

Plasma Structures in the High-Latitude Ionosphere

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The High-Latitude Ionosphere

Highly Structured

Small scale structures associated with large scale structures

Large scale structures vary:

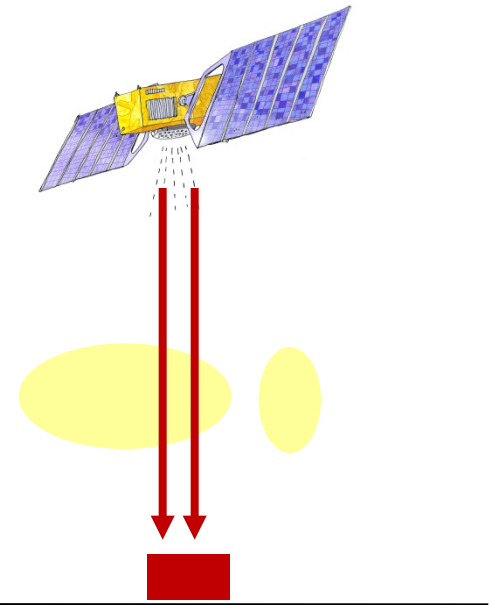
Season

Solar cycle

Geomagnetic activity

Solar wind conditions

Location, UT

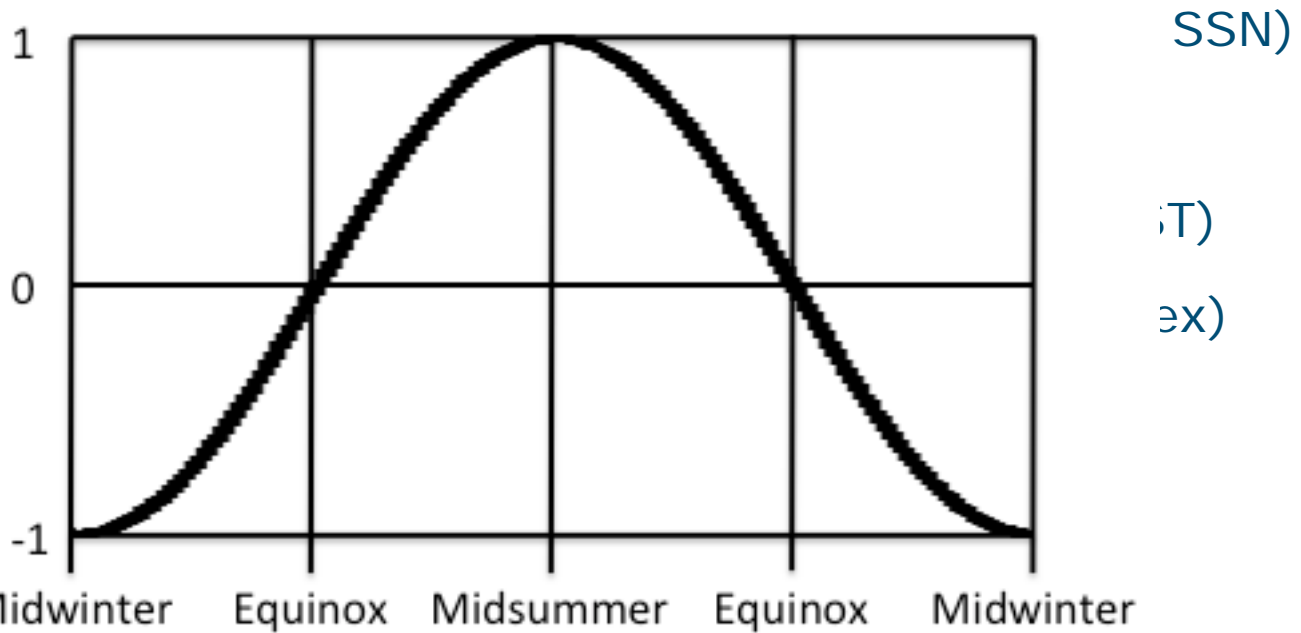


Relative importance is a more open question

Generalised Linear Modelling

Dependant Variable: Amount of plasma structuring

Independent



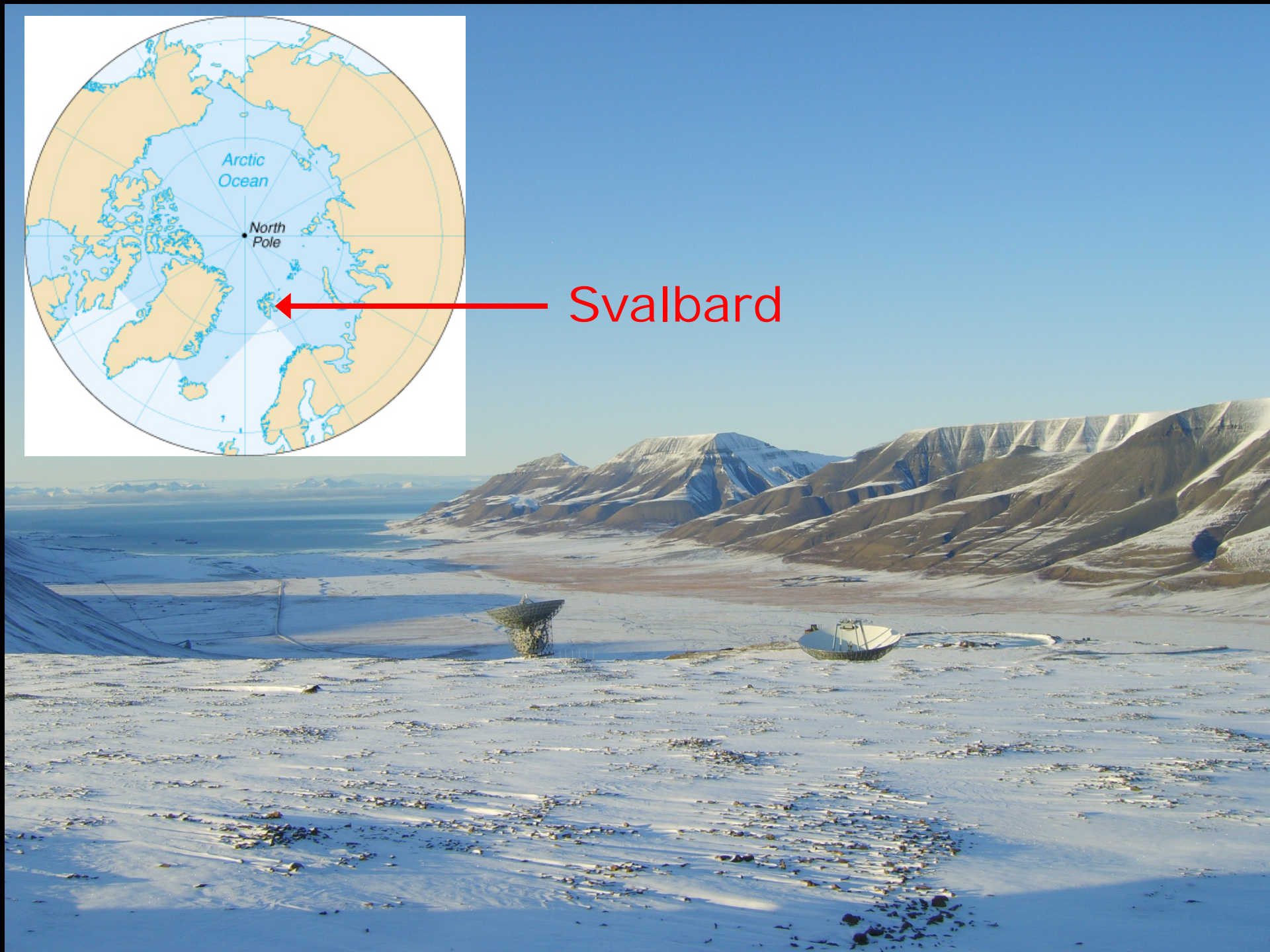
Modelling

Add most significant parameter

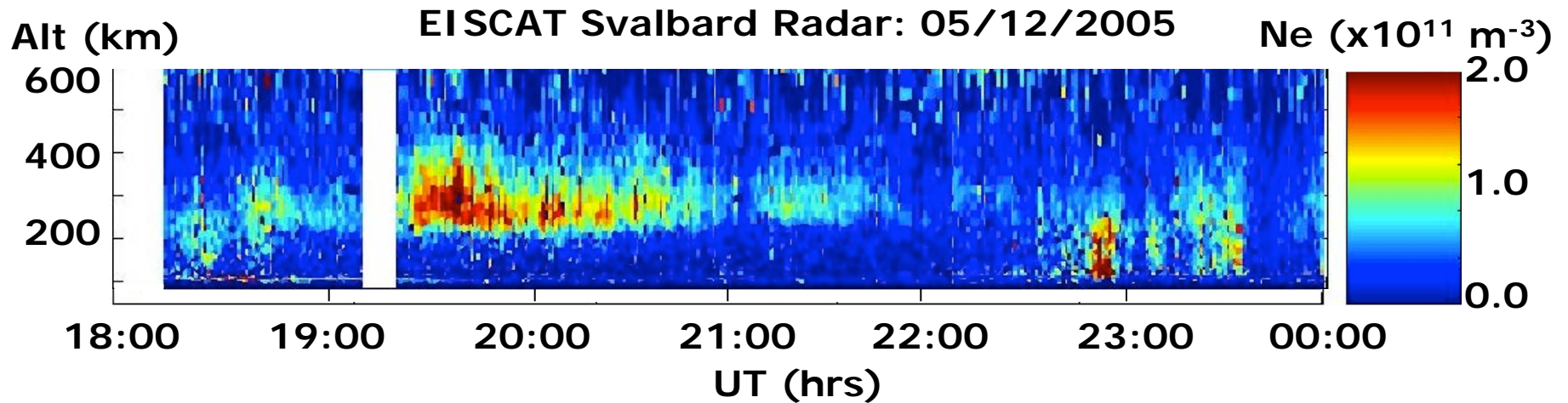
Continue until nothing else is significant at the 5 % level



Svalbard

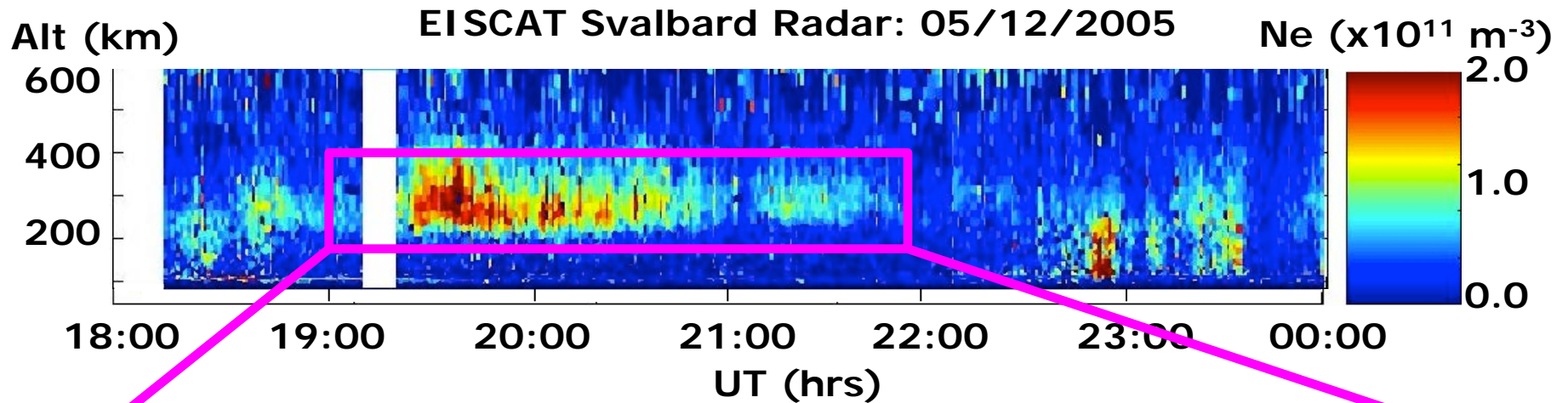


Structuring Ratio



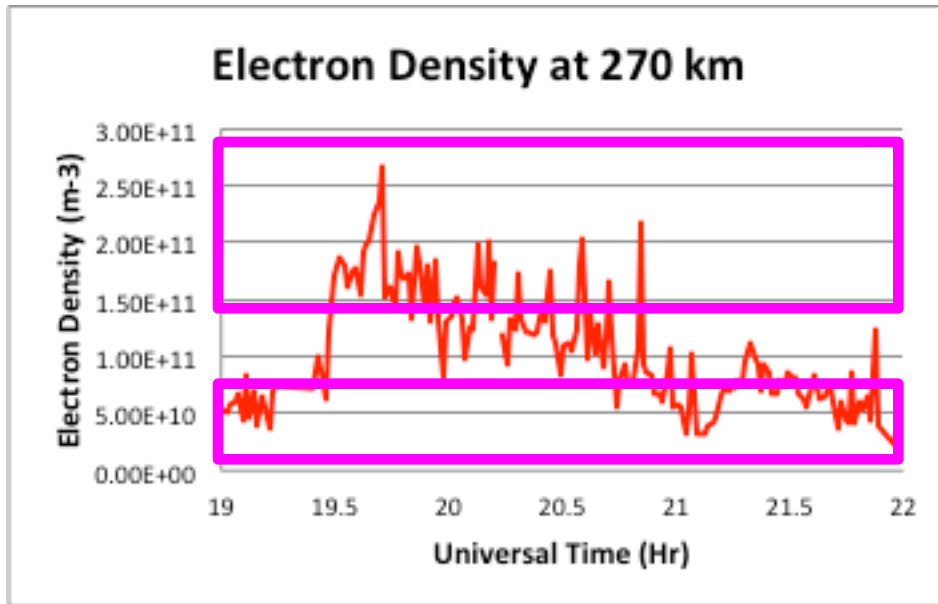
How highly structured is this plasma?

Structuring Ratio



Peak Altitude: 270 km

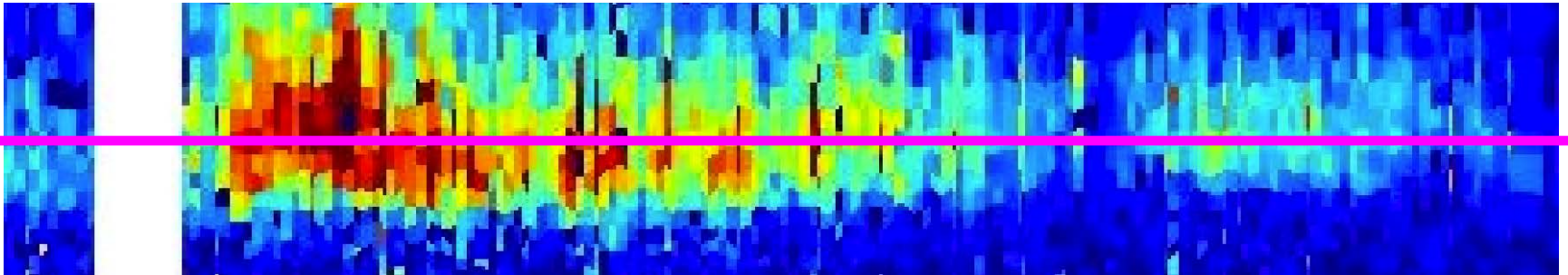
Structuring Ratio



Larger Ne: $1.8 \times 10^{11} m^{-3}$

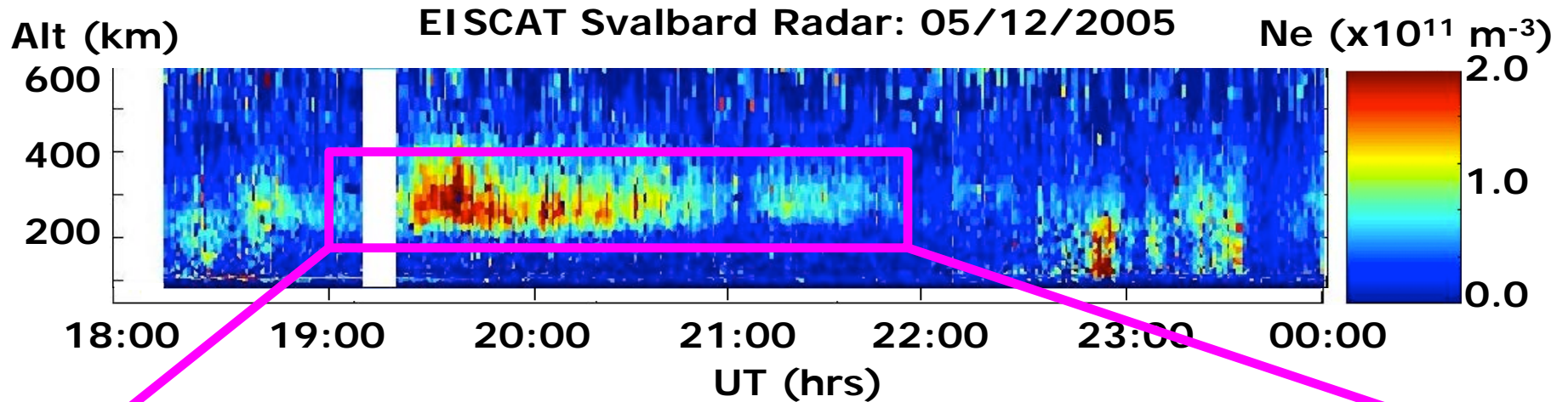
Smaller Ne: $0.5 \times 10^{11} m^{-3}$

Structuring Ratio: 3.6



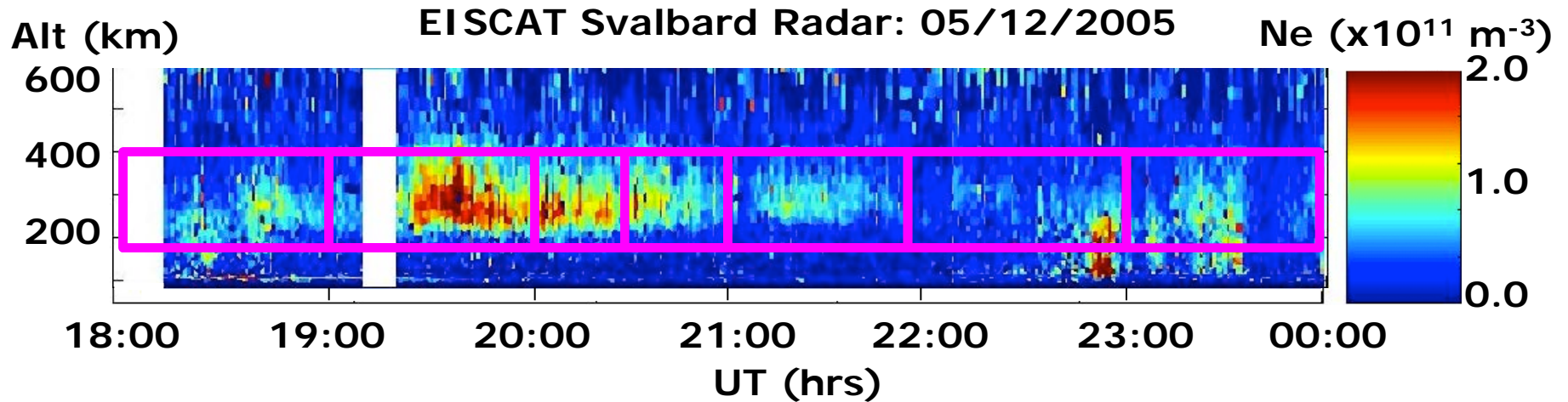
Peak Altitude: 270 km

Structuring Ratio



Peak Altitude: 270 km

Structuring Ratio



Time Intervals:

3 hours

EISCAT Svalbard Radar

Six hours of quasi-continuous observations in noon (06-12 UT), dusk (12-18 UT), midnight (18-00 UT) or dawn (00-06 UT) sector

Archive data: 1997-2013



Number of observations

Noon: 290

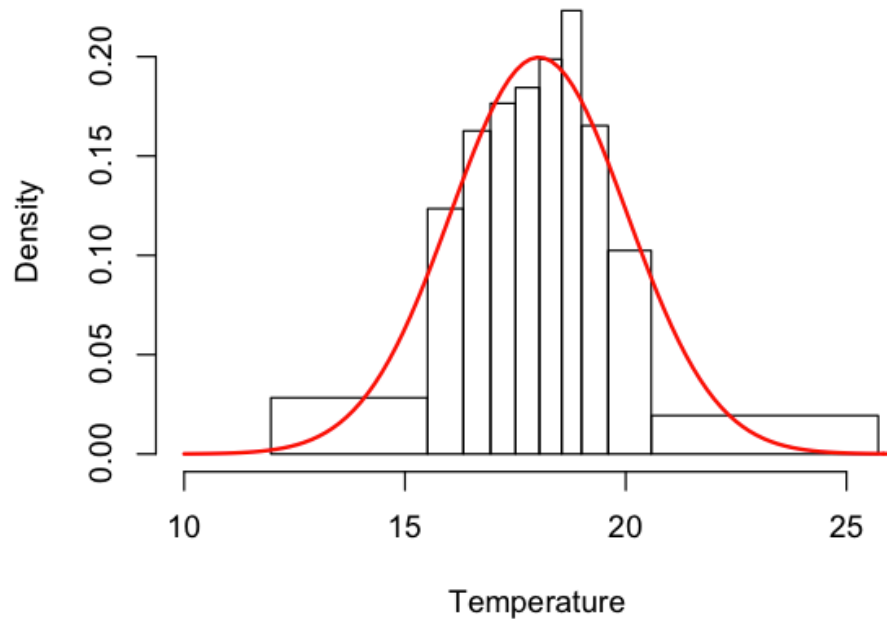
Dusk: 293

Midnight: 301

Dawn: 308

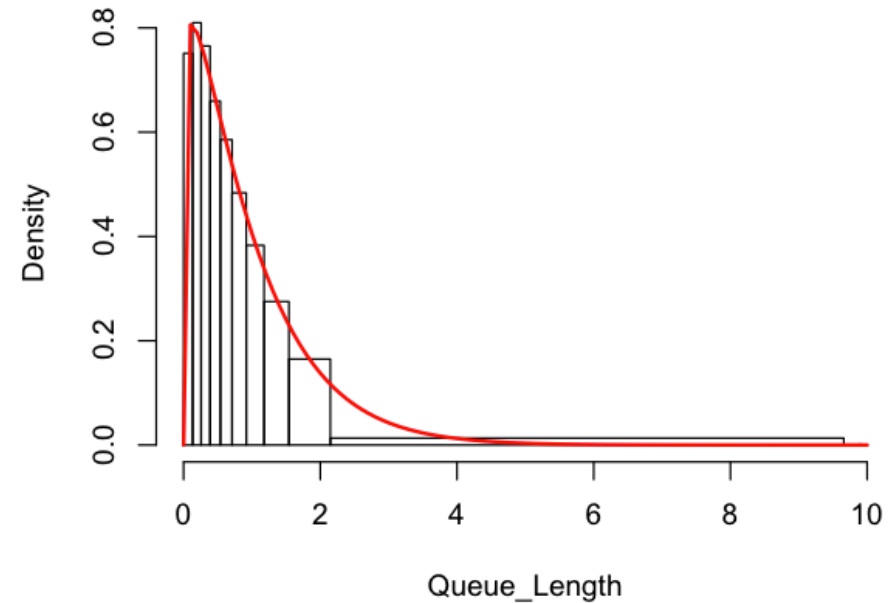
Distributional Form

Temperature of Conference Centre



Normal distribution

Number of People in Queue

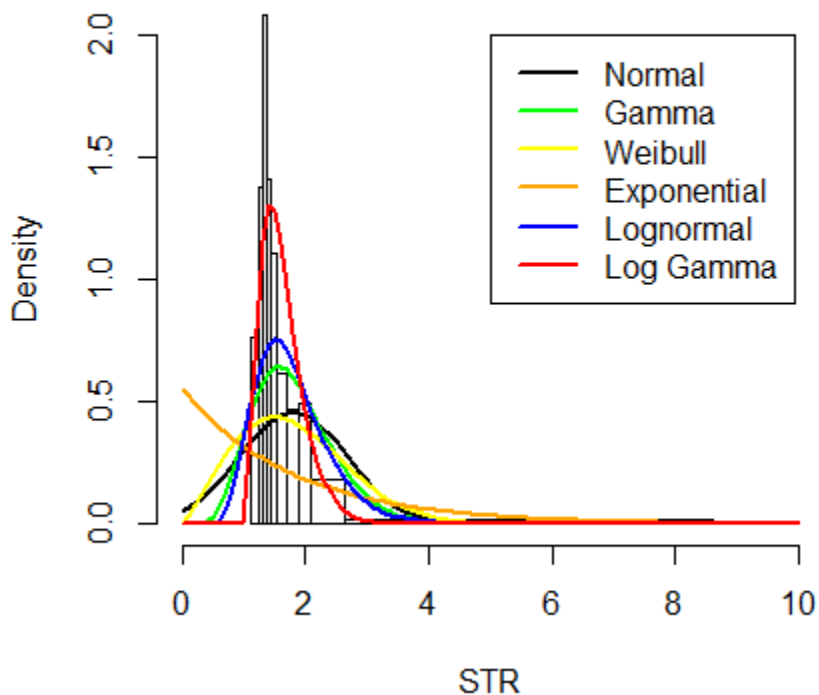


Weibull distribution

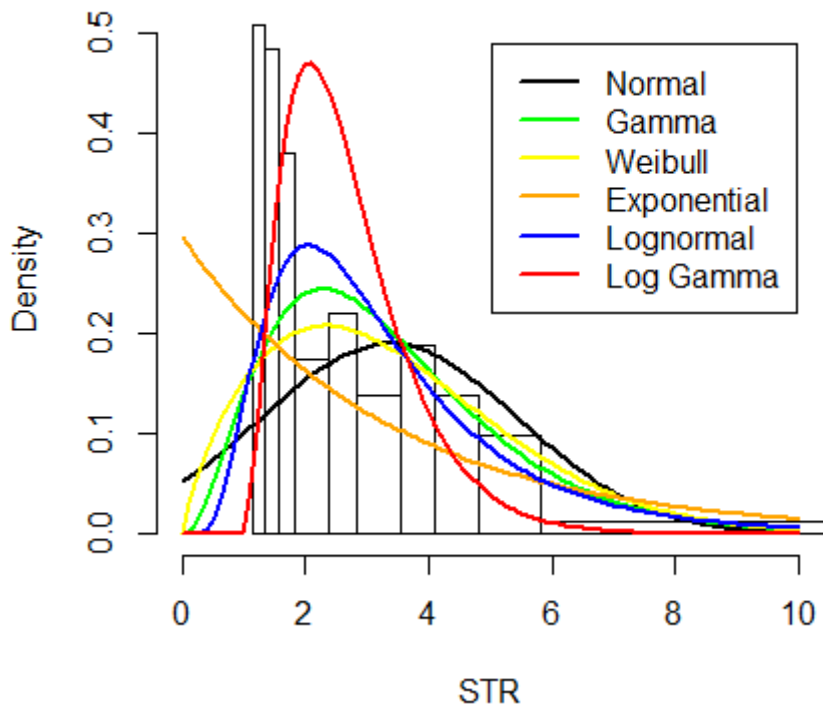
Tried: Normal, Gamma, Exponential, Weibull, Lognormal & Log Gamma distributions

Distributional Form

F-region, dusk, around peak, 2 hours



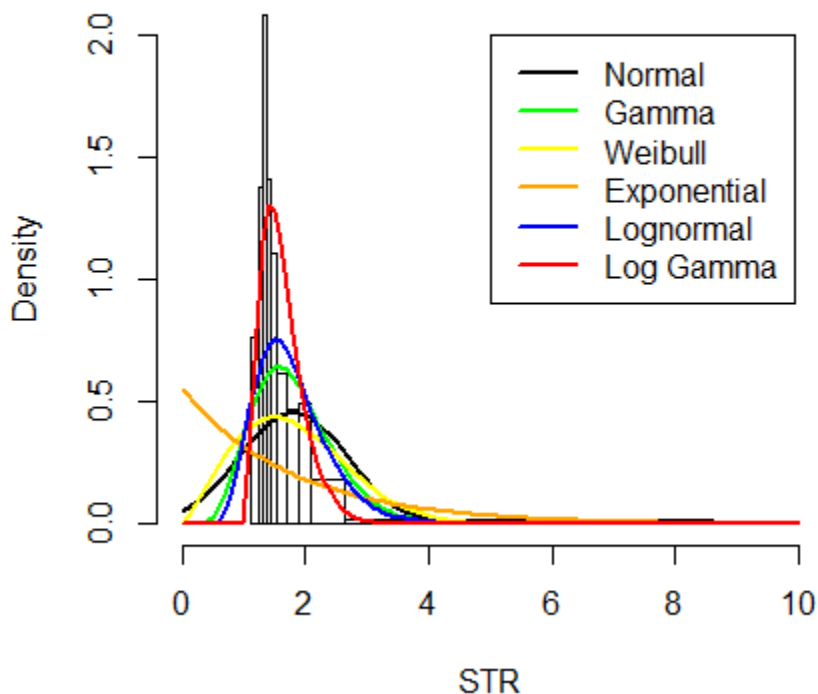
F-region, midnight, around peak, 6 hours



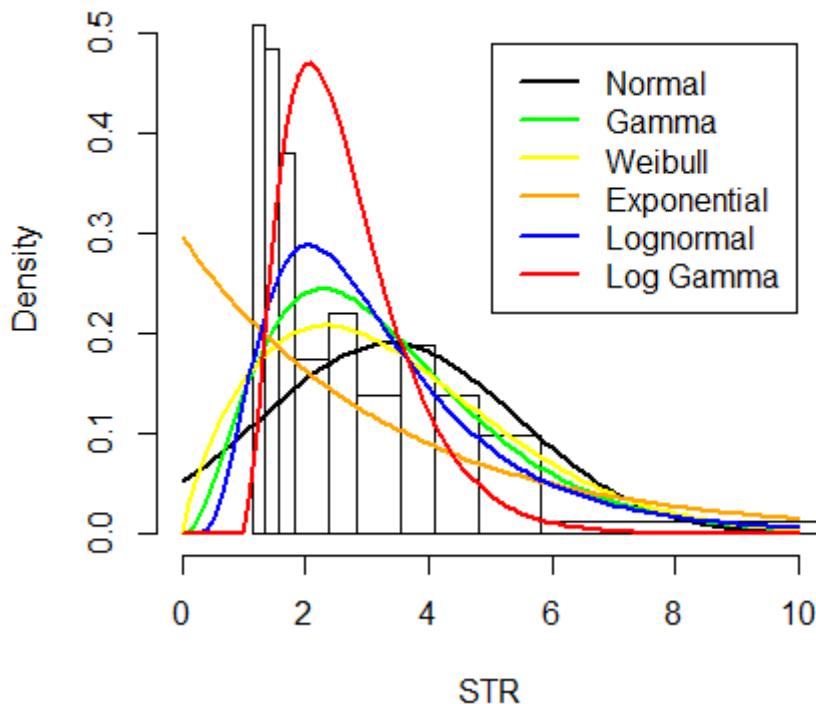
Log Gamma fits best, but does not fit everywhere

Distributional Form

F-region, dusk, around peak, 2 hours



F-region, midnight, around peak, 6 hours



Expect Chi-sq ~ 1 per degree of freedom

Inspect all plots: If Chi-sq more than x3 expected value, or less than x1/3 expected value, reject

Link Function

	Link Function	Form of equation
Identity	$X\beta = \mu$	$\log(STR) = a + b \cdot x$
Inverse	$X\beta = \mu^{-1}$	$\log(STR) = \frac{1}{a + b \cdot x}$
Log	$X\beta = \ln(\mu)$	$\log(STR) = a \cdot x^b$

Find which produces, on average, the most statistically significant results

$$\begin{aligned} \textit{Significance_Index} = & (4 \cdot \textit{Sig_0.01\%}) + (3 \cdot \textit{Sig_0.1\%}) \\ & + (2 \cdot \textit{Sig_1\%}) + (1 \cdot \textit{Sig_5\%}) \end{aligned}$$

Generalised Linear Modelling

Dependant Variable: Amount of plasma structuring

Independent Variables: Solar activity (F10.7 Solar Flux & SSN)

Solar wind (IMF: Bx, By & Bz)

Geomagnetic activity (Kp, AE)

Plasma convection (Polar cap index)

Season (days from midwinter)

Modelling Method

Test statistical significance of all parameters

Number of Models

$$3 \times 7 \times 4 \times 2 = 168$$

Relative Importance of Geophysical Parameters

Svalbard

Parameter	Importance
Season	5.0
F10.7	3.5
Dst (average)	3.4
PCI (average)	2.9
SSN	2.9
Kp (average)	2.8
IMF Bz (average)	1.7
IMF Bz (stdev)	1.6
IMF By (stdev)	1.4
PCI (stdev)	1.3
AE (stdev)	1.2
Dst (stdev)	1.2
AE (average)	1.0
IMF Bx (stdev)	0.7
IMF Bx (average)	0.4
Kp (stdev)	0.1
IMF By (average)	0.0

Importance

Significance better than ...

... 5% Importance 1

... 1% Importance 2

... 0.1% Importance 3

... 0.01% Importance 4

... 0.001% Importance 5

Season is the most important parameter

Seasonal Variation

Parameter Estimate varies with altitude range and time range ...

Altitude	Time						
	6-hr	5-hr	4-hr	3-hr	2-hr	1-hr	30-min
Peak	1.11±0.05	1.13±0.06	1.18±0.06	1.19±0.06	1.27±0.07	1.3±0.1	1.5±0.2
Around Peak	1.17±0.05	1.20±0.06	1.27±0.06	1.29±0.06	1.41±0.08	1.5±0.1	1.6±0.2
Height Integrated	1.33±0.07	1.41±0.08	1.47±0.09	1.56±0.09	1.8±0.1	2.1±0.1	2.4±0.2

F region, dusk

... but this can be largely explained by the variation in the amount of structuring

Altitude	Time						
	6-hr	5-hr	4-hr	3-hr	2-hr	1-hr	30-min
Peak	2.33±0.12	2.22±0.11	2.17±0.11	2.06±0.11	1.97±0.11	1.97±0.13	2.25±0.33
Around Peak	2.18±0.07	2.08±0.06	2.03±0.07	1.90±0.05	1.81±0.05	1.77±0.06	2.04±0.23
Height Integrated	1.90±0.05	1.83±0.05	1.79±0.05	1.68±0.03	1.60±0.03	1.48±0.03	1.41±0.02

F region, dusk

Seasonal Variation

Variation with sector and region shows significant differences

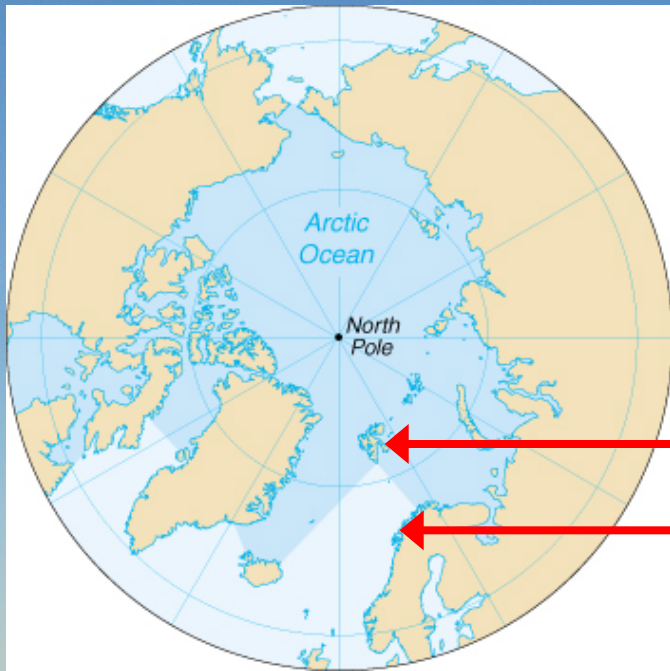
3 hour, around peak

Sector	F-region	E-region
Dusk	1.29±0.06	0.63±0.05
Midnight	1.01±0.05	0.36±0.03
Dawn	0.77±0.04	-0.70±0.05

5 hour, TEC

Sector	F-region	E-region
Dusk	1.40±0.08	0.51±0.04
Noon	1.08±0.09	0.35±0.04

Different dependence upon season at different MLT and regions



Svalbard
Tromsø



Relative Importance of Geophysical Parameters

Svalbard

Parameter	Importance
Season	5.0
F10.7	3.5
Dst (average)	3.4
PCI (average)	2.9
SSN	2.9
Kp (average)	2.8
IMF Bz (average)	1.7
IMF Bz (stdev)	1.6
IMF By (stdev)	1.4
PCI (stdev)	1.3
AE (stdev)	1.2
Dst (stdev)	1.2
AE (average)	1.0
IMF Bx (stdev)	0.7
IMF Bx (average)	0.4
Kp (stdev)	0.1
IMF By (average)	0.0

Tromsø

Parameter	Importance
Season	5.0
Dst (average)	2.6
PCI (average)	1.8
Kp (average)	1.3
F10.7	1.2
IMF By (stdev)	1.1
IMF Bz (average)	1.1
PCI (stdev)	0.9
AE (stdev)	0.8
IMF Bz (stdev)	0.8
Dst (stdev)	0.7
IMF Bx (stdev)	0.7
SSN	0.7
AE (average)	0.5
IMF Bx (average)	0.5
Kp (stdev)	0.4
IMF By (average)	0.3

Season still the most important parameter

Relative importance of other parameters changes

In the F-region 24 % of parameter estimates are significantly different

In the E-region 36 % of parameter estimates are significantly different

Relative Importance of Geophysical Parameters

Season: Parameter Estimate

F-region: Significantly different in 1 out of 31 cases

Predict more structuring in winter than summer

E-region: Significantly different in 9 out of 33 cases

Predict more structuring in winter than summer

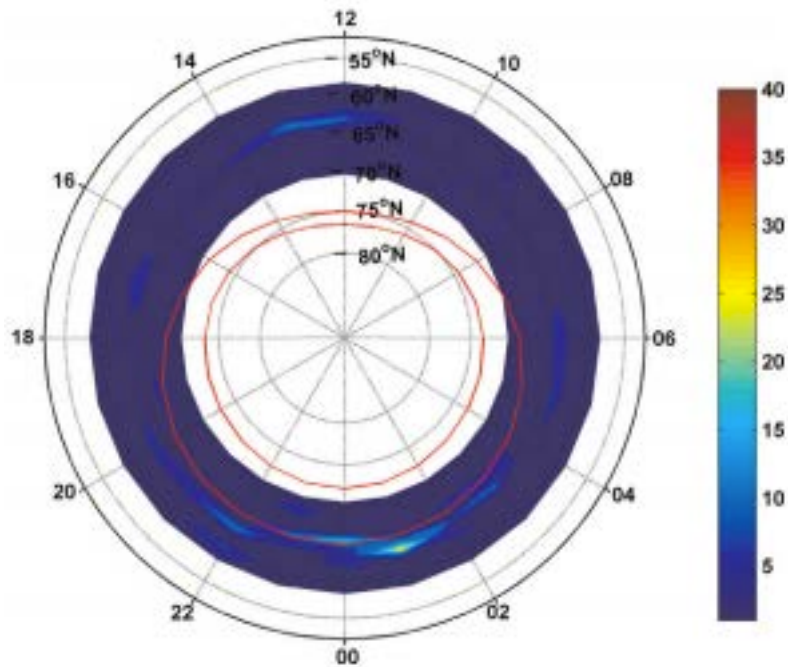
Where differences occur in the noon sector predict more variation in E-region than F-region

Where differences occur in the dawn sector predict less variation in E-region than F-region

Phase Scintillation

Data supplied by Nottingham Geospatial Institute

Occurrence of phase scintillation > 0.3



Binomial Model

1 year of data: 02.2012-01.2013

Restrict to EI > 70 degrees



Sreeja & Aquino, JASTP, 2014
Data from 2011-2013

Relative Importance of Geophysical Parameters

Svalbard

Tromsø

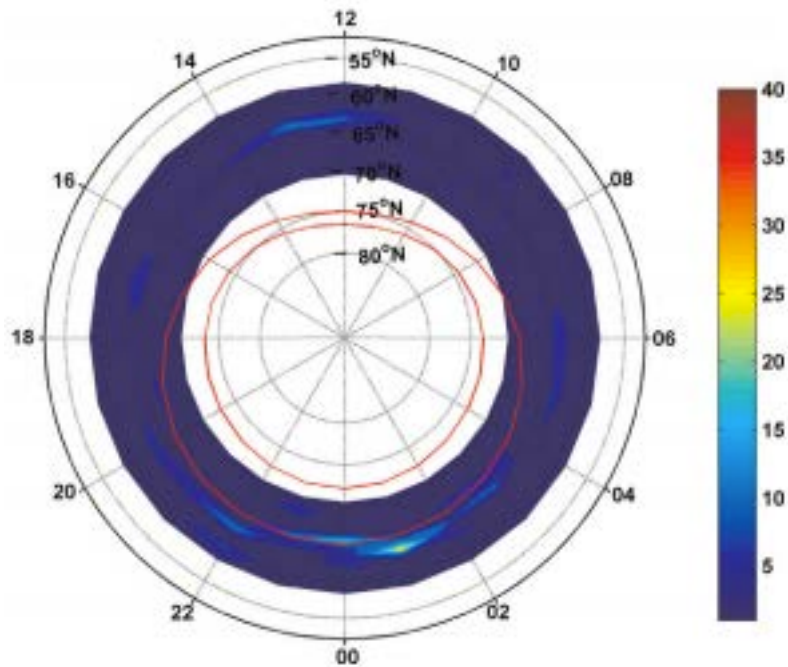
Bronnoysund

Parameter	Importance	Parameter	Importance	Parameter	Importance
Season	5.0	Season	5.0	F107	4.0
F10.7	3.5	Dst (average)	2.6	Solar Wind Density (average)	4.0
Dst (average)	3.4	PCI (average)	1.8	Season	3.8
PCI (average)	2.9	Kp (average)	1.3	SSN	3.0
SSN	2.9	F10.7	1.2	Kp (average)	2.8
Kp (average)	2.8	IMF By (stdev)	1.1	AE (average)	2.8
IMF Bz (average)	1.7	IMF Bz (average)	1.1	DST (average)	1.0
IMF Bz (stdev)	1.6	PCI (stdev)	0.9	PCI (average)	0.5
IMF By (stdev)	1.4	AE (stdev)	0.8	Solar Wind Velocity (Average)	0.0
PCI (stdev)	1.3	IMF Bz (stdev)	0.8	IMF Bz (average)	0.0
AE (stdev)	1.2	Dst (stdev)	0.7	IMF By (average)	0.0
Dst (stdev)	1.2	IMF Bx (stdev)	0.7	IMF Bx (average)]	0.0
AE (average)	1.0	SSN	0.7	IMF Abs(By) (average)	0.0
IMF Bx (stdev)	0.7	AE (average)	0.5		
IMF Bx (average)	0.4	IMF Bx (average)	0.5		
Kp (stdev)	0.1	Kp (stdev)	0.4		
IMF By (average)	0.0	IMF By (average)	0.3		

Phase Scintillation

Data supplied by Nottingham Geospatial Institute

Occurrence of phase scintillation > 0.3



Binomial Model

1 year of data: 02.2012-01.2013

Restrict to EI > 70 degrees

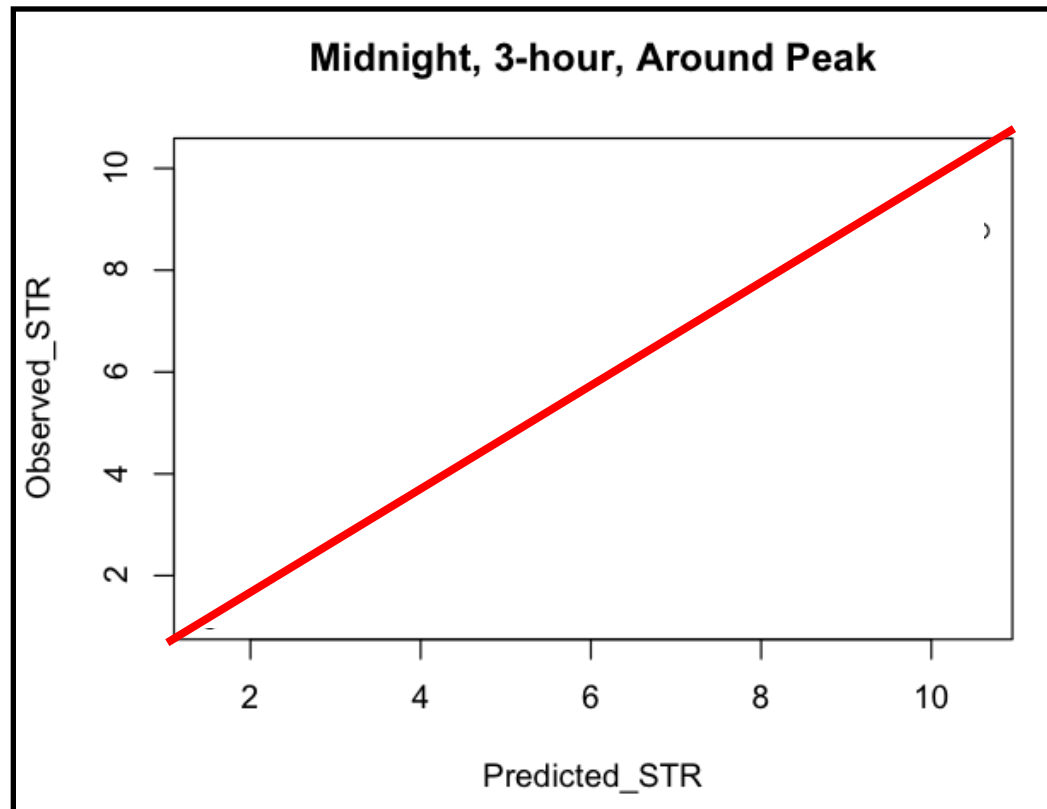
Sector	Parameter Estimate	Standard Error
Noon	0.035	0.003
Dusk	-0.028	0.014
Midnight	-0.032	0.011
Dawn	-0.063	0.010

Dusk & Midnight not significantly different; all other comparisons are significantly different

Sreeja & Aquino, JASTP, 2014
Data from 2011-2013

Model Results

$$\log(\text{STR}) = \frac{1}{1.28 \cdot \text{Season} - 0.09 \cdot \text{IMF_Bz_Stdev} + 2.36}$$



Seasonal Variation

EISCAT Svalbard Radar observations

Around solar maximum

IMF B_z predominantly negative

ESR in antisunward cross polar flow

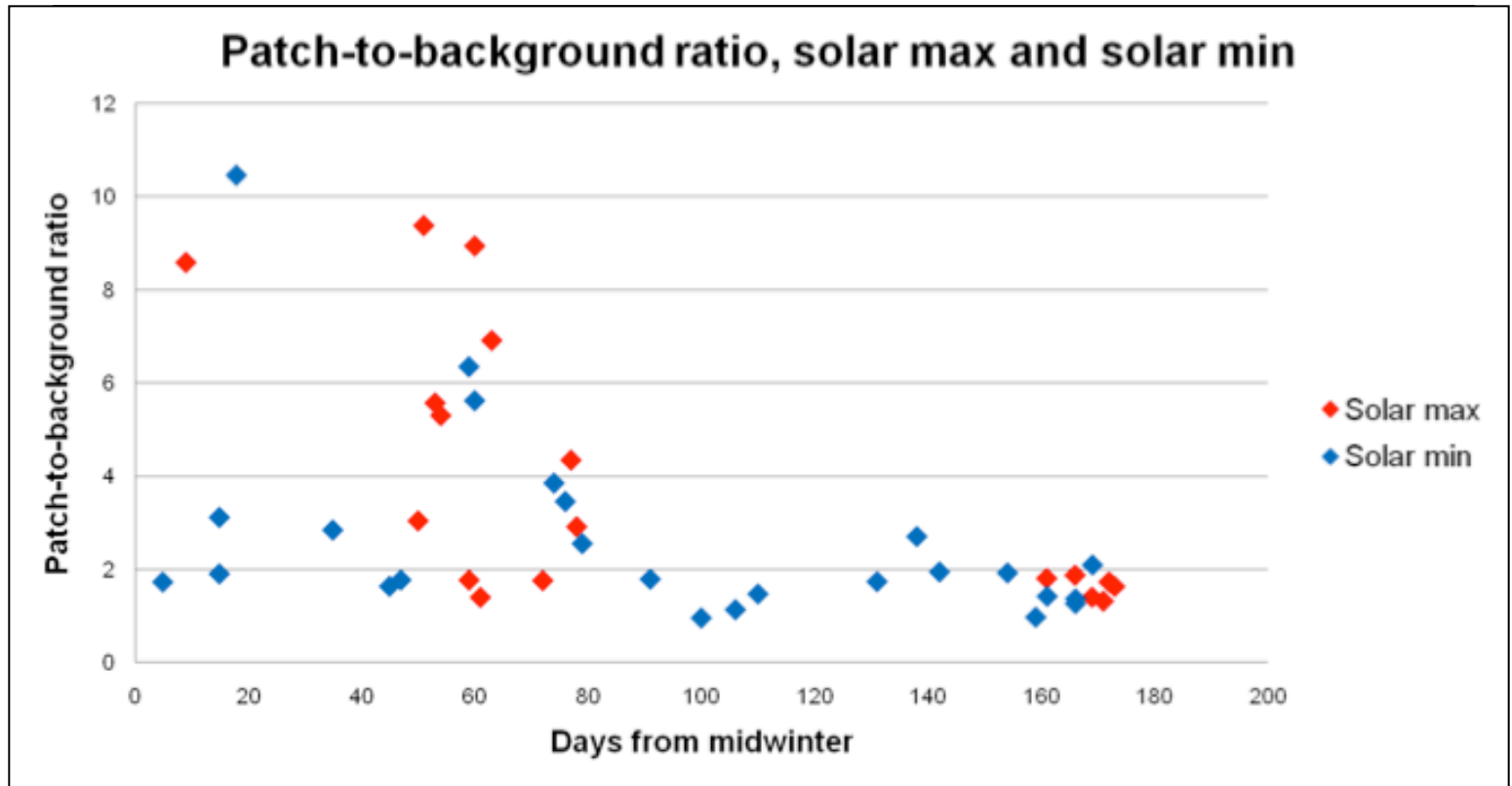
ESR poleward of Harang discontinuity

No evidence of precipitation

42 studies from 6 years of data



Seasonal Variation



Relatively little variation between solar maximum & solar minimum
Enhancements weaker in summer

Next Steps: Statistical Models

Next steps: Model the 'upper limit'

Next steps: Thermosphere (Amy Ronksley)

Comparison of thermospheric & ionospheric parameters using over a solar cycle of data

Look for statistically significant relationships between thermospheric and ionospheric parameters

Break statistical models into times of Sudden Stratospheric Warming (SSW) & non-Sudden Stratospheric Warming

Summary

Large scale structures

Season is the most significant parameter; the parameter estimate varies between sectors

Small scale structures

F10.7 is the most significant parameter; the parameter estimate varies between sectors

Thermosphere

Seems that at least one parameter is missing.

Next step: Include thermospheric parameters