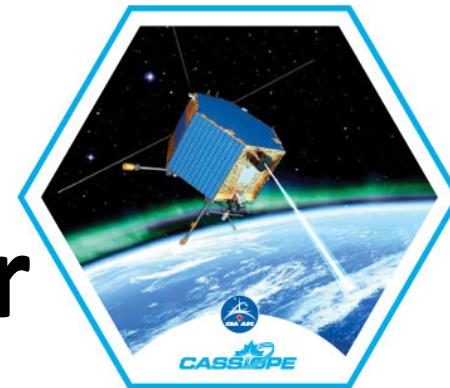




UNIVERSITY OF  
CALGARY



# e-POP GAP Occultation Receiver (GAP-O) onboard CASSIOPE

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# Outline

- CASSIOPE satellite & e-POP payload, including GAP
- GAP-O Occultations
- Ionosphere electron density profiles
- GAP-O receiver bias
- Ground/GAP-O observations of high latitude structures
- Visualization tools and data availability

- Designed to study solar wind-magnetosphere-ionosphere (SW-M-I) coupling processes in polar regions, including polar ion/neutral outflow, ionospheric structure, and radio propagation

- **CASSIOPE:** Cascade Smallsat and Ionospheric Polar Explorer

- **e-POP:** Enhanced Polar Outflow Probe

- **GAP:** GPS Attitude, Positioning, and Profiling Experiment

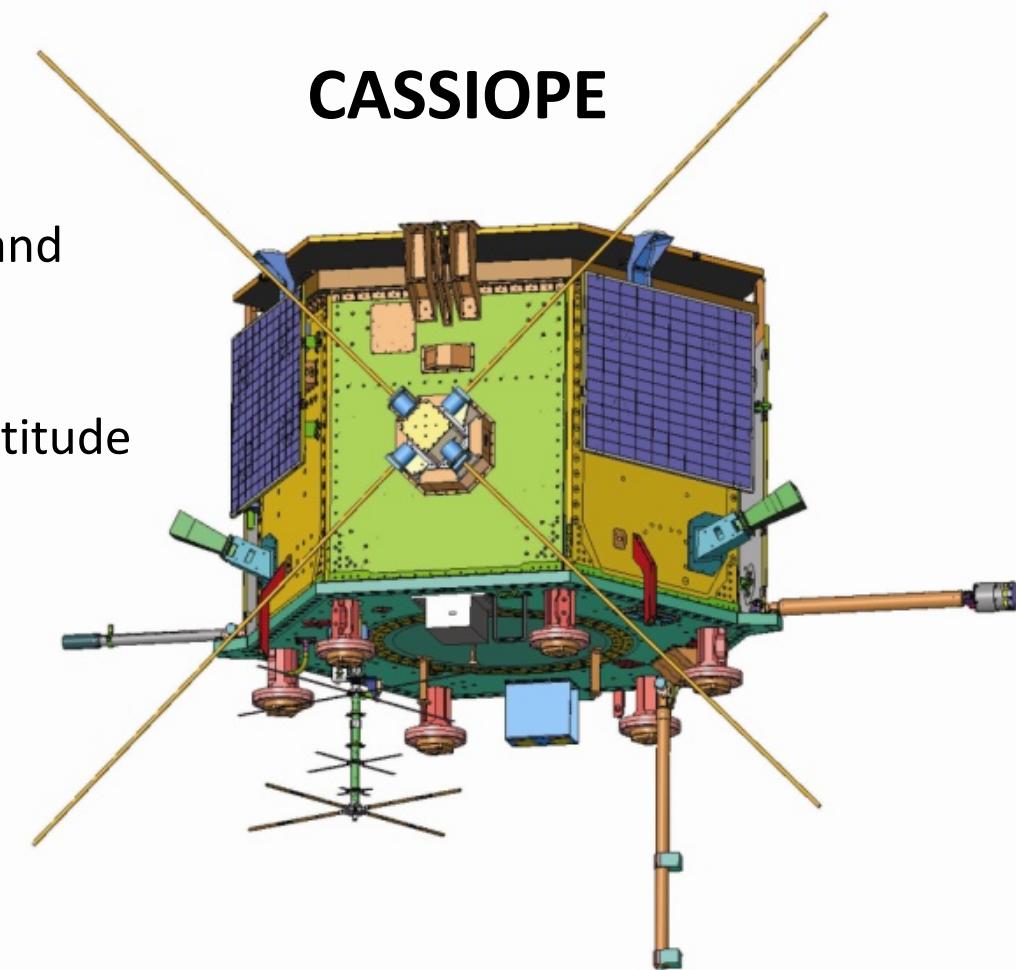
- **P.I.:** Richard Langley (University of New Brunswick Geodesy and Geomatics Engineering)

- **GAP-A:** 4 zenith facing antennas for navigation, spacecraft attitude and orbit determination

- **GAP-O:** 1 anti-ram facing antenna for occultation

- Dual frequency NovAtel OEM-4 receivers

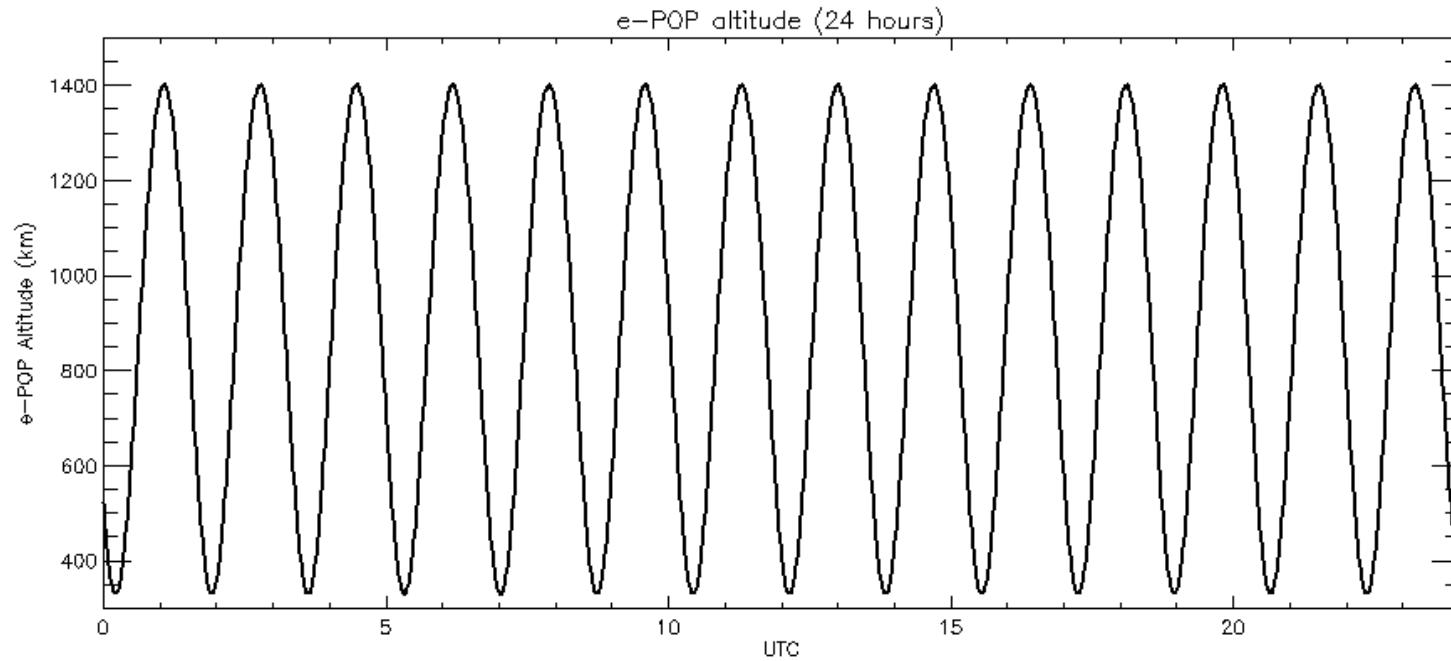
- L1 & L2 carrier phase, pseudorange, nominal C/N<sub>0</sub> at 20-100 Hz resolution



# CASSIOPE orbit

- Launched in late 2013
- Initial perigee: 325 km
- Initial apogee: 1500 km
- 81° Inclination
- ~1h 40min orbital period

e-POP Altitude(24 hours) – January 1, 2015



# GAP-O Publications

**D. Kim and R.B. Langley;**

The GPS attitude, positioning, and profiling experiment for the enhanced polar outflow probe platform on the Canadian CASSIOPE satellite

Geomatica

(2010), 64(2),  
233-243

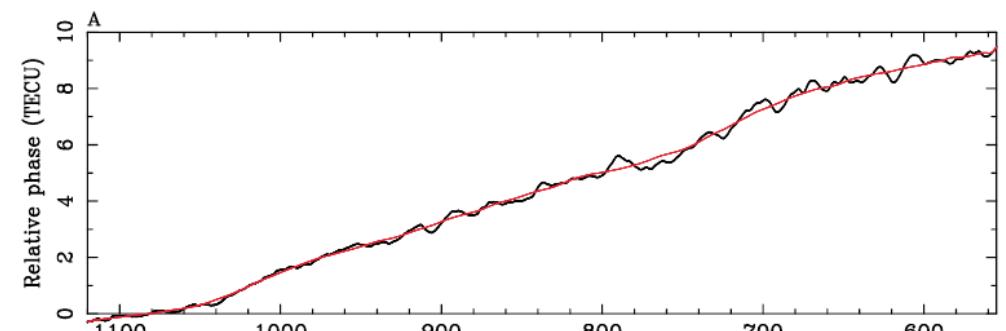
**E.B. Shume, A. Komjathy, R.B. Langley, O. Verkhoglyadova, M.D. Butala, and A.J. Manucci;**

Intermediate scale plasma irregularities in the polar ionosphere inferred from GPS radio occultation

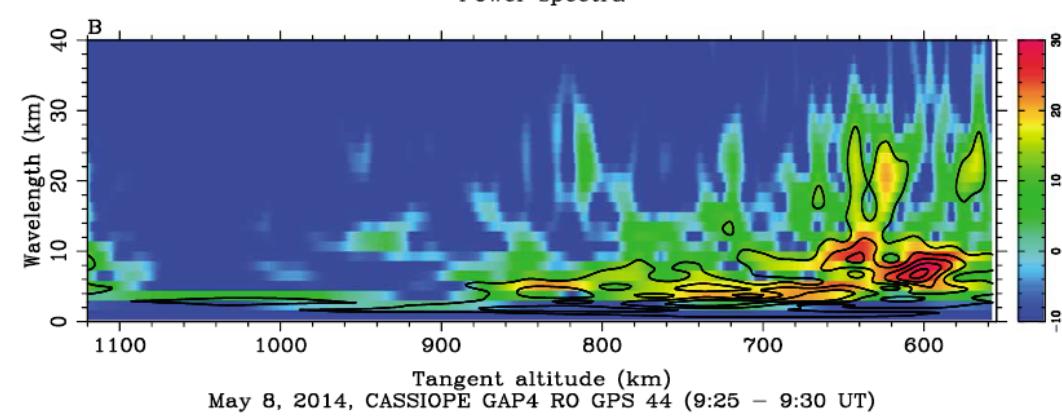
Geophys. Res.  
Lett. (2015) DOI:  
[10.1002/2014GL  
062558](https://doi.org/10.1002/2014GL062558)

Shume et al. [2015]

Dec 8, 2013, CASSIOPE GAP RO GPS 41 (21:42 – 21:54 UT)

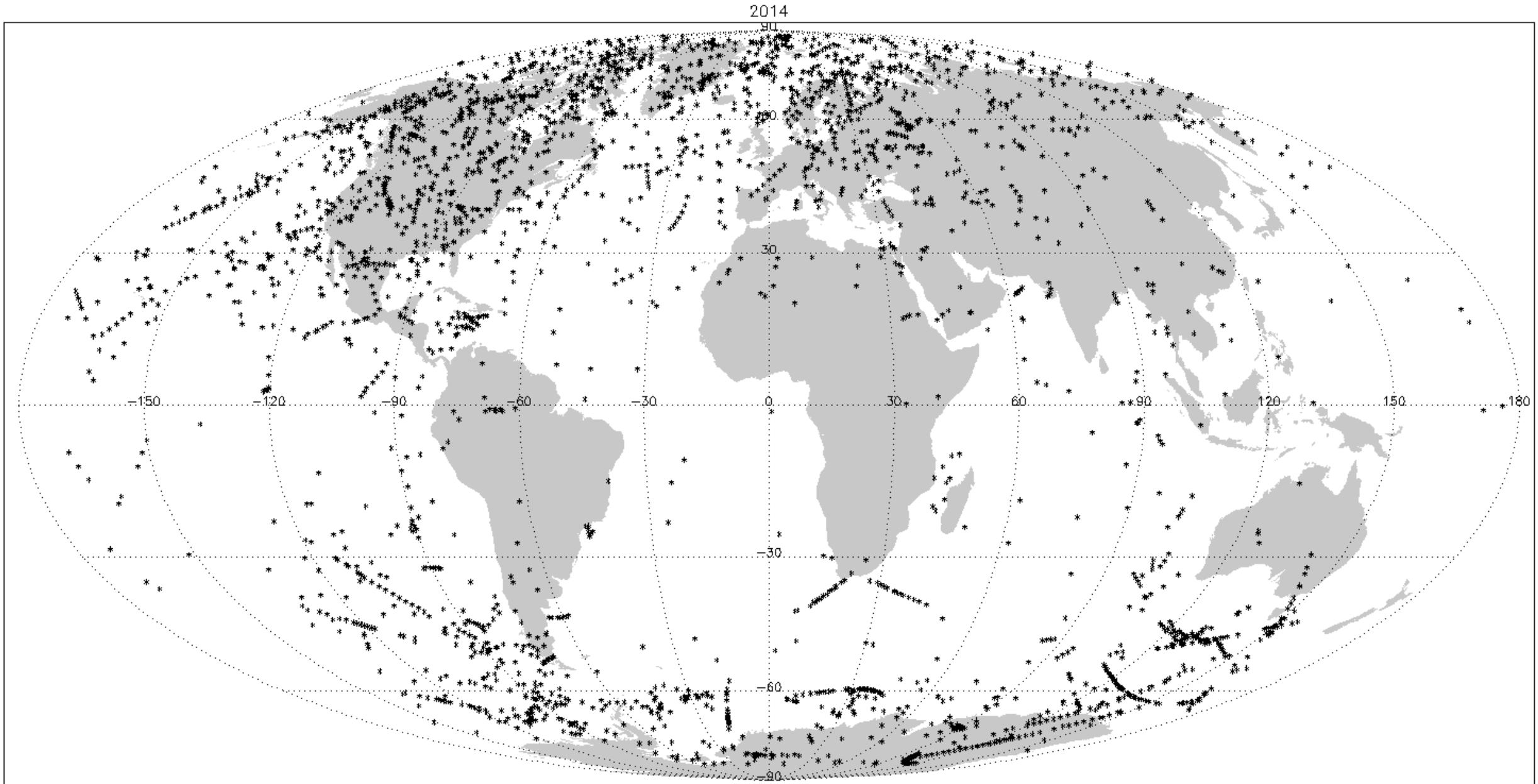


Power spectra



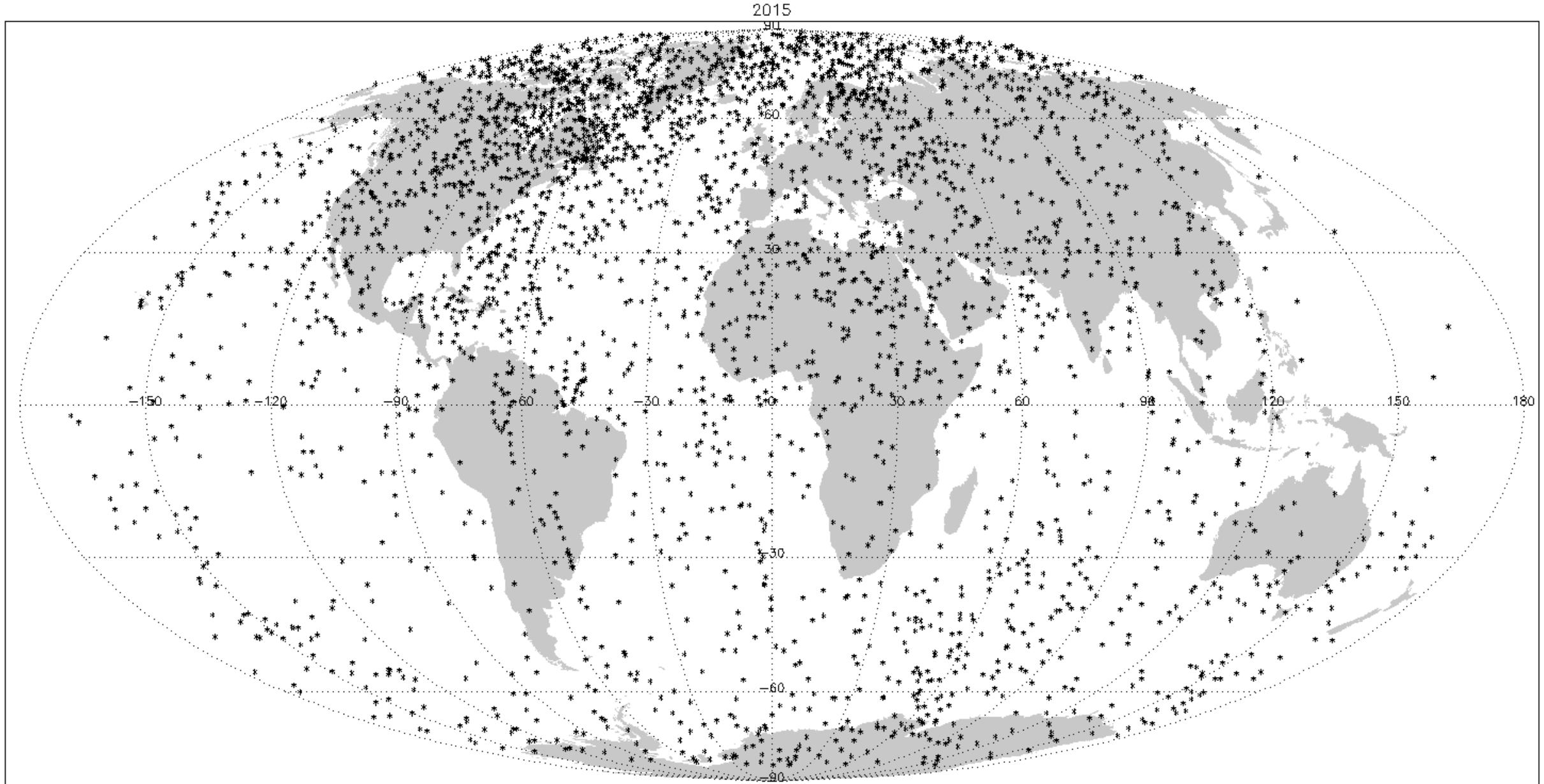
# GAP-O Occultations – Year 2014

## (Only events with useable data)



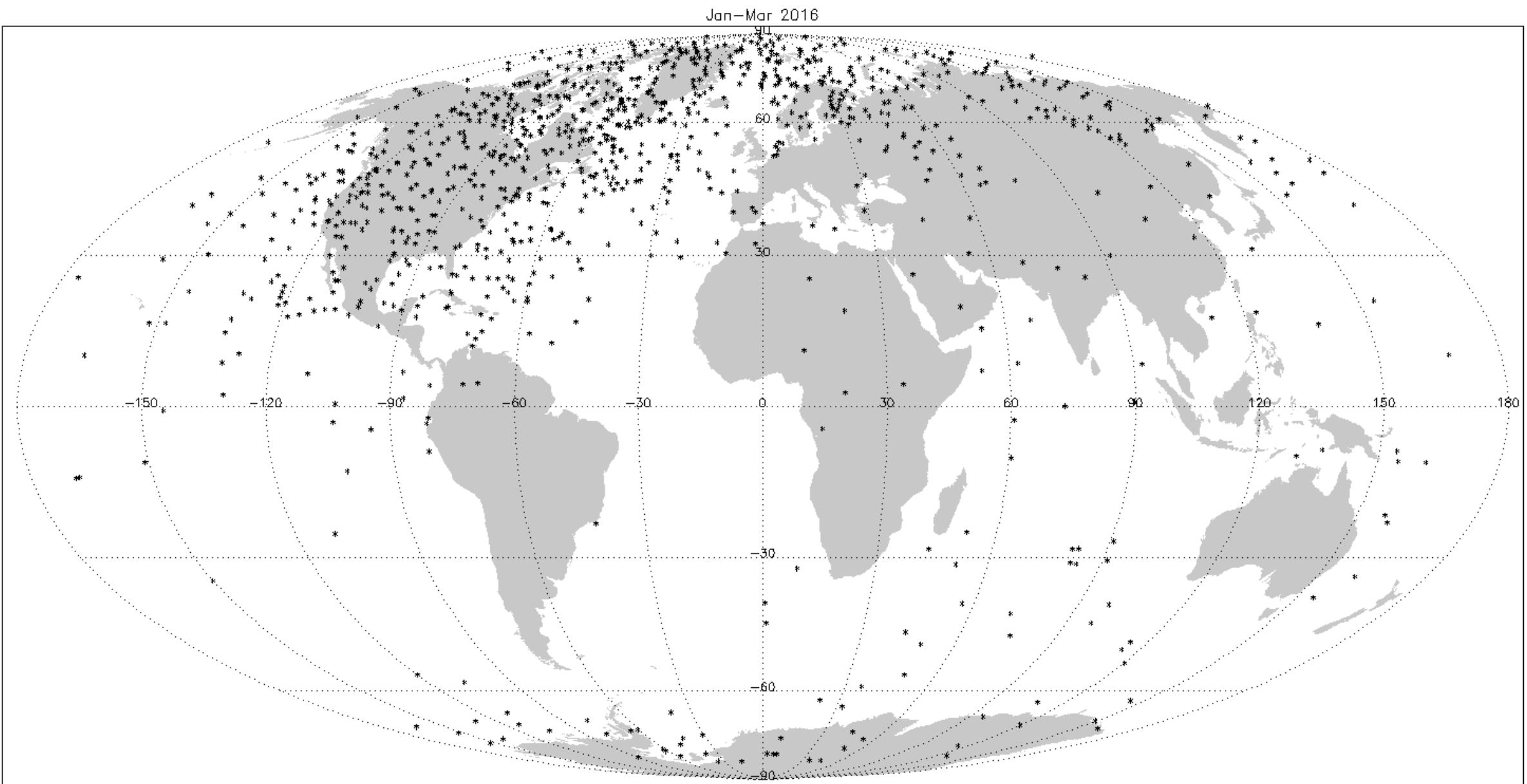
# GAP-O Occultations – Year 2015

## (Only events with useable data)

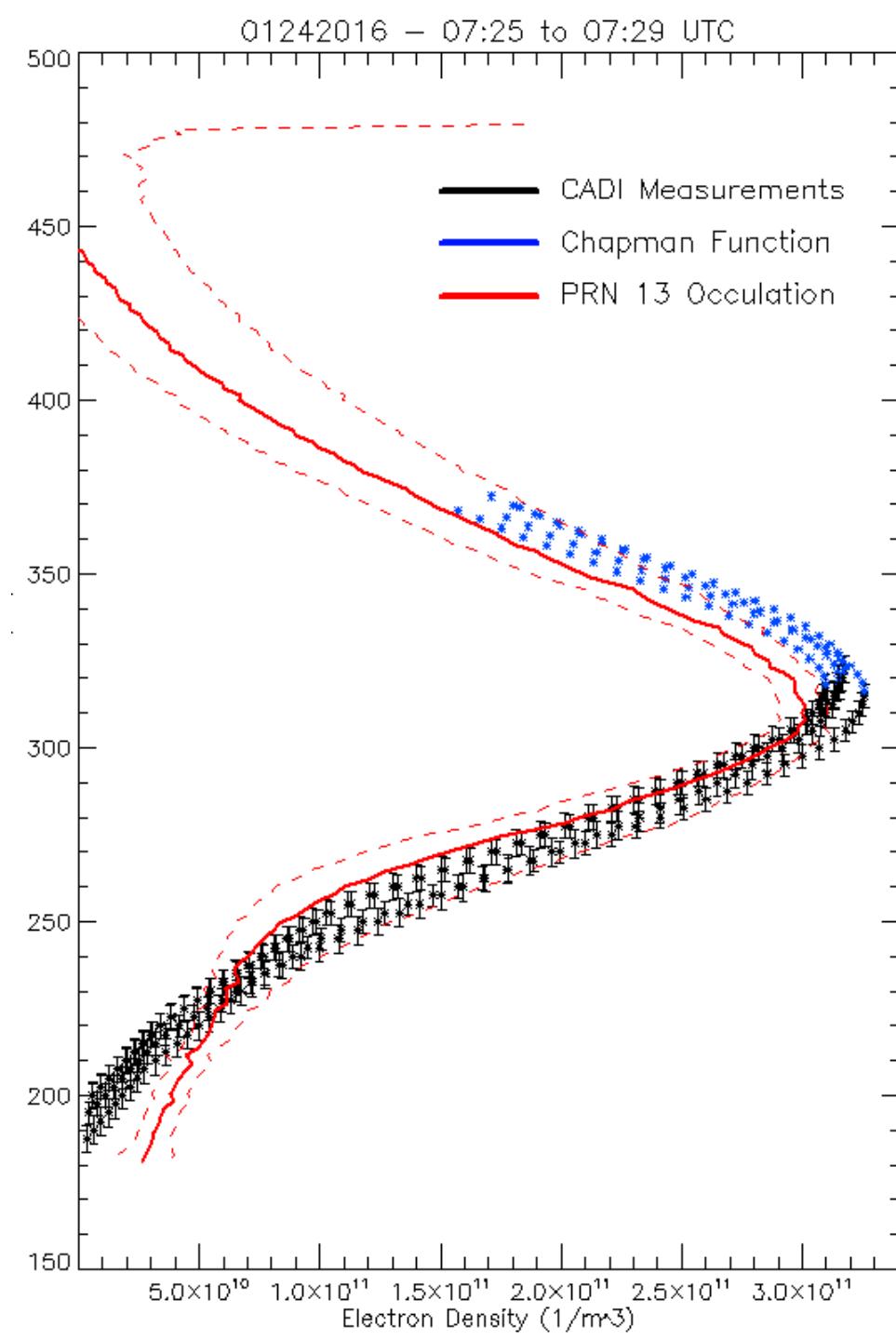
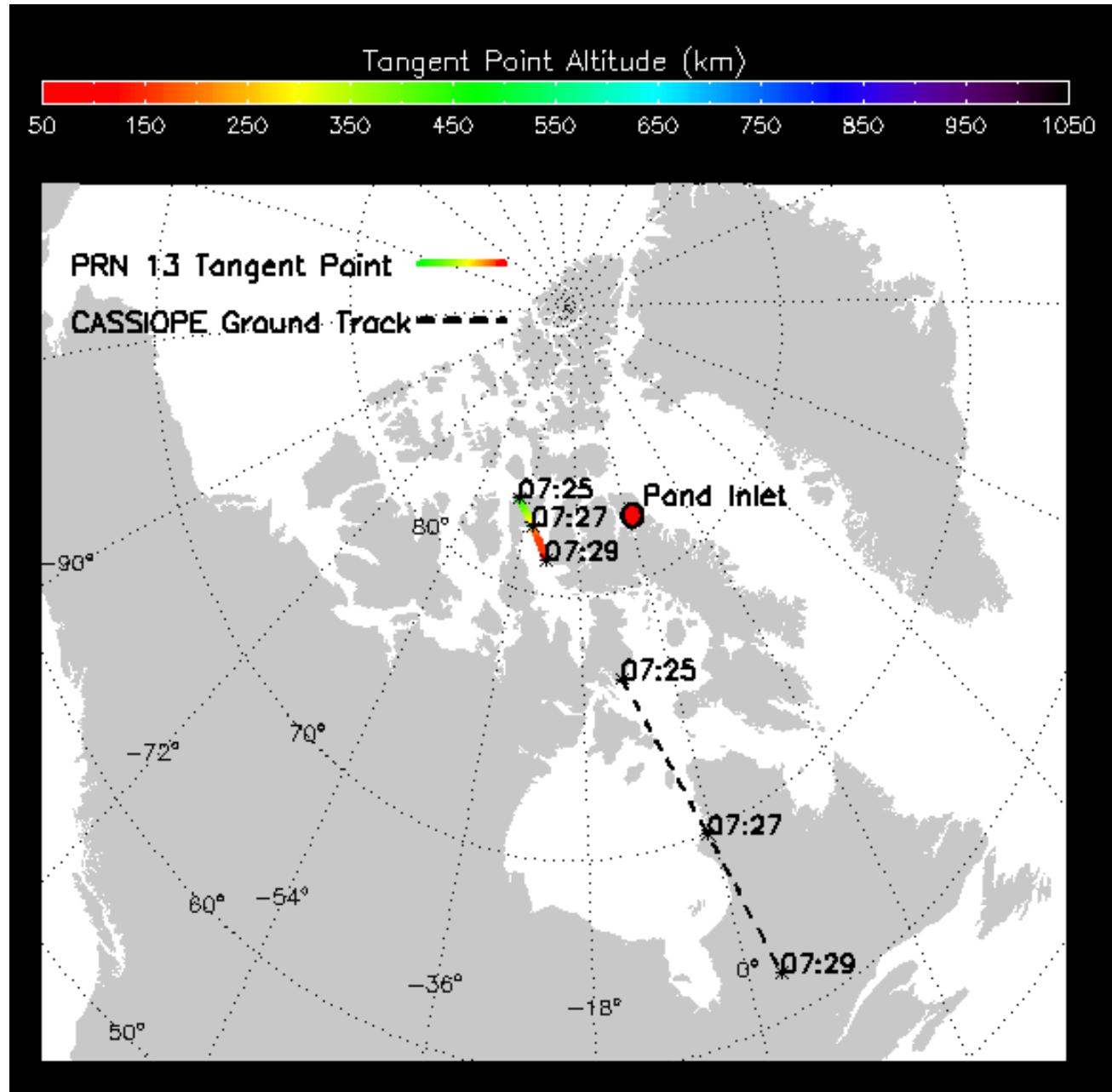


# GAP-O Occultations – January to March 2016

(Only events with useable data)



# Density profiles – Abel Inversion



# Differential code biases (DCBs)

- Frequency dependent differential delays:
  - GPS Satellite biases ( $DCB_s$ ) – e.g. <ftp://ftp.unibe.ch/aiub/CODE/>
  - GPS Receiver biases ( $DCB_r$ ) - unknown

TEC derived from carrier phase (L):  $TEC_{L,i} = TEC_{abs,i} + \frac{40.3(f_1^2 - f_2^2)}{f_1^2 f_2^2} \left( \frac{N_{1,i}c}{f_1} - \frac{N_{2,i}c}{f_2} \right) + \varepsilon_{L,i}$

TEC derived from pseudorange (P):  $TEC_{P,i} = TEC_{abs,i} - DCB_{s,i} - DCB_r + \varepsilon_{p,i}$

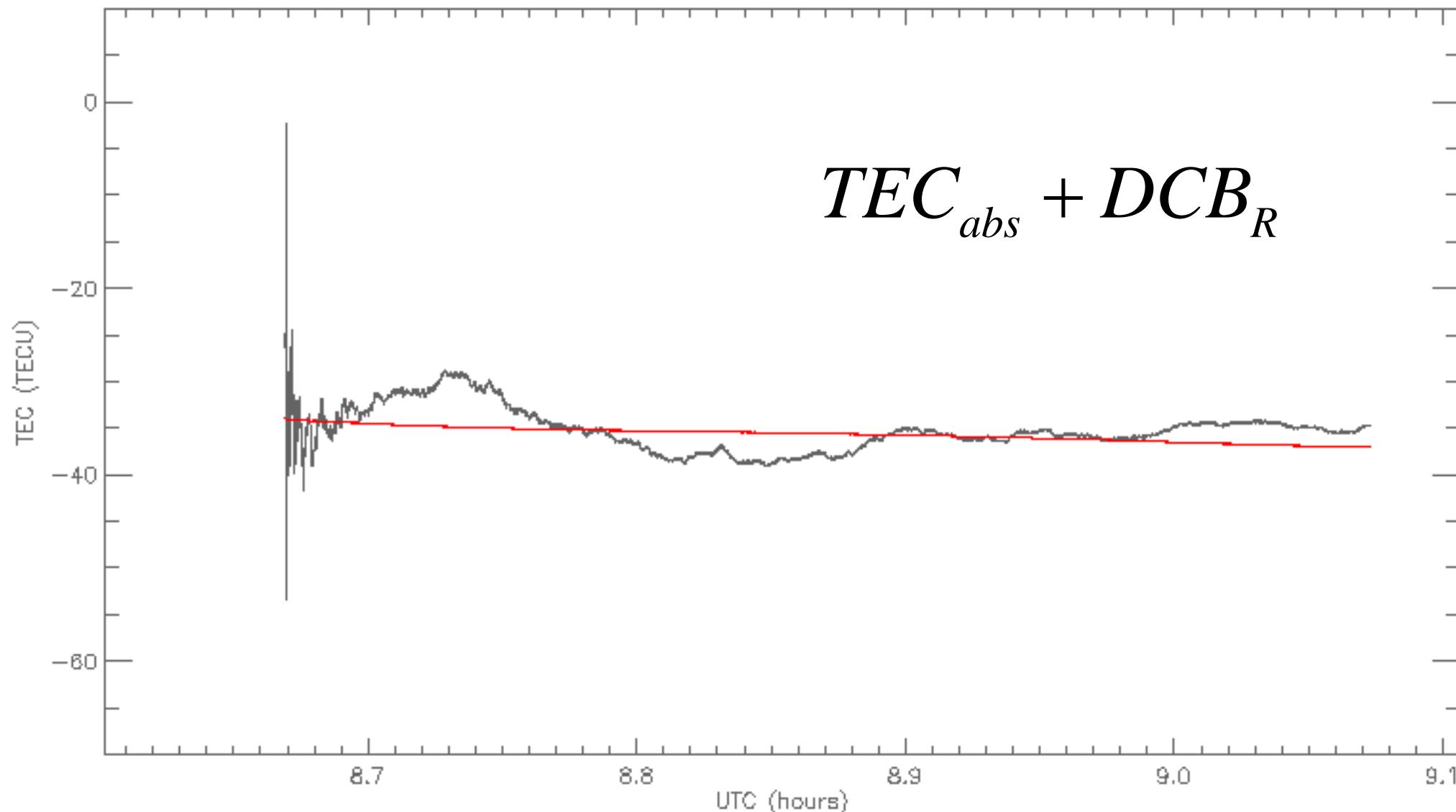
Unknown integer cycle ambiguity

Error terms

- DCB estimation for LEO receivers is tricky:
  - Fast moving satellites, which sample range of locations over a short time
  - Multipath and cycle slips reduce accuracy of DCB estimation methods
  - Typically only an hour or two of GAP-O measurements each day

# GAP-O Phase levelled TEC – Jan 6, 2014

PRN 01



Phase levelling algorithm: *Stephens et al. (2011), New leveling and bias estimation algorithms for processing COSMIC/FORMOSAT-3 data for slant total electron content, Radio Science, 46, doi:10.1029/2010RS004588.*

# Receiver bias estimation techniques

- Minimization of Standard Deviations
- Zero-TEC Method(s)
- Least Squares (LSQ) Method
- **Validation of receiver bias estimate using ground-based ionosonde & GPS TEC**

## Minimization of Standard Deviations

*Ma and Maruyama (2003), Determination of GPS receiver differential biases by neural network parameter estimation method, Radio Science, 40, doi:10.1029/2004RS003072.*

- Using topside TEC measurements (above e-POP):
  - Assume that the ionosphere is regionally homogenous
  - Ideally, vertical TEC (VTEC) calculated from different GPS satellite ray paths will be equal
  - Calculate the time-averaged mean and standard deviation of VTEC from multiple GPS satellites.
  - Find the receiver DCB that minimizes the standard deviation.

$$VTEC = (TEC_{biased} - DCB_r) \cdot M(E)$$

# Minimization of Standard Deviations – Mapping Functions for LEO

Thin layer model (TLM) [e.g.  
*Zhong et al., 2015*]:

$$M(E_{sat}) = \sqrt{1 - \left(\frac{R_{orbit}}{R_{shell}}\right)^2 \cos^2(E_{sat})} \quad R_{shell} = R_{orbit} + 300 \text{ km}$$

F&K [*Foelsche and Kirchengast, 2002*]:

$$M(E_{sat}) = \frac{\sin(E_{sat}) + \sqrt{\left(\frac{R_{shell}}{R_{orbit}}\right)^2 - \cos(E_{sat})}}{1 + \frac{R_{shell}}{R_{orbit}}}$$

Lear [*Lear, 1987*]:

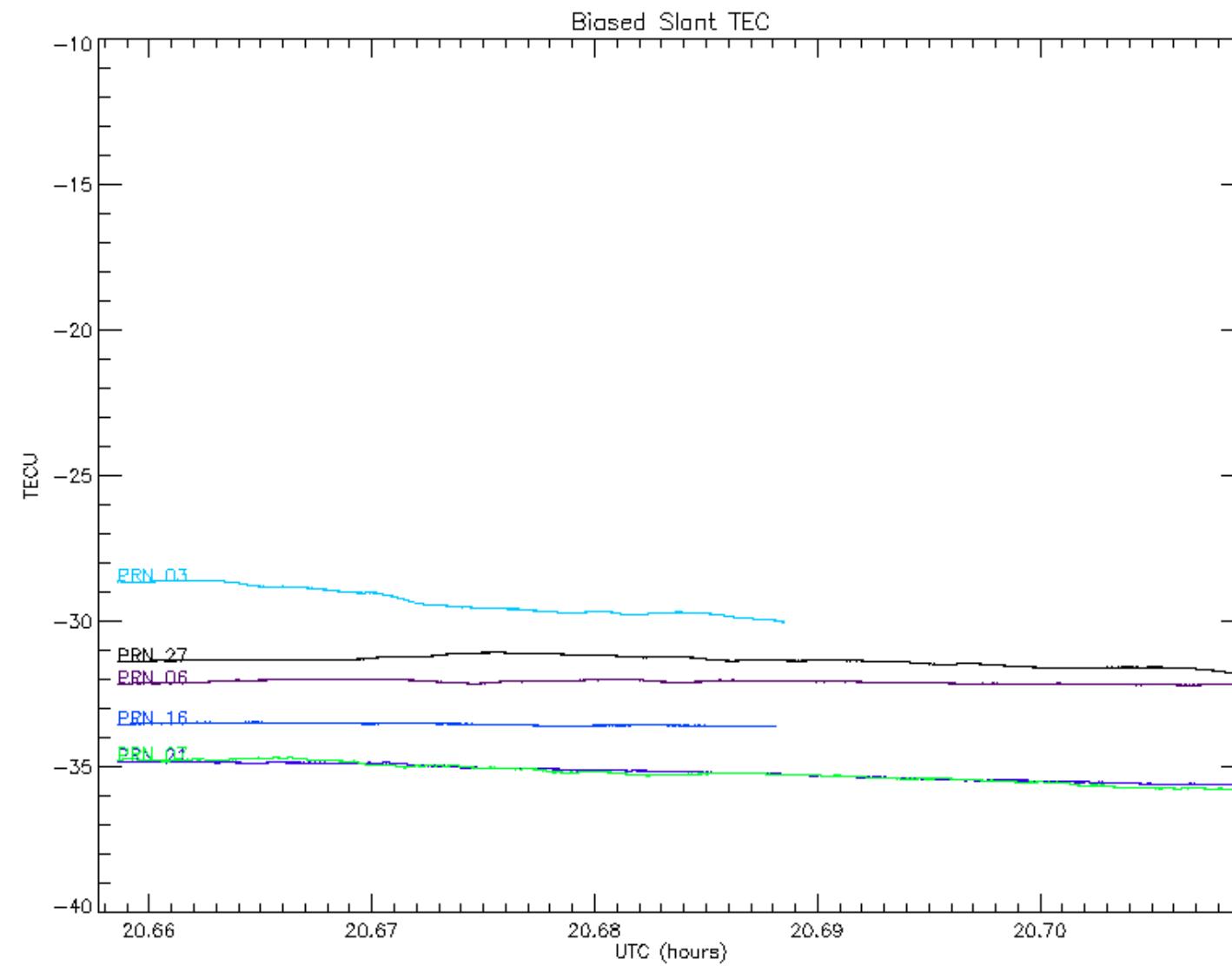
$$M(E_{sat}) = \frac{2.037}{\sin(E_{sat}) + \sqrt{1.076 - \cos^2(E_{sat})}}$$

A simple sine projection:

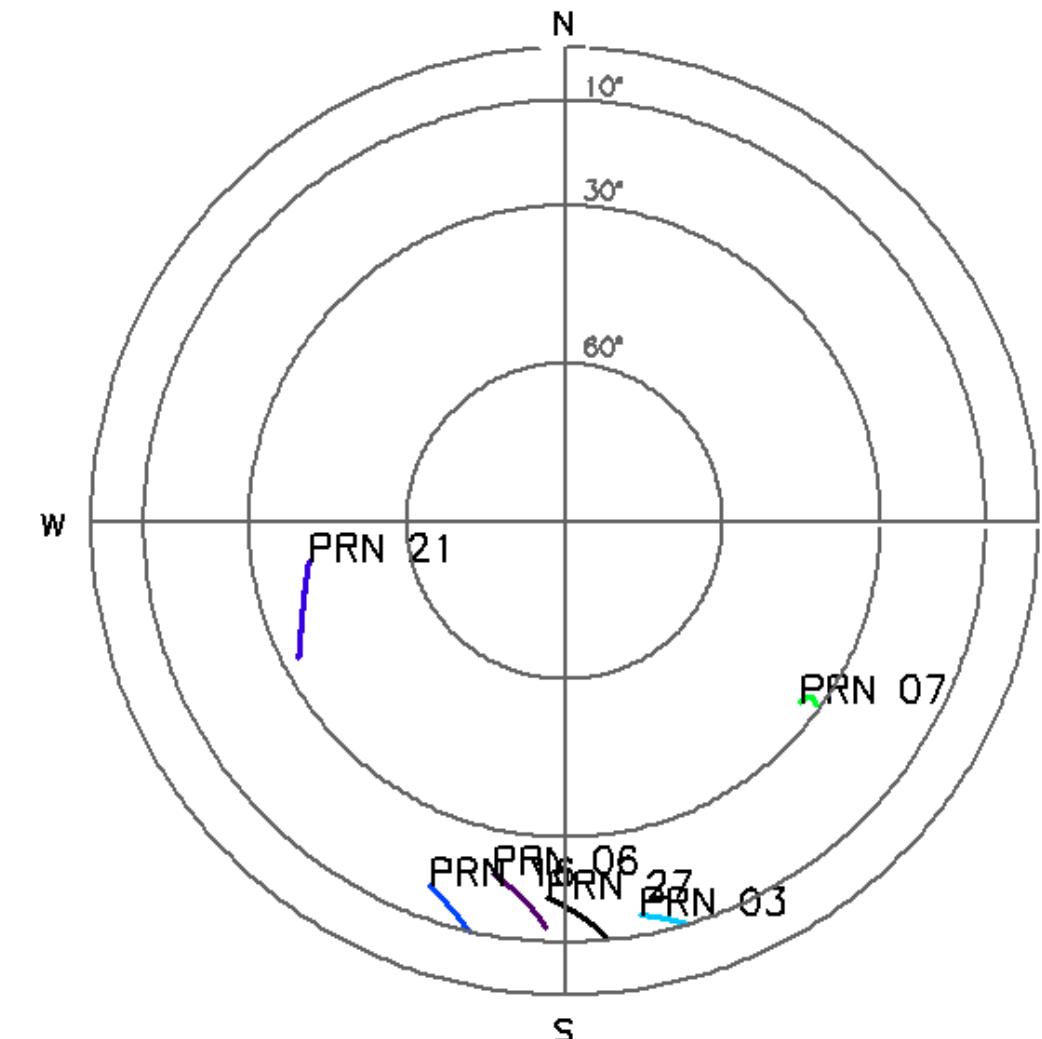
$$M(E_{sat}) = \sin(E_{sat})$$

# Minimization of Standard Deviations – Jan 2, 2014

Biased Slant TEC



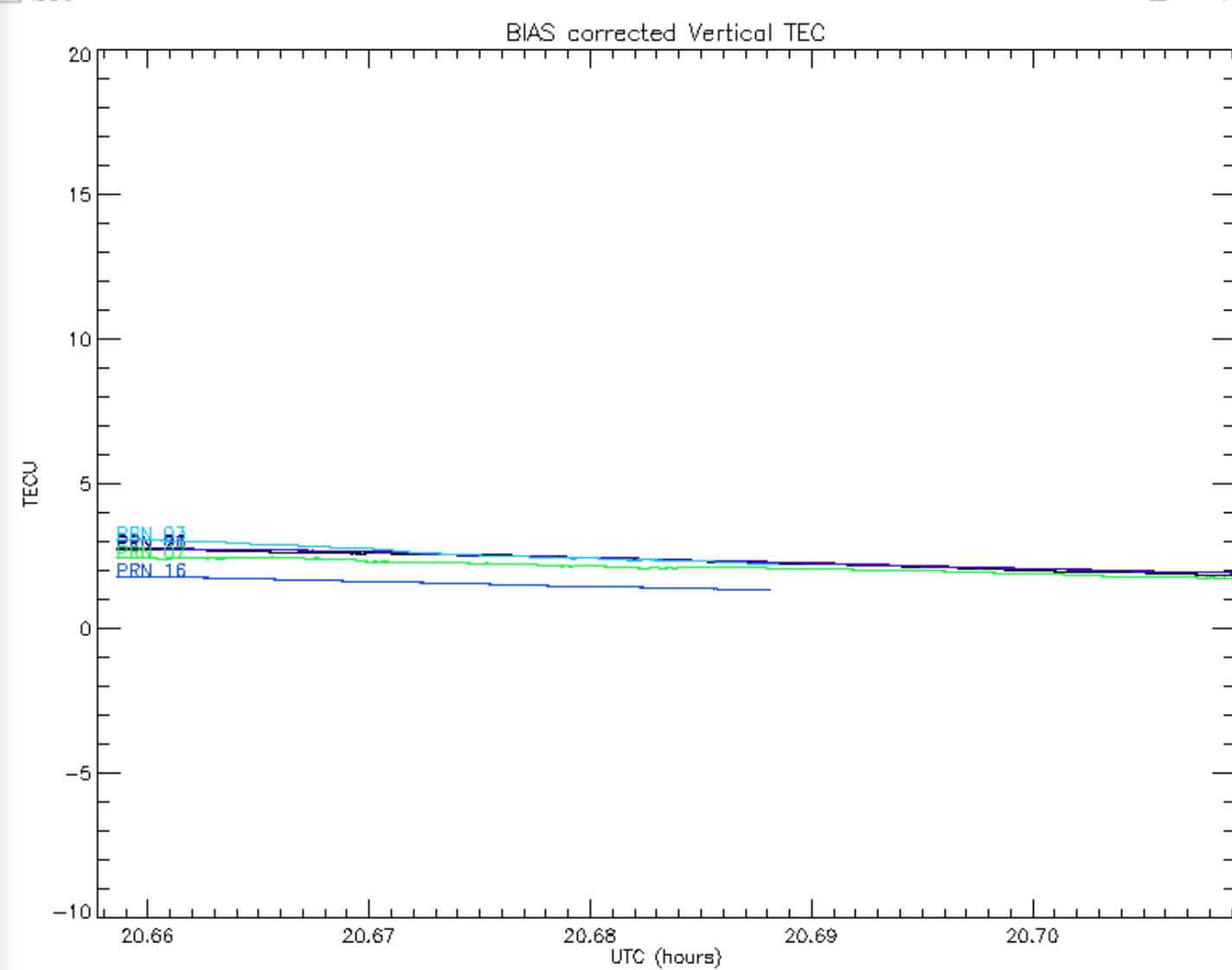
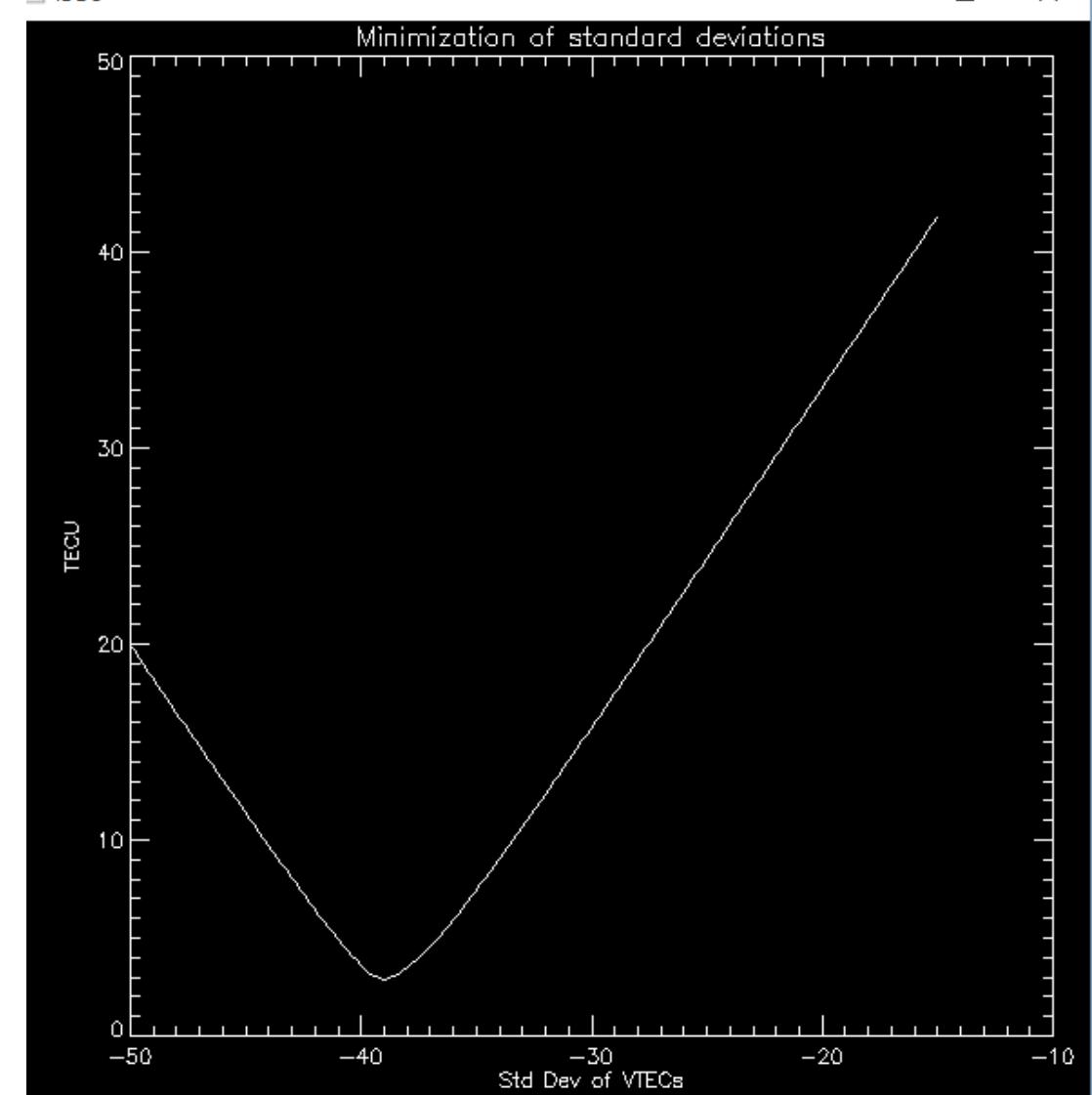
Skyplot (10 degrees elevation cutoff)



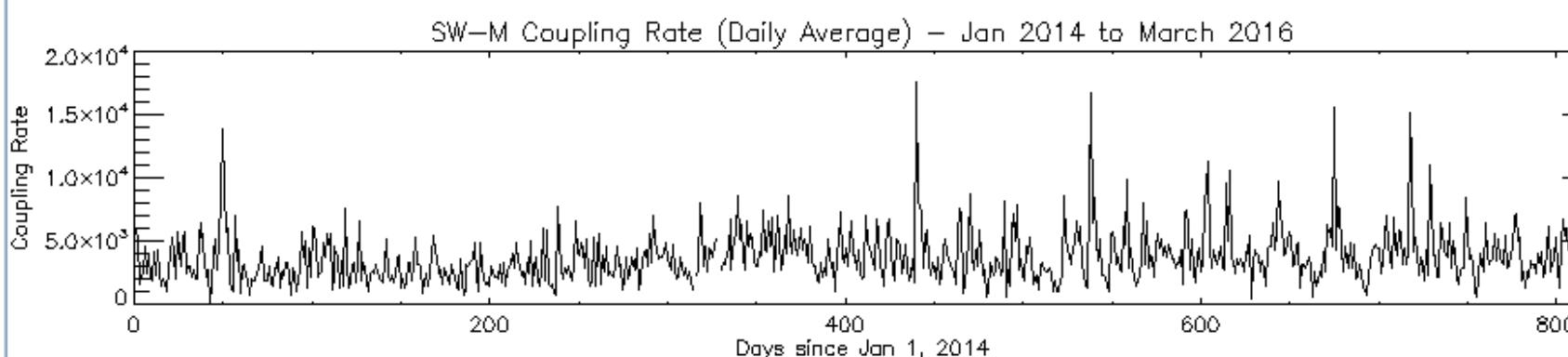
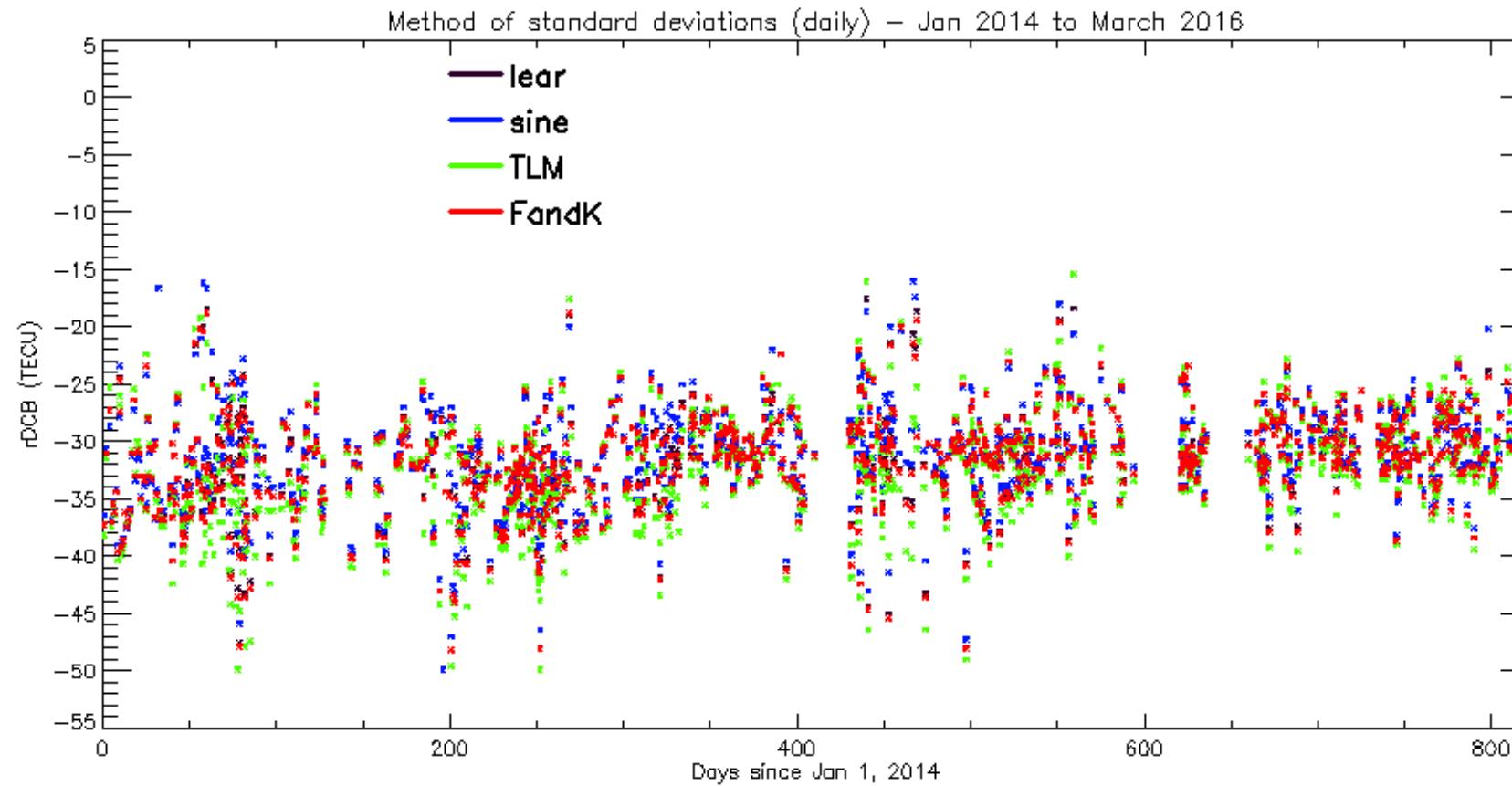
# Minimization of Standard Deviations – Jan 2, 2014

Time averaged  
standard deviation:  $\sqrt{\frac{1}{N} \sum_{i=0}^N (VTEC_i - \bar{VTEC})^2}$

Projected vertical TEC  
(using TLM projection) for  
DCB=-39 TECU



# MSD rDCB estimates (daily) – Jan 1, 2014 to April 1, 2016



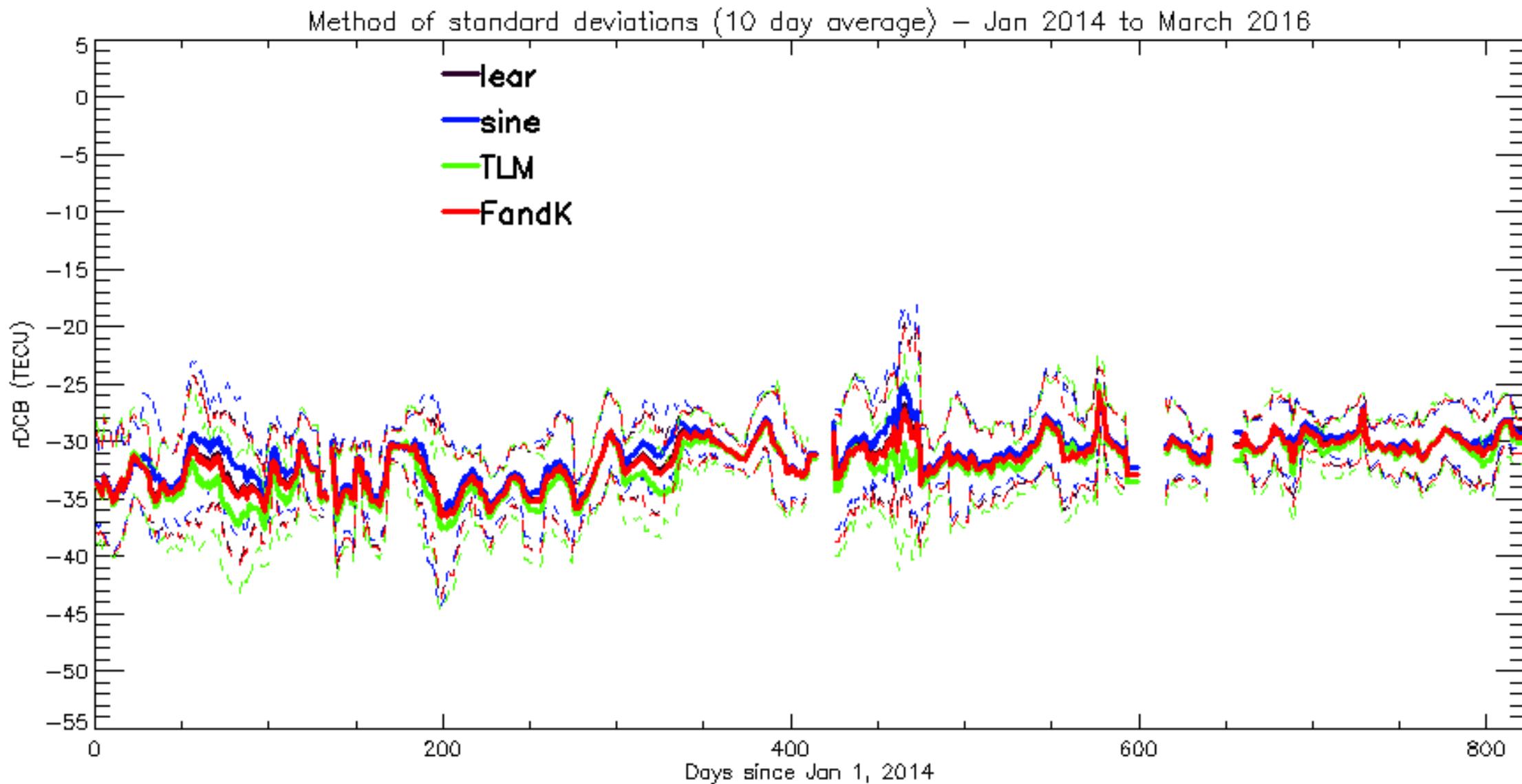
- No clear dependence of day-to-day rDCB variability on solar wind conditions or geomagnetic indices (K<sub>p</sub>, DST, etc.).
- Real temporal variations in rDCB (e.g. associated with temperature changes), or due to errors inherent to estimation method?

**SW-M Coupling Rate**  
**[Newell et al., 2007]:**

$$\frac{d\Phi}{dt} = v_{sw}^{\frac{4}{3}} \left( \sqrt{B_y^2 + B_z^2} \right)^{\frac{2}{3}} \sin^{\frac{8}{3}} \left( \frac{\phi}{2} \right)$$
$$\phi = \tan^{-1} \left( \frac{B_y}{B_z} \right)$$

## MSD rDCB estimates (10 day running average) – Jan 1, 2014 to April 1, 2016

- Apparent long period trend in receiver bias: -35 TECU to -30 TECU over 2 years

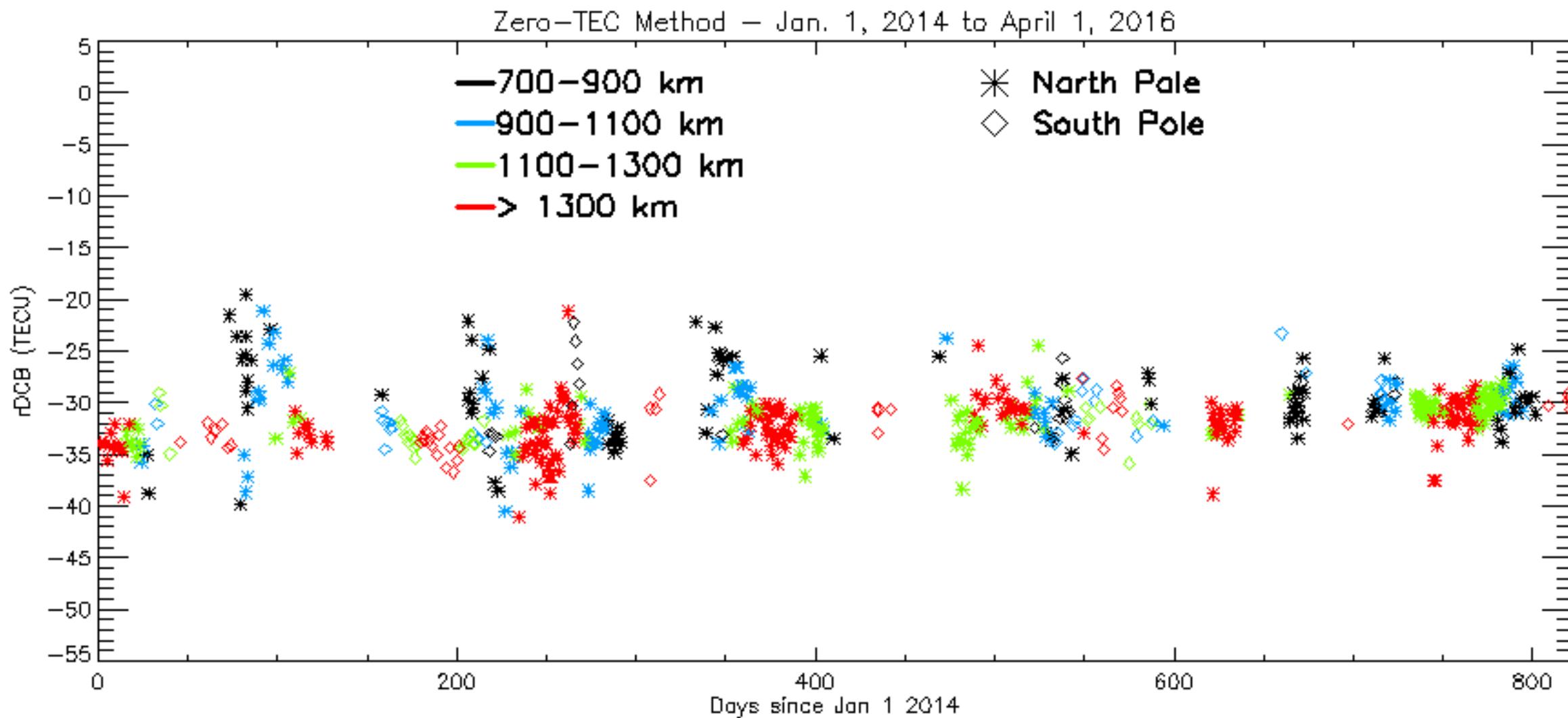


## Zero-TEC Method

- Assume that TEC above e-POP is zero
  - Ideally at apogee, winter polar cap, near solar minimum
- Estimate receiver bias from lowest observed topside TEC
- Advantage: Fast and simple
- Disadvantage:
  - TEC above e-POP is non-zero
  - e-POP not always at apogee in the polar cap
  - Outliers

## Zero-TEC Method

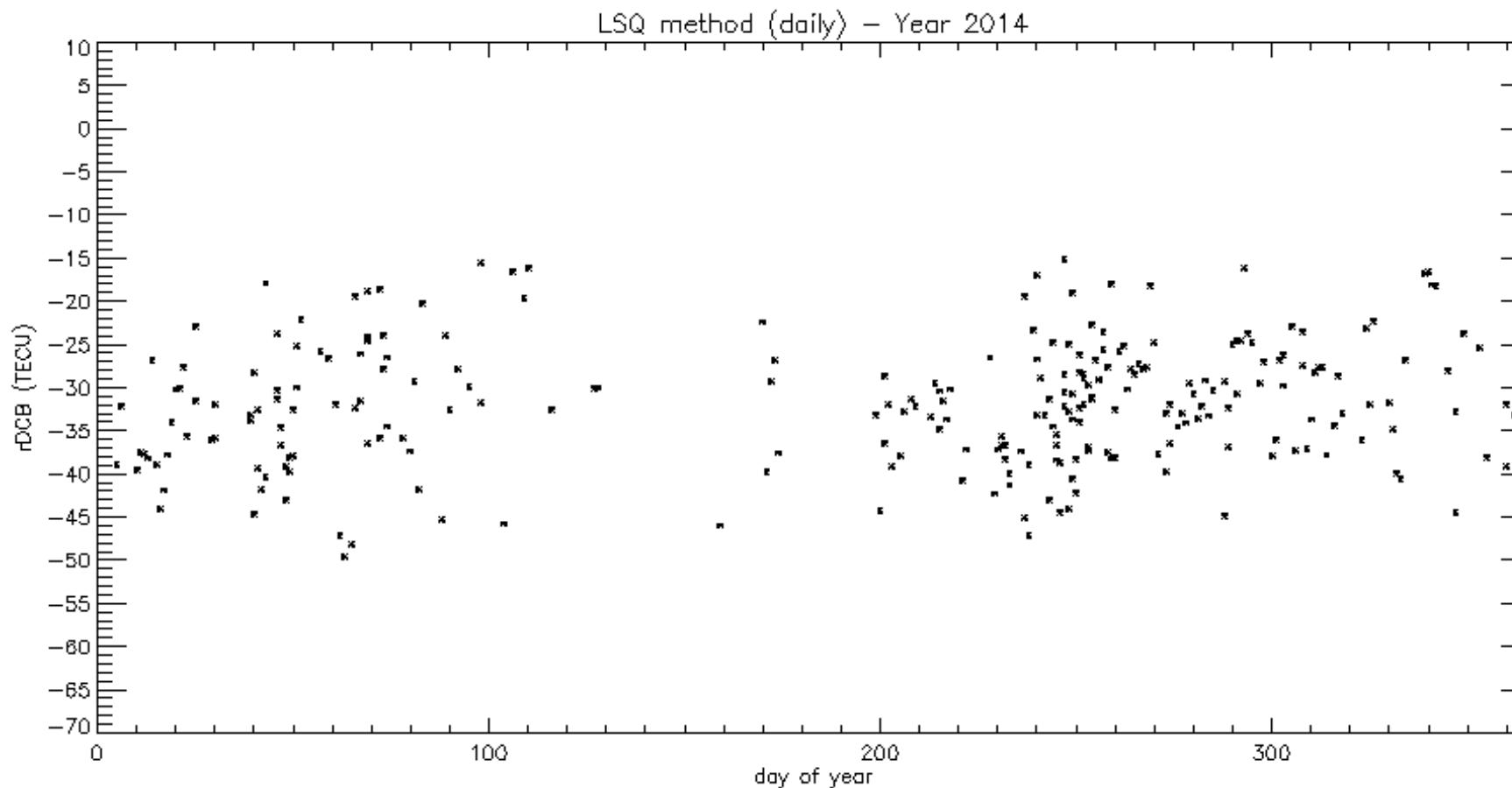
- For polar cap passes at  $> 70^\circ$  and  $< -70^\circ$  degrees magnetic latitude:  
calculate average TEC for GPS satellites at  $> 50^\circ$  degrees elevation



## Least Squares (LSQ) Method

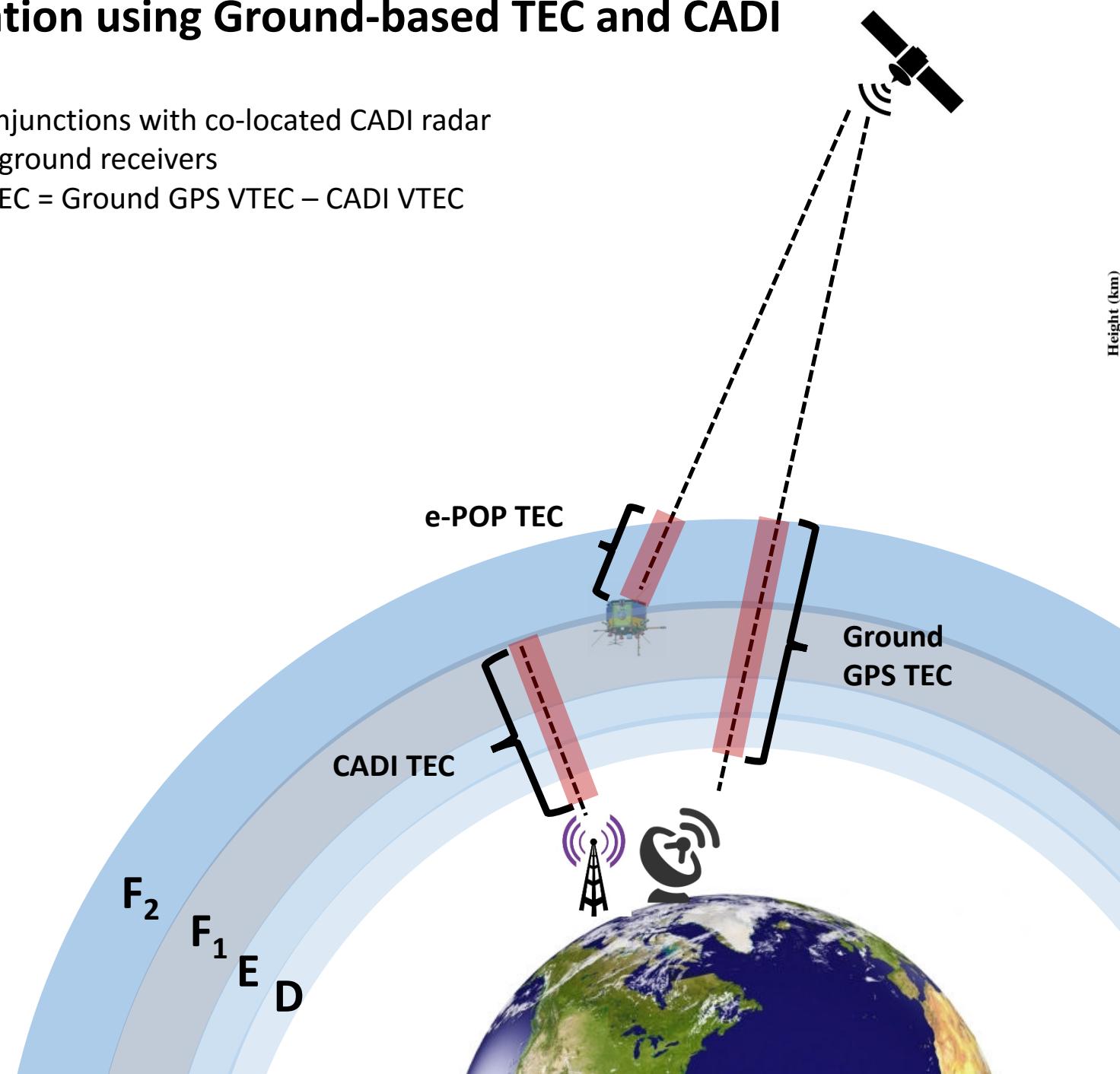
- Assume equivalent VTEC from satellite pairs with near-proximity ray paths
- Combine each satellite pair to obtain a daily set of linear equations
- Solve resulting system of linear equations using LSQ method

$$(TEC_{biased,1} - DCB_r) \cdot M(E)_1 = (TEC_{biased,2} - DCB_r) \cdot M(E)_2$$

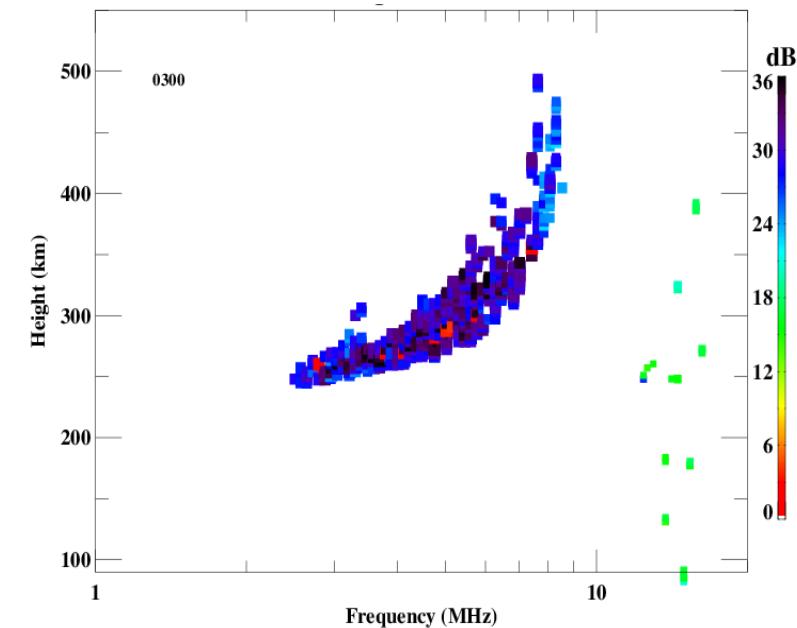


# rDCB validation using Ground-based TEC and CADI

- e-POP conjunctions with co-located CADI radar and GPS ground receivers
- e-POP VTEC = Ground GPS VTEC – CADI VTEC



CADI ionogram (bottom-side electron density)



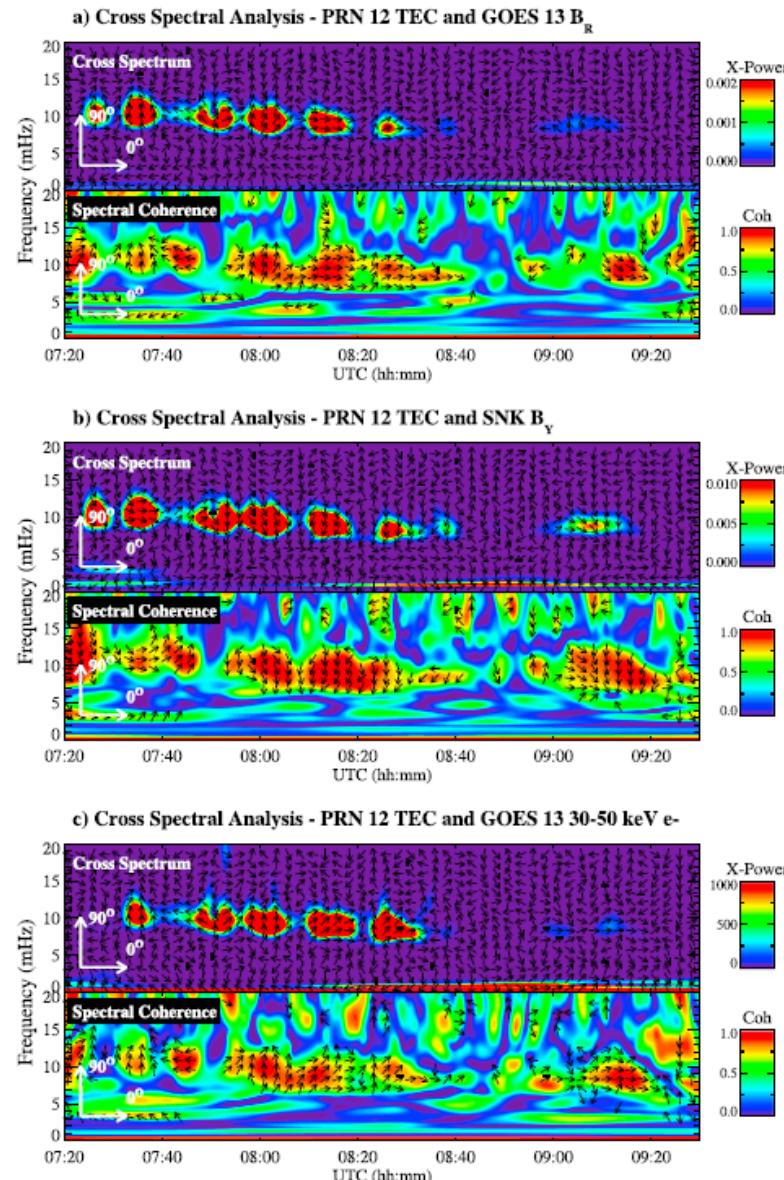
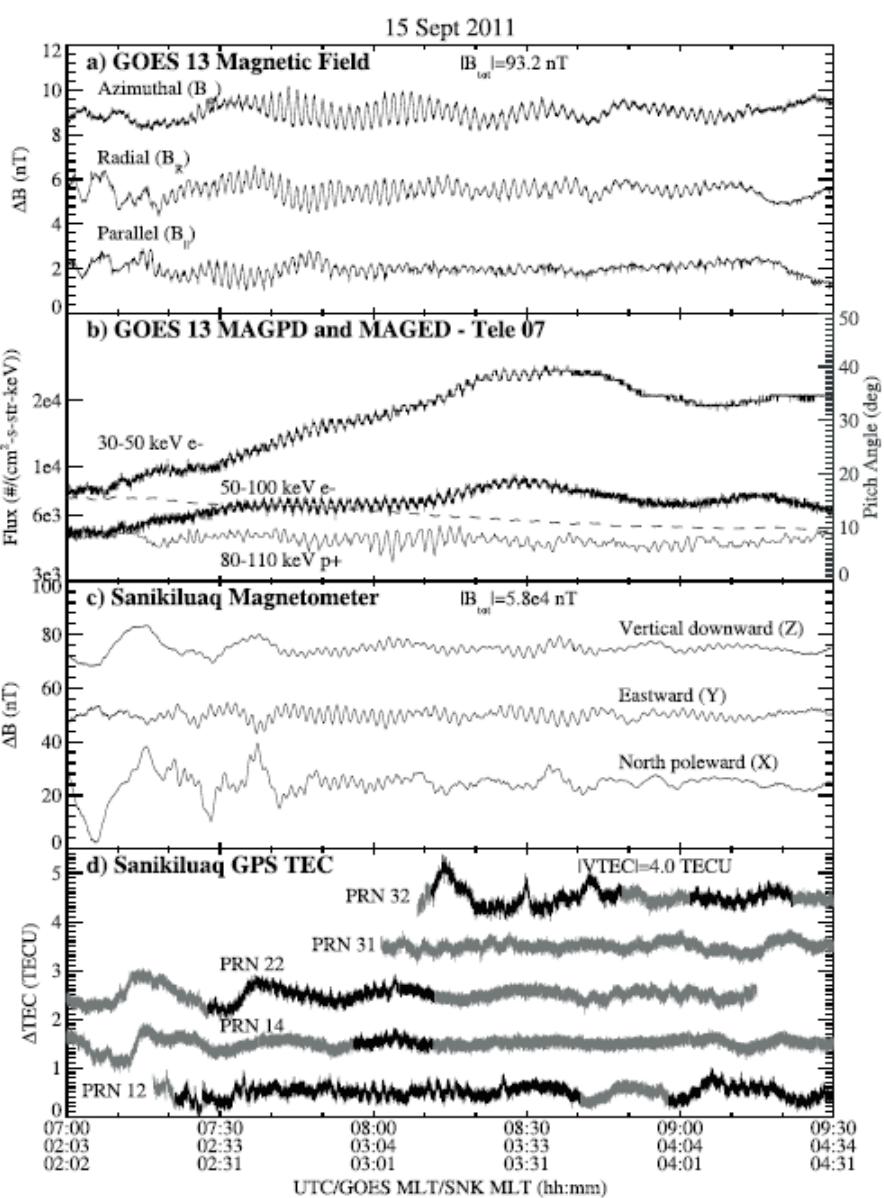
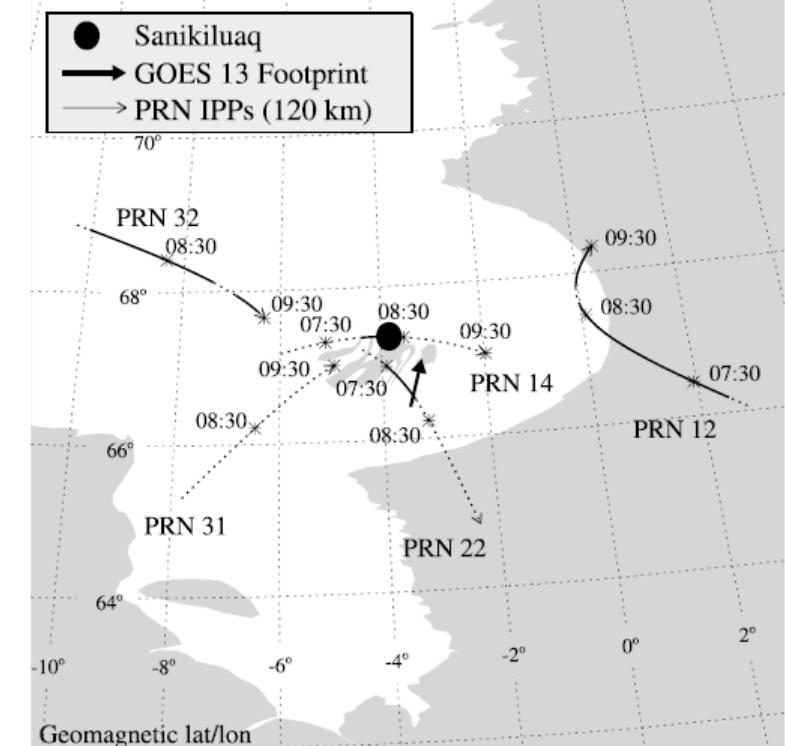
## e-POP – Resolute CADI conjunctions (within 3 degrees latitude/longitude)

DATE	TIME (UTC)	e-POP Altitude (KM)	MSD e-POP VTEC (TECU)	GPS-CADI TEC (TECU)	CADI TEC (TECU)	Kp Index
02/27/2014	07:39:38- 07:39:59	375.3-380.0	5.94±0.88	10.86	4.01 (good)	4-5
03/01/2014	07:38:57- 07:39:13	354.9-357.4	5.11±1.99	12.11	3.38 (very patchy)	1-3
03/03/2014	07:38:03- 07:38:14	342.5-343.5	5.09±2.26	11.31	3.15 (very patchy)	1-2
03/05/2014	07:36:55- 07:36:56	337.4-337.4	9.12±0.56	10.89	3.72 (some patches)	1-2
03/07/2014	14:32:36- 14:32:59	397.1-403.5	10.61±1.08	14.49	8.75 (good)	0-1
03/10/2014	05:48:00- 05:48:03	337.1-337.1	5.85±2.13	8.90	3.23 (good)	1-2

# Combined Ground-Occultation TEC Observations of High Latitude Ionization Structures

Watson et al [2016], GPS  
TEC response to  $Pc4$  Giant  
Pulsations, JGR Space  
Physics

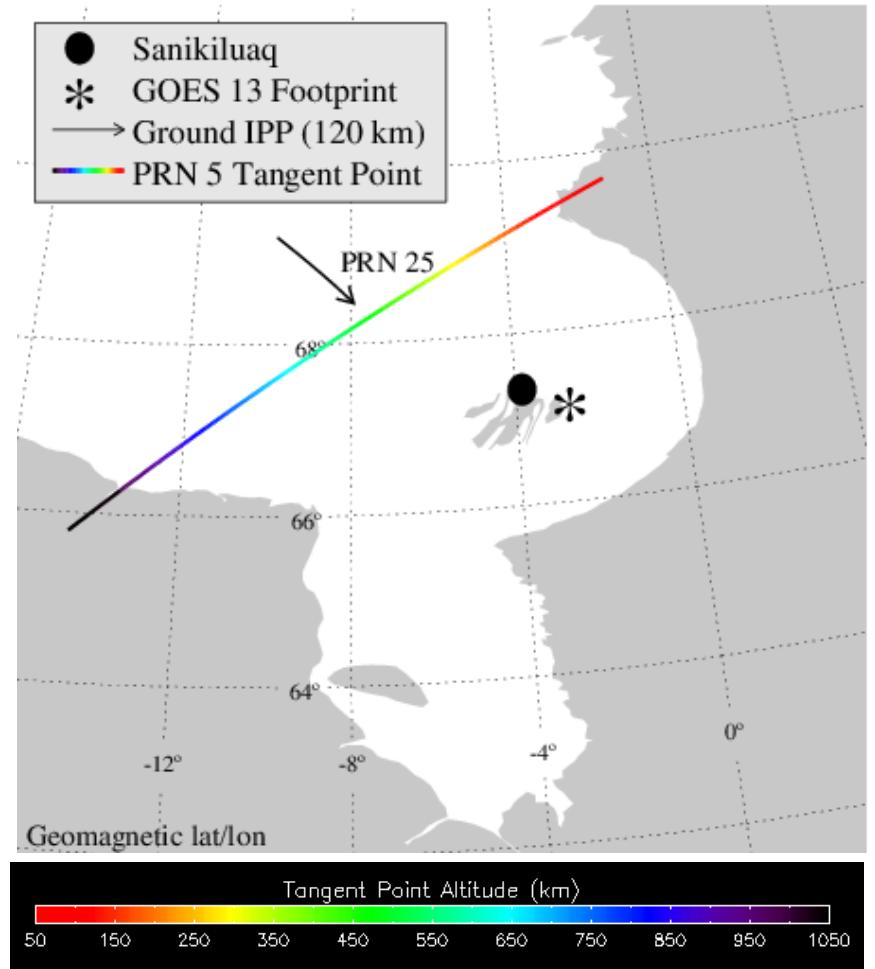
15 Sept 2011 - 07:00 to 09:30 UTC



# Combined Ground-Occultation TEC Observations of High Latitude Ionization Structures

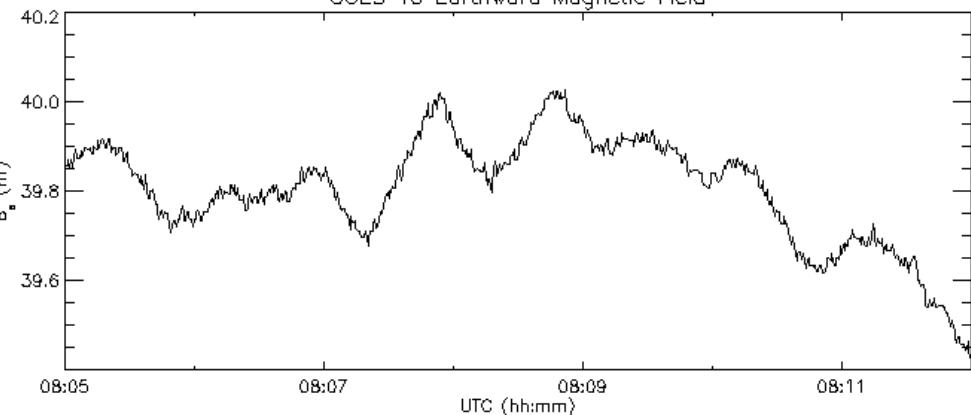
## PRN 5 TEC (Occultation)

12 Feb 2016 - 08:05 to 08:12 UTC



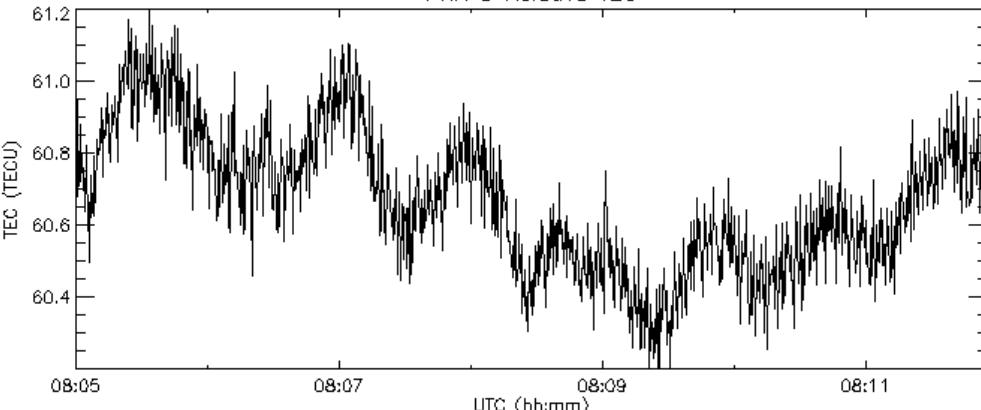
## GOES 13 Magnetic Field

GOES 13 Earthward Magnetic Field

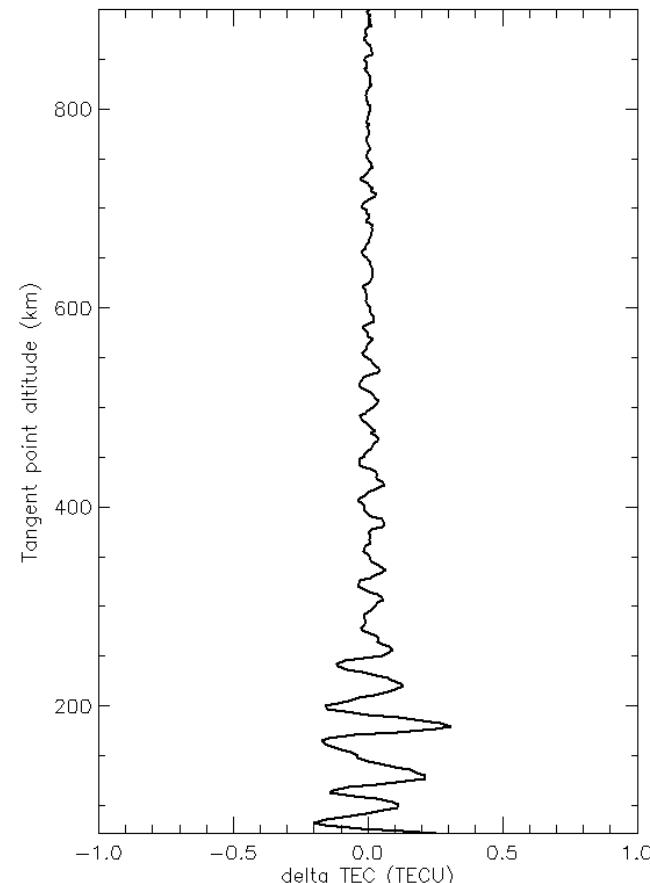


## PRN 25 TEC (Ground Receiver)

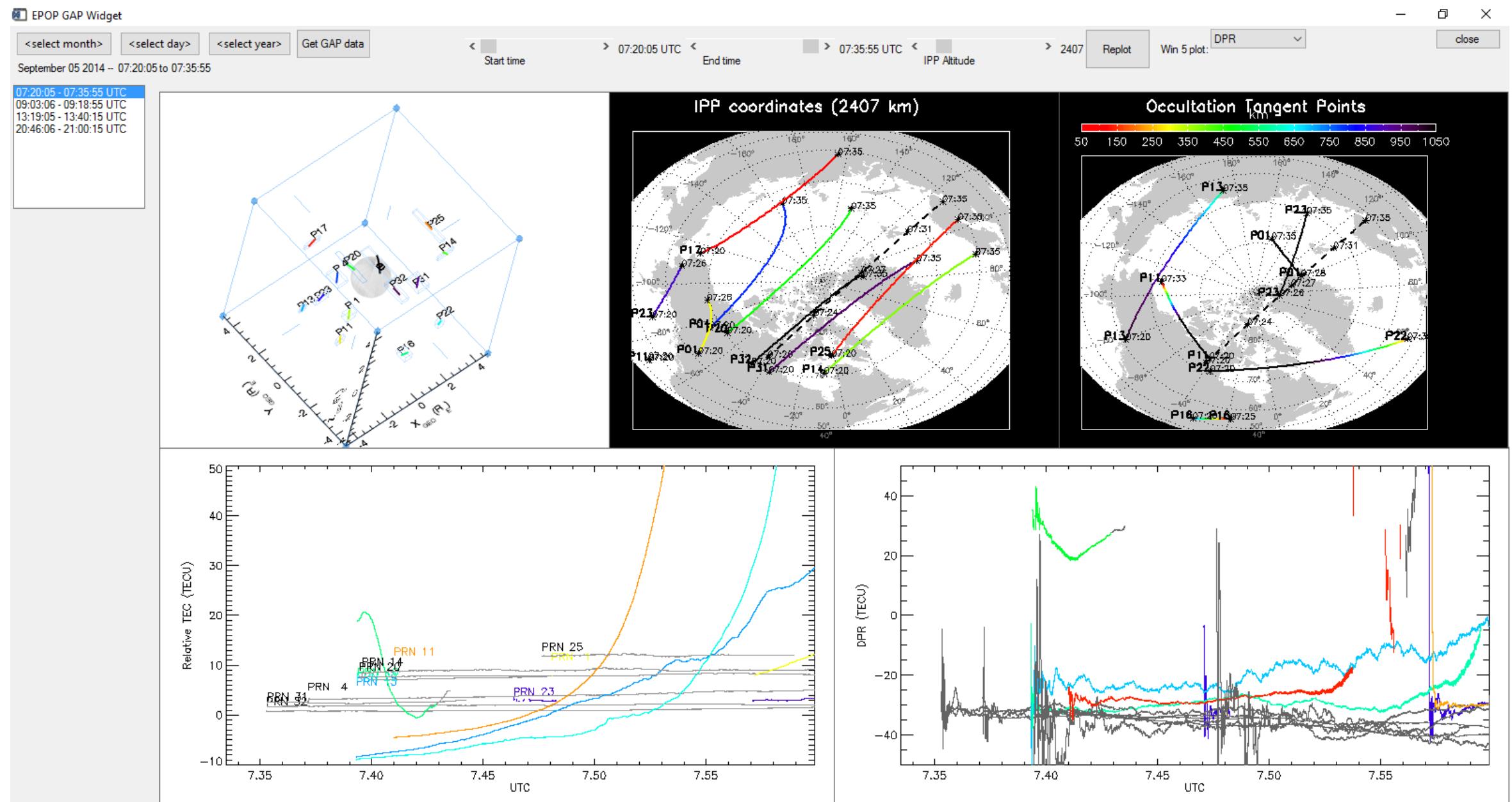
PRN 5 Relative TEC



PRN 5 – 02122016 – 08:05 to 08:12 UTC



# GAP-O Data Availability and Visualization tools



## GAP-O Data Availability and Visualization tools

- sFTP RINEX files via: <http://epop-data.phys.ucalgary.ca/>
- Research teams working on development of e-POP GAP data:
  - Richard Langley (University of New Brunswick)
  - Ionospheric Group at JPL (Esayas Shume, Attila Komjathy)
  - e-POP team at the University of Calgary
- My Email: [Watson@phys.ucalgary.ca](mailto:Watson@phys.ucalgary.ca)

## Conclusions and Future Work

- Ionospheric density profiles inverted from GAP-O occultation TEC currently viable on a case-by-case basis.
- Automated TEC inversion for high latitude regions will require a more advanced method accounting for ionization structure & density gradients.
- Minimization of standard deviations (MSD) is currently the most reliable method for receiver DCB estimation.
- Further investigation required to determine whether temporal variations in rDCB are real or are due to errors inherent to the estimation technique.
- Comparison to CADI & ground based GPS TEC indicates that MSD method tends to underestimate receiver bias.
- Planned studies involving GAP-O data include combined occultation-ground TEC observation of high latitude ionization structures.

## References

- Kim, D. and R. B. Langley (2010), The GPS attitude, positioning, and profiling experiment for the enhanced polar outflow probe platform on the Canadian CASSIOPE satellite, *Geomatica*, 64(2), 233-243.
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- Newell, P. T., T. Sotirelis, K. Liou, C.-I. Meng, and F. J. Rich (2007), A nearly universal solar wind-magnetosphere coupling function inferred from 10 magnetospheric state variables, *J. Geophys. Res.*, 112, A01206, doi:10.1029/2006JA012015.
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- Zhong, J., L. Jiuhou, and Y. Xinan (2016), Determination of Differential Code Bias of GNSS Receiver Onboard Low Earth Orbit Satellite, *IEEE Transactions on Geoscience and Remote Sensing*, doi:10.1109/TGRS.2016.2552542.