ; PIUR:<br>PUER: 0.08; TACN: 0.05          | 0.03             | 15                          |
| =4 | JAN03'09 | @11.6LT<br>164.23-61.62= 102.61        | ~34        | BOGT~ 32<br>CORR~ 31 | CUZC: 0.16 ; HUAN:<br>JICA: 0.07 ; PIUR:<br>PUER: ; TACN:                  |                  | 28                          |
| >5 | MAR01'08 | @10.21LT<br>99.07-44.59= 54.48         | ~24        | BOGT~ 44<br>CORR~ 35 | CUZC: ; HUAN: 0.14<br>JICA: ; PIUR: 0.4<br>PUER: 0.12; PUCA: 0.07          | 1.1              | 29                          |
| >5 | APR06'08 | @11.21LT<br>159.32-43.66=115.66        | ~39        | IQTS~ 40<br>CORR~52  | CUZC: 0.1; ANCN: 0.37<br>JICA: 0.1; PIUR: 0.39<br>PUER: 0.07; TACN: 0.03   | 0.79             | 30                          |
| >5 | AUG18'08 | @12.6LT<br>138.7-58.556= 80.14         | ~31        | BOGT~ 37<br>CORR~ 22 | CUZC: 0.14 ; HUAN: 0.06<br>JICA: ; PIUR: 0.07<br>PUER: 0.07 ; TACN:        |                  | 27                          |



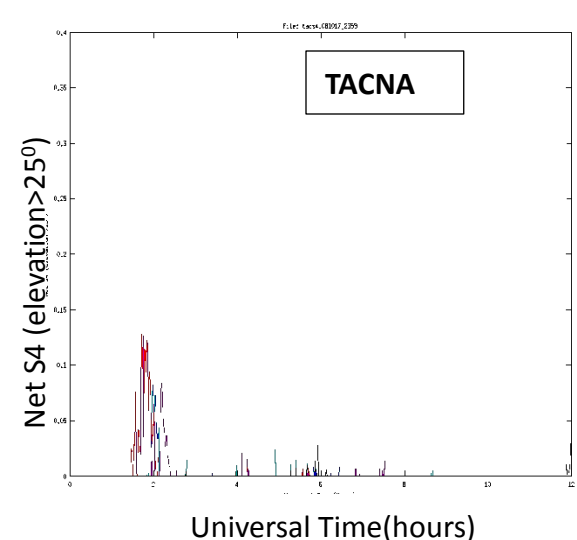
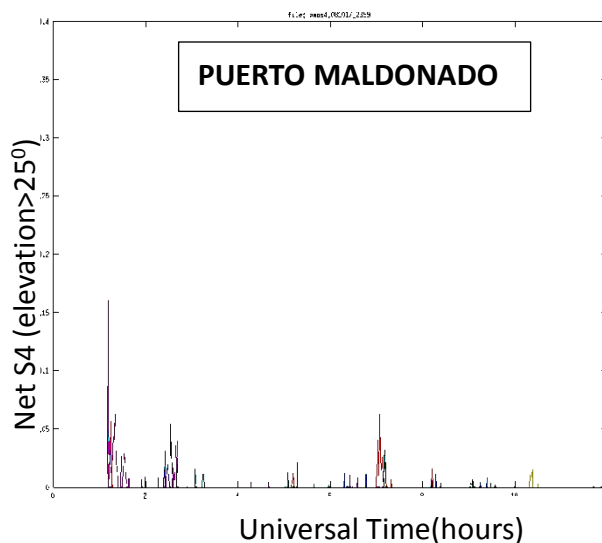
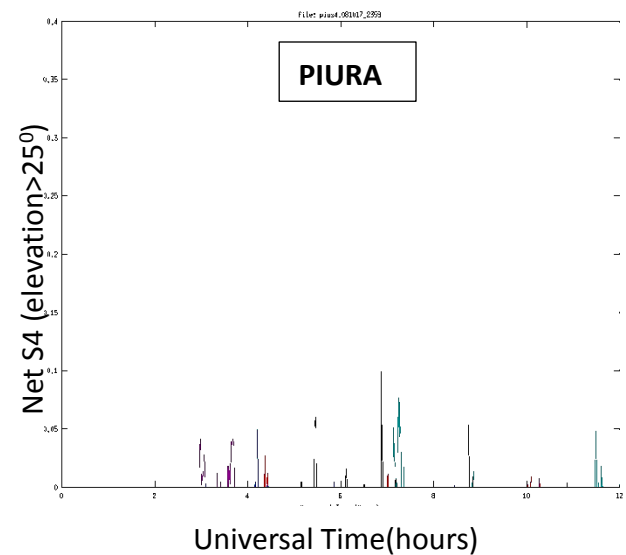
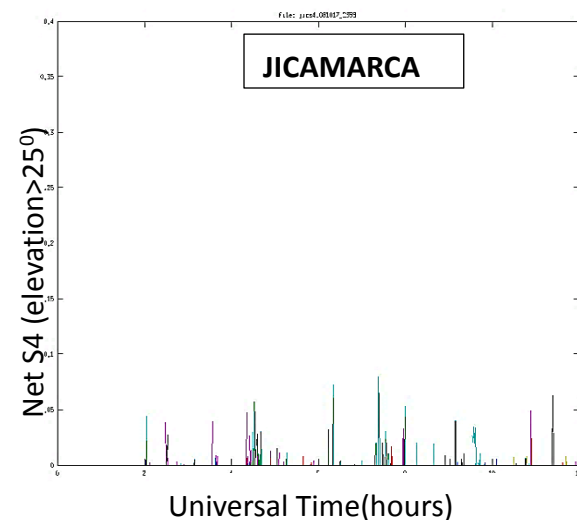
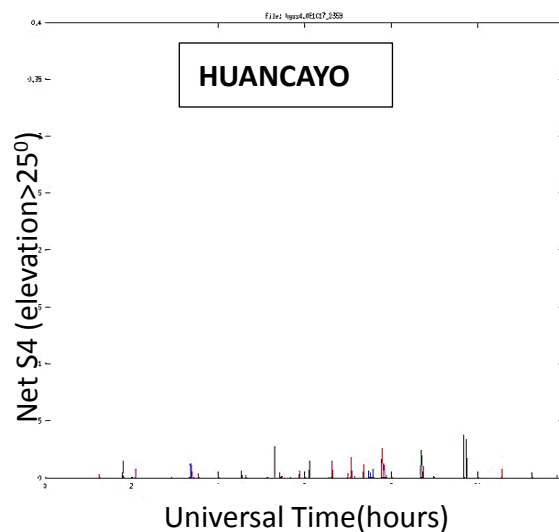
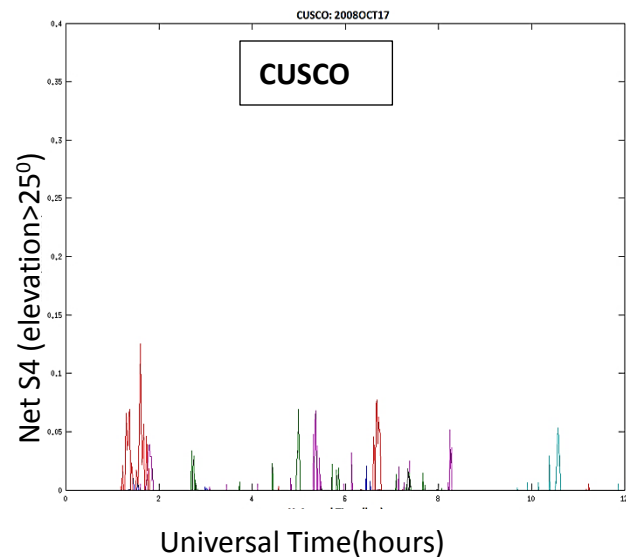
# Strong EEJ & Vertical ExB Drift: Quiet Day (Oct-16, 2008) $K_p < 3$



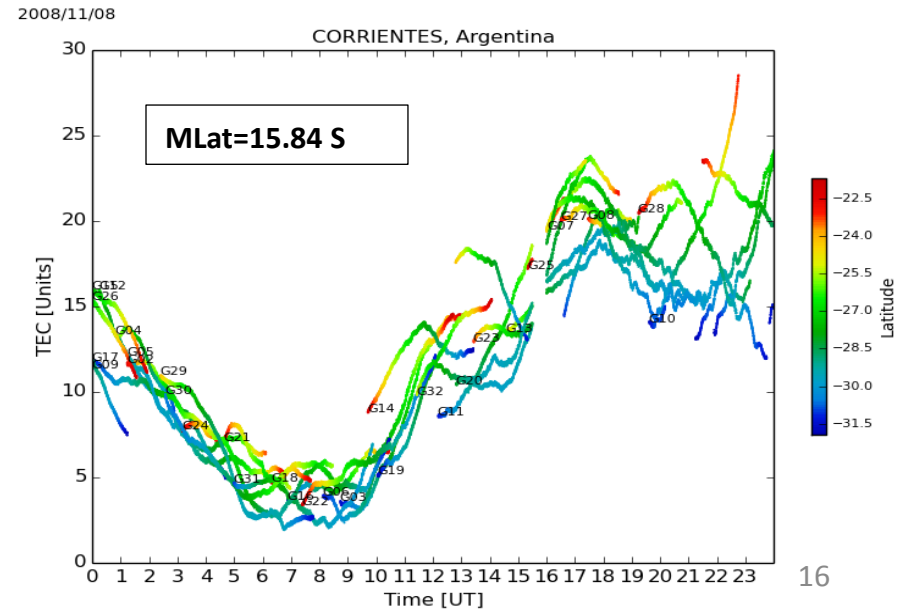
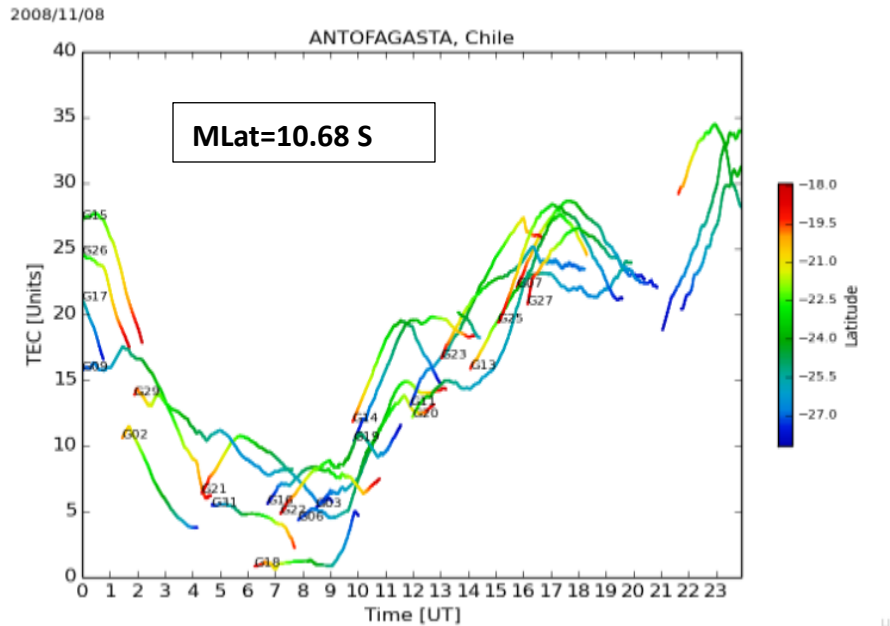
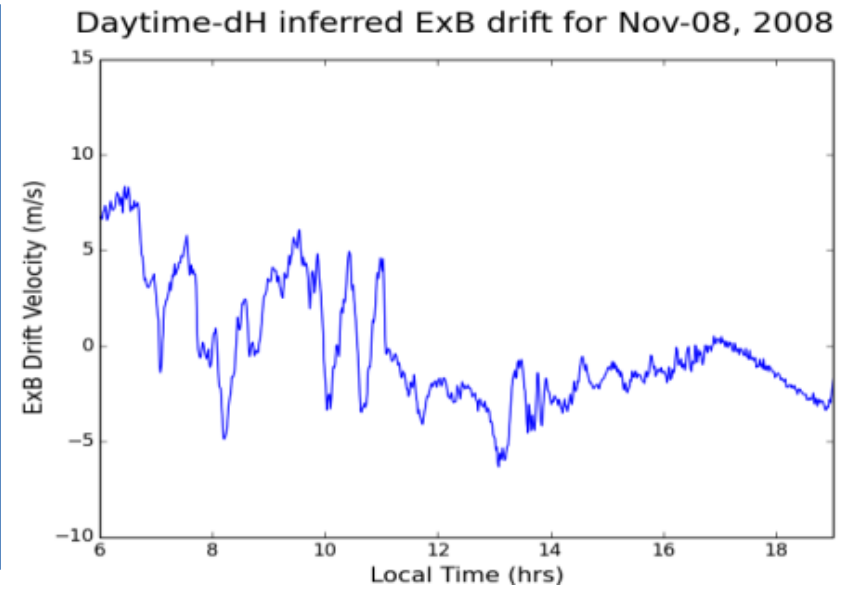
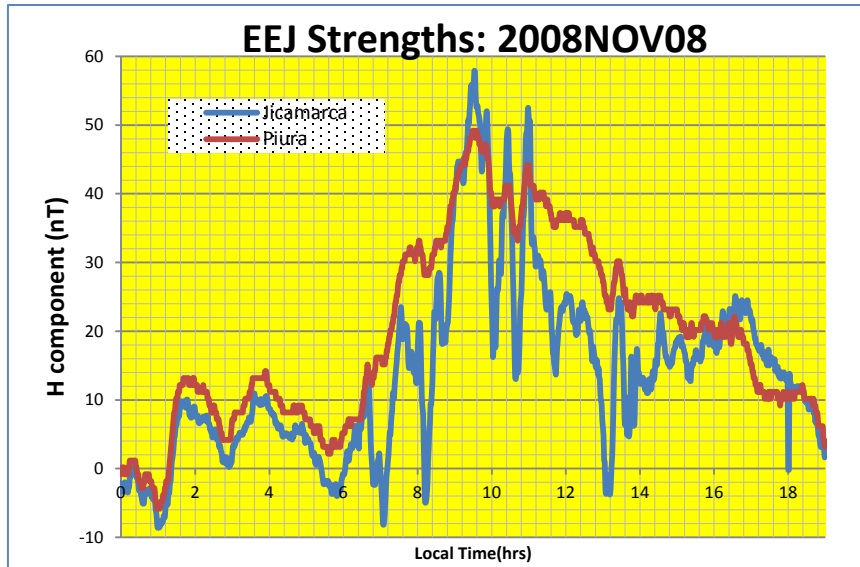


# Net Scintillation(Elevation>25°): Oct-17, 2008

(Strong EEJ & Vertical ExB Drift/ Quiet Day: Kp<3)



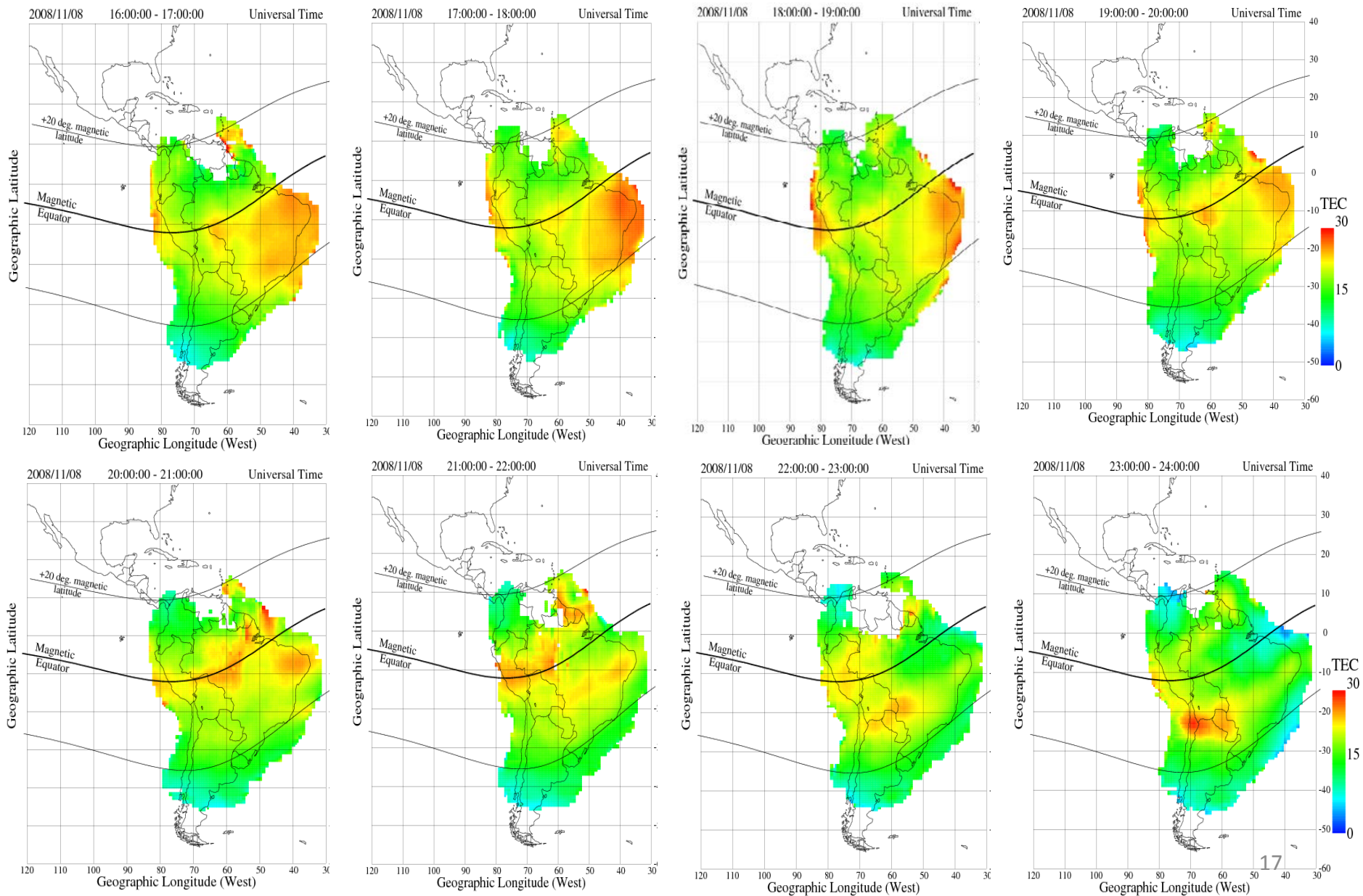
# Weak EEJ & Vertical ExB Drift: Non-Quiet Day (Nov-08, 2008) Kp=4





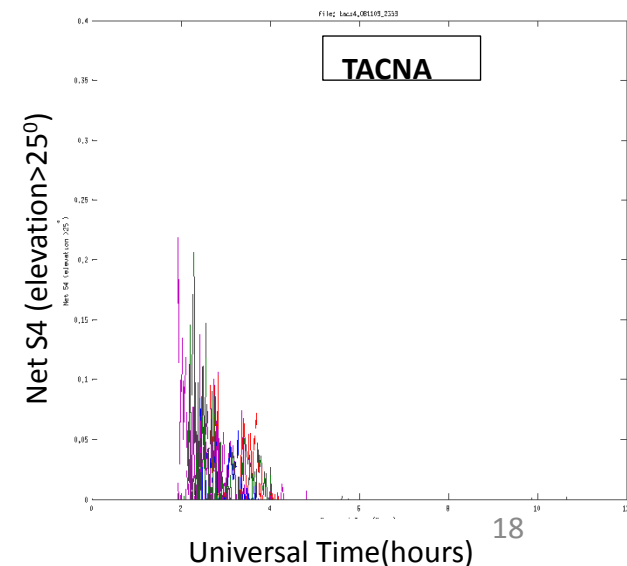
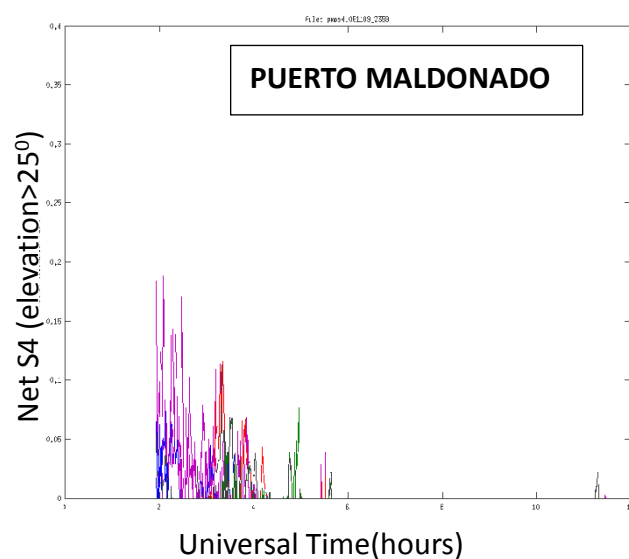
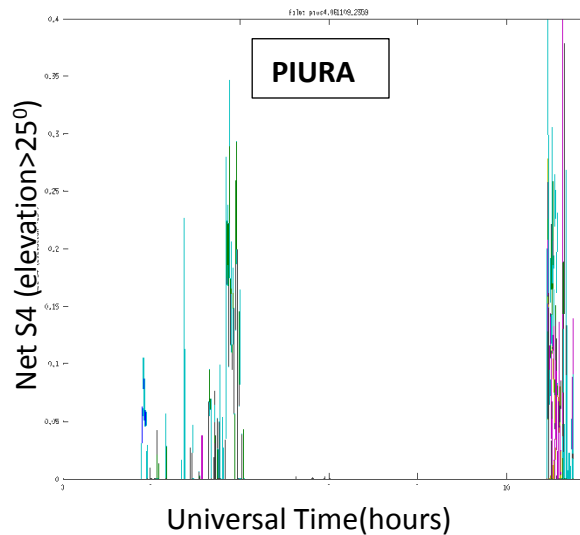
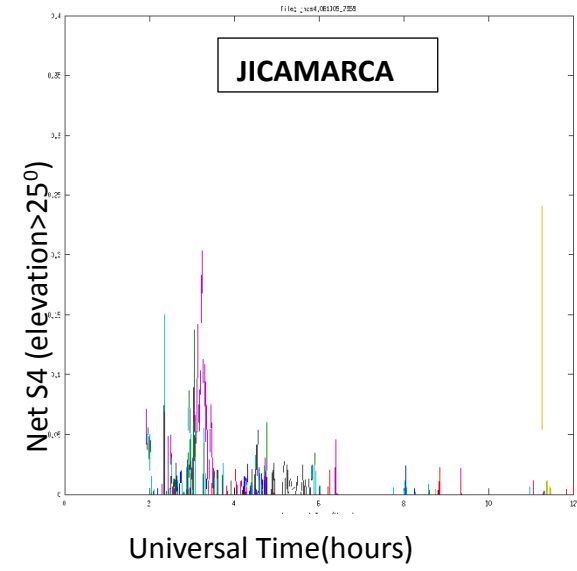
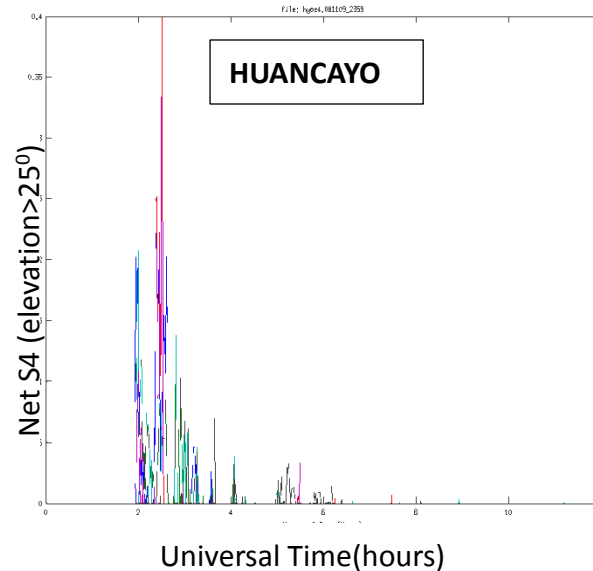
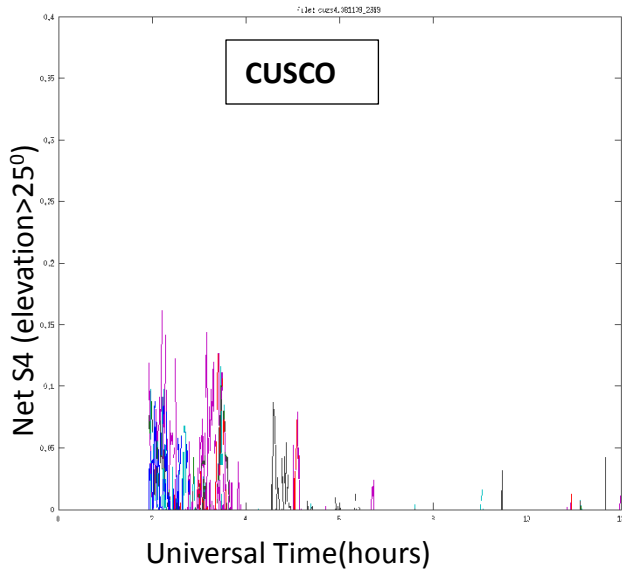
# TEC Fitted Map: Nov-08, 2008

(Weak EEJ & Vertical ExB Drift/ Non-Quiet Day: Kp=4)



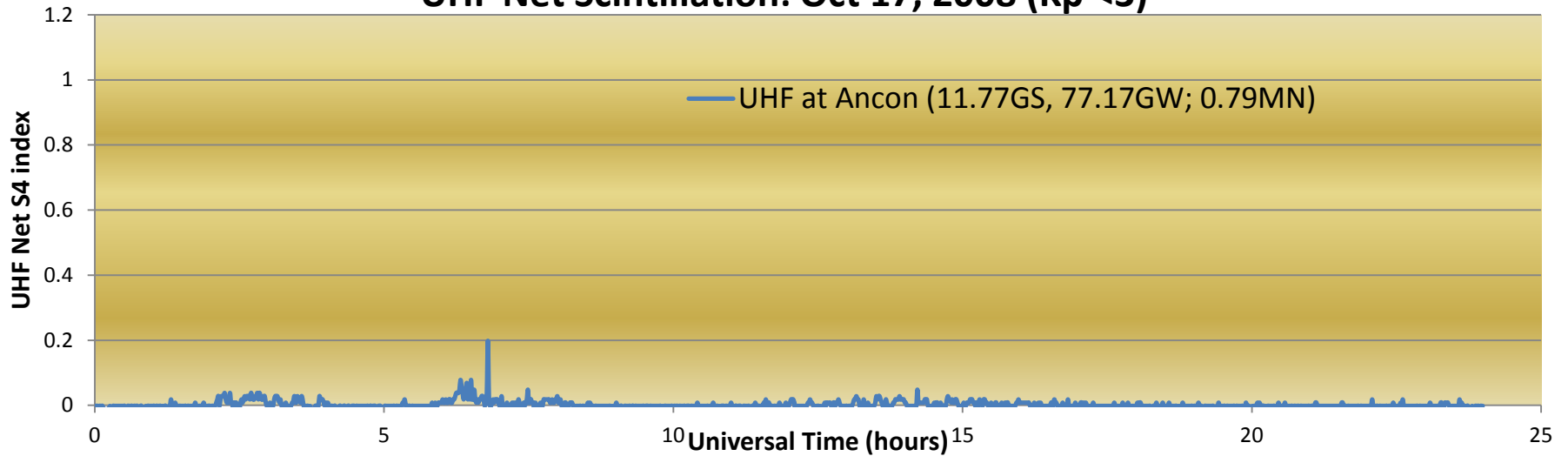
# Net Scintillation(Elevation>25°): Nov-09, 2008

(Weak EEJ & Vertical ExB Drift/ Non-Quiet Day: Kp=4)

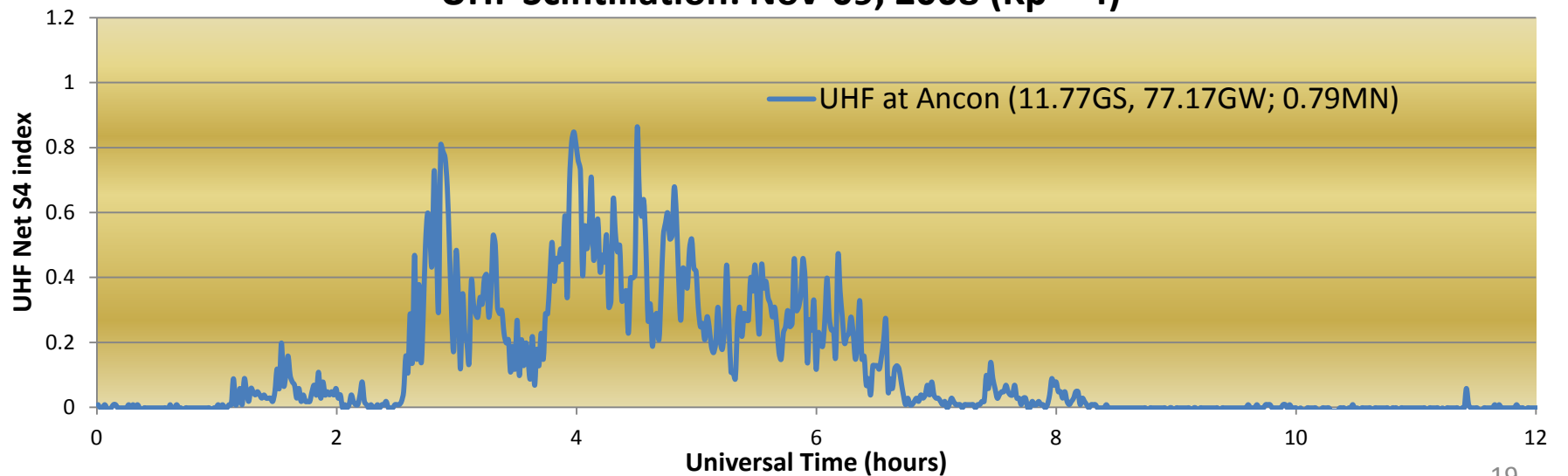


# Ancon UHF Scintillation

## UHF Net Scintillation: Oct 17, 2008 (Kp <3)

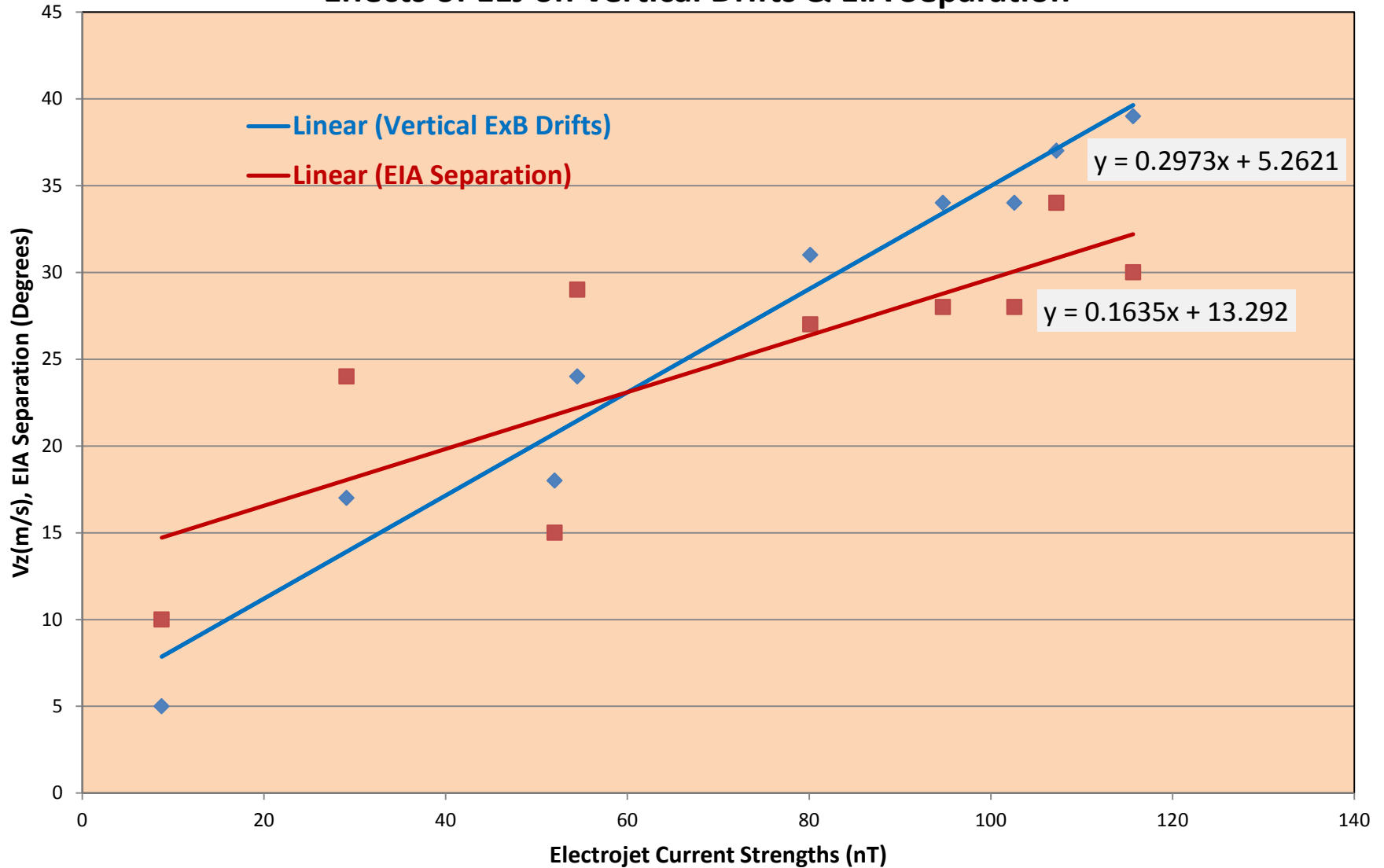


## UHF Scintillation: Nov-09, 2008 (Kp = 4)



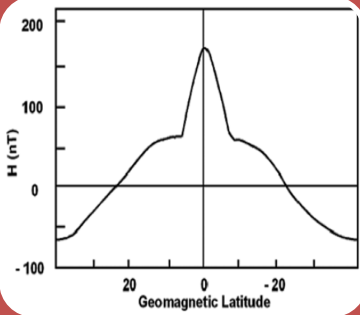
# Results

## Effects of EEJ on Vertical Drifts & EIA Separation

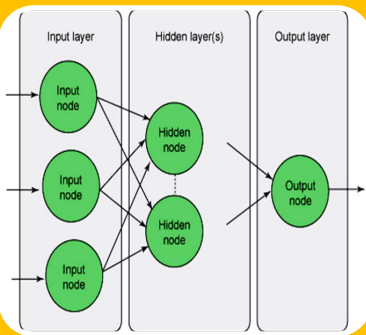




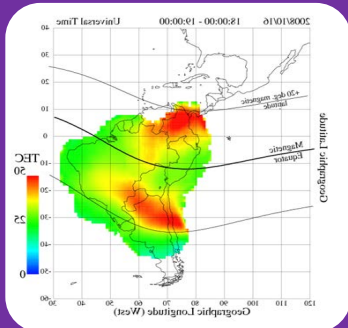
# Summary



There is a quantifiable relation between daytime vertical ExB drift and EEJ current. Hence, magnetometer observation can be used to infer vertical ExB drift.



Neural network technique with suitable inputs for the estimation of daytime vertical drift potentially gives better result than other approaches, based on the RMS errors comparison.



EEJ strengths and daytime vertical ExB drifts have noticeable connection to GPS-TEC. The EEJ strength controls the shape and location of EIA crests. But, their relation with the increase of the S4 index after sunset is not as obvious.

**Thank you!**